

# EIC IR Design Meeting

Draft Minutes for Friday, May 8, 2020

## Agenda

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## 1 Solenoid compensation schemes—V. Morozov

Title: “Solenoid Compensation Schemes”

File: [solenoid\\_compensation\\_08may20.pptx](#)

### 1. Baseline Compensation Scheme (Case 6) [slide 2]

#### (a) Pros:

- i. No impact on the detector acceptance
- ii. No orbit excursion in the rear, modest orbit excursion in the forward
- iii. Low corrector strength, the strengths are practically independent of energy
- iv. Moderate technical complexity with only one non-conventional  $y$  corrector integrated into B0A

#### (b) Cons:

- i. Orbit excursion in the forward part of up to 9 mm at 41 GeV but comparable to  $\sim 20$  mm horizontal shift of the orbit with energy. Should not impact the dynamic aperture significantly.
- ii. Requires a nested  $y$  corrector in B0A but seems technical feasible (see talk by Holger [section 2])

### 2. Backup Compensation Scheme (Case 7) [slide 3]

(a) Both forward  $x/y$  correctors shifted further downstream to locations after all of the forward IR elements

(b) Simplifications compared to the baseline scheme:

- i. Conventional correctors only

(c) Additional complication compared to the baseline scheme:

- i. Closed orbit excursion over a wider range
3. Comments [slide 4]
    - (a) Chose a baseline solution and a backup solution
    - (b) The baseline solution is more technically complicated but feasible
    - (c) The backup solution is technically simpler but may be slightly worse for the dynamic aperture
    - (d) The final choice depends on the dynamic aperture
    - (e) Need to combine solenoid kick and forward quadrupole shifts and rotations
    - (f) Need to check the forward acceptance of the combined system
    - (g) Continue coupling compensation studies
  4. Do not have a final solution for the shifted quadrupoles, but with enough knobs, should be able to correct orbit in the general case.

## 2 B0ApF corrector—H. Witte

Title: “B0ApF Corrector”

File: [2020-05-07\\_b0apf\\_corr.pdf](#)

1. Integrating B0ApF into what V. Morozov calls “case 6” [slide 2].
2. Change inner radius of coil from 50 mm to 60 mm [slide 4].
3. Corrector coil is 1 mm thick [slide 4].
4. Didn’t let the optimization shown on slide 5 finish; so there could be further improvements to harmonics.
5. Integrated corrector field of 67 mT m with a current density of 200 A/mm<sup>2</sup> [slide 6].

## 3 Update on equal divergence studies—V. Ptitsyn

Title: “Luminosity with equal proton divergencies”

File: [lumi\\_with\\_equal\\_divergencies.pdf](#)

1. Main constraints used for luminosity optimization [slide 2]
  - (a)  $I_p < 1$  A,  $I_e < 2.5$  A
  - (b)  $\zeta_p < 0.015$ ,  $\zeta_e < 0.1$
  - (c)  $\sigma'_{x/ye} < 220$   $\mu$ rad
  - (d)  $\beta_{px}^* > 90$  cm,  $\beta_{py}^* > 5$  cm
  - (e)  $\beta_{ex}^* > 35$  cm,  $\beta_{ey}^* > 4$  cm
  - (f) IBS growth time  $> 2$  h

2. Beam flatness is the ratio of the vertical to the horizontal beam size at the interaction point.
3. Scan with equal horizontal and vertical divergencies [slide 3]
  - (a) Blue curve:
    - i. Beam flatness  $\sim 0.04$ – $0.05$
  - (b) Yellow curve:
    - i. Constant beam flatness =  $0.085$
4. More detailed scan at lower divergencies [slide 4]
  - (a) There is a luminosity maximum
  - (b) Left slope: emittance limited (by IBS/beam-beam)
  - (c) Right slope: beta-function limited (beta\*'s at min)
  - (d) Issue: this beam flatness ( $0.085$ ) may be not acceptable because of emittance deterioration in beam-beam interactions.
5. Equal divergence data, for different beam flatness [slide 5]
  - (a) Present recommendation from Beam-beam experts to have the flatness no less than  $0.12$ .
  - (b) Although beam-beam studies continue.
6. Luminosity on beam flatness [slide 6]
  - (a) Difference between highest luminosity and luminosity with  $110\ \mu\text{rad}$   $x/y$  divergence reduces with the flatness increase.
7. Leftmost point on orange (“Highest lumi”) curve is preCDR point [slide 6].
8. E.C. Aschenauer: Losses in luminosity are more than compensated for by increases in acceptance.
9. Further possibilities [slide 7]
  - (a) In further plans: to see if constraints on min proton beta\*'s can be decreased (proposed by R. Palmer), with luminosity improvements. Proton DA studies by Yun Luo will explore these constraints.
  - (b) In pCDR the electron current  $2.5\ \text{A}$  corresponded to  $9\ \text{MW}$  SR power. Do we have some room for current increase? Vacuum, HOMs.
10. E.C. Aschenauer: If we can do this, it will make a big difference for the forward physics.
11. Parameter files:
  - (a) [eq\\_div\\_110\\_flatn\\_120.txt](#)
  - (b) [eq\\_div\\_110\\_ibs\\_flatn.txt](#)

**4 All other business**

None

**5 Draft agenda for Friday, May 15, 2020 from 2:30 to 3:30 p.m.**

1. Solution to parameter set with equal  $\beta^*$  for protons and electrons—B. Palmer
2. Physics simulations with collider magnets inside detector region—friends from Physics
3. Status and to-do list for CDR IR section—E.C. Aschenauer, C. Montag, and H. Witte
4. All other business