# Calorimeter calibration and related data production in the 2023 run

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### overview

### Calibrations

- Cosmic muon calibration of HCals
- Pi0 and "tower slope" of the EMCal
- Software implementation

#### Hardware relevant studies

- Understanding pathological calorimeter data
- Software solutions

### State of calorimeter simulations

- Noise simulation
- Photon energy scale and resolution

### Calorimeter data preparation

# Signal processing

# Same raw data processing in the Hcals and EMCal

#### ≻Raw data for 2023

 31 14-bit samples per a channel, digitized 6x beam clock

### Signals are fit with template waveform

- The saved output of the fit is
  - Amplitude (floating point ADC max of the waveform)
  - peak time (saved in integer milli-sample precision)
  - pedestal float point precision
  - Least-squares difference between the fit and waveform data divided by number of degrees of freedom
- We currently fit 16 samples (not well studied but no large degradation in performance for 12 samples)



### **Template fit performance**

#### >Chi2 vs. ADC with beam template.

- Noise levels in the HCals are at the level of 2-5 ADC, which indicates low energy signals are consistent with a good template description.
- Towers can have distorted waveforms for an entire run (example at chi2=1e6)
- Some towers have distorted ~1% of the time

#### >Both cases of problematic towers are caught by a Chi2 cut

#### >The other calorimeters look very similar

#### Computing speed

chi2

- To fit all channels is several seconds per an event
- This can be speed up by an order of magnitude by software zero suppression.



### **EMCal calibration**

# $\succ$ The EMCal calibration is factorized into two dimensions, $\eta$ and $\phi$

#### $ightarrow \Phi$ dimension: The tower slope method

- Azimuthal symmetry is assumed in the tower energy spectra, up to a gain factor.
- The slope (dN<sub>twr</sub>/dE) is a function of E and thus a spectra can be fit with another to find the relative gain factor.
- All towers with the same  $\eta$  are fit with the  $\varphi$ -integrated spectra yielding a relative calibration for all towers with the same  $\eta$  acceptance

#### If a dimension: pi0 mass spectra, binned in the leading tower of the leading cluster

- The pi0 mass peak is extracted via a fit and the ratio of this with a target mass is the calibration factor.
- Because the mass is a product of many calibration factors an iterative approach is to used.
- The result is an absolute data-driven calibration in the  $\eta$  dimension.
- The mass target is arbitrary at this point and is commented later.

### **Tower slope method**

Fit a particular  $\eta$ - $\varphi$  with a template(E\*gain) where the template is the sum of towers spectra for the particular  $\eta$ 

• Fit range E=[0.12,0.7] GeV

### >Need to improve QA on fits and extracted coefficients

#### Justin Ohio U



# **Pi0 calibration**

#### >Example: 2.1M events

Runs: 21598, 21609, 21608, 21599, 21616, 21615

### >Speed: non-parallelized

- Tower slope 5 hours
- Pi0 10 iterations: 24 hours (parallelized 2h)

### ≻Cuts

- $p_{T1} > 1.3 + 1.4 * (N_{clus}-30)/200$
- $p_{T2} > 0.4 + 1.4 * (N_{clus}-30)/200$
- Only add  $p_T$  for  $N_{clus} > 30$
- ΔR < 1.1
- Asym < 0.6

>3% peak error! This means there is an additional 3% relative uncertainty/smearing due to calibrations which is not expected to improve



### **η asymmetry in current calibration**

- It has been presented in PC meetings of an obvious asymmetry in EMCal tower energies.
- This mis-calibration arose from the unavailability of a proper vertex at the time of the QM calibration (the current used in DSTs)
- A default value of 0 was used which biased the extracted pi0 mass
- Right: he pi0 mass extracted by the same clusters (detector eta/phi and E) but a different vtx used to assign the physics eta



### HCal cosmic muon calibration overview <sup>9</sup>

- > The calibration for both the inner and outer HCals is derived by comparing energy deposits from cosmic muons between data and MC
- Data is collected with single trigger tower threshold trigger with calorimeters in high gain mode
  - Unprescaled rates of 5-20 kHz
- MC: We use a dedicated MC generator called <u>EcoMug</u>
- > Offline signal criteria is applied to select muons with a long path length, passing through all tiles in a tower, seen on the right.

#### Hanpu, Shuhang, Columbia





#### **HCal cosmic muon calibration results** 10

- The distribution is characterized with a fit and the peak position is used for calibration.
- The ADC is compared to the simulated light yield from the scintillator.
- The calibration factor is then scaled by the high-low gain ratio and by the MC-derived pion sampling fraction
  - Said another way, the calibration matches data-MC MIPs scintillator energy of a particular kinematics/species and uses the MC to translate that into energy deposited in scintillator +absorber by beam-like particles.
- Bellow are the comparison of η-φ cosmic MPV in data (left) and MC (right)
  - · Pre-installation calibrations already applied to data
  - Top sPHENIX iphi=15-16





8.5

6.5

60



### **Current status of calibration**

### >Hcal cosmics

- Analysis is on git <u>sPHENIX-Collaboration/analysis/HCalCosmicCalib</u>
- Working on updating the calibration with cosmics taken during the months of beam data-taking which will account for changes in pulse fitting – expected in the coming weeks
- Now comparing to official sphenix cosmics simulation
  - CreateFileList.pl -type 22 DST\_CALO\_CLUSTER -run 11
- Exploring dependence on muon trajectory

#### **EMCal calibration**

- Analysis is on git <u>sPHENIX-Collaboration/macros/calibrations/calo/calib\_nov23</u>
- Original calibration for QM was done without vertex info which caused issues
- Automated fully PR'ed process has been implemented and new results are expected to be available in the next weeks

### What is in the DSTs

#### Raw and calibrated <u>TowerInfov2</u> objects

- Data fields
  - Amplitude (floating point ADC max of the waveform)
  - peak time (saved in integer milli-sample precision)
  - pedestal float point precision
  - Least-squares difference between the fit and waveform data divided by number of degrees of freedom
- Status fields
  - get\_isBadChi2: if the chi2 is above 1e4 for that event
  - get\_isHot: all towers with isBadChi2 > 5% of hits during the run
  - get\_isBadTime: set to true is the event/tower has a peak time greater than 2 (1) from that runs hit time mean for the HCal (EMCal).
  - get\_isGood: if all the above are false, will return true
- No modification of data fields (all masking is done but users using status bits)

#### >EMCal clusters created with a photon template

- Clustered with towers which have no quality cuts
- No hot cluster removal right now

# **Timing cuts**

### >Left: mean EMCal time for hit (> 0.5 GeV)

- There is effect of a rounding error when detector is timed in due to disagreement of FPGA time and GTM (experts can confirm)
- Middle: single EMCal channel (ieta=20 iphi=62) timing distribution for hists (top) all events (bottom)
- Right: inner and outer mean channel time



### **Bad chi2 values**

- Right: fraction of hits that have a bad chi2 value
- >> 5% is considered a hot tower
- Bad towers are easily determined by this metric
- Emma has found towers with noisy bits that lead to bad chi2s at the 1-5% of events (hard to see in this plot)



## Identifying pathological tower behavior 15

### >Hot towers were initially identified statistically

Inordinately large average energy or number of hits

#### The underlying pathological behaviors were identified in both EMCal and HCal

- Towers with jumbled bits for an entire run
- Towers with sometimes rare single bit flips
- Appears read-out related

#### Developed event-by-event identification strategy

- Pulses are fit with template waveforms, which is used to generate a very poor goodness-of-fit metric for jumbled/flipped bits
- coresoftware was updated to calculate this and save it to tower objects

# Passed information to hardware experts for possibly identifying hardware or operational solutions

### **Examples of pathological towers**

 $\langle RMS \rangle \pm \sigma_{RMS}$ 

 $11.9635 \pm 3.75301$ 

 $12.1037 \pm 4.15951$ 

 $12.0316 \pm 3.70411$ 

 $12.0461 \pm 3.70546$ 

 $12.056 \pm 3.69148$ 

 $12.2218 \pm 3.72039$ 

 $12.0275 \pm 3.77564$ 

 $12.053 \pm 3.71444$ 

 $12.0946 \pm 3.71661$ 

 $12.6984 \pm 3.72108$ 

 $12.6266 \pm 3.70581$ 

 $12.6904 \pm 3.69699$ 

 $12.7552 \pm 3.77643$ 

 $12.657 \pm 3.6937$ 

#### Pedestal data: Pedro Nieto-Marin Iowa State



### State of MC

> The philosophy leading up to QM and, depending on person power, the strategy in the future is to calibrate to the existing MC, this way previous truth-reco physics definitions stay the same, such as the JES.

#### There are some additions that need to be made, given what was learned from run 23

- The noise in the simulation is 1 ADC, this is a large underestimate and is channel dependent
- Run 23 conditions such as live towers

#### >Additionally, there are current features of the MC that make it complex to calibrate to.

- Pi0 mass has a nontrivial eta-phi dependence
- Inaccurate intrinsic or noise related resolution leads systematic effects

# EMCal

#### $\triangleright$ As previously said the current calibration calibrates to an arbitrary pi0 mass

- > The correct pi0 mass depends on
  - What the mass is in MC
  - Account for any difference in data and MC

#### $\succ$ We need to measure the resolution in data and compare it to MC.

- Any discrepancy need to be propagated to the MC or account for it in the calibration
  - For example, a lower resolution produces a large observed pi0 mass given a falling pT spectrum
- > Need to finish up single photon energy closure in MC i.e. finalize Position dependent correction
- Left: pi0 mass in HIJING without position dependent correction
- > Right: example of truth studies of resolution effects on pi0 mass

#### <u>Sijan + Justin</u>



#### <u>Nikhil Kumar</u>



# Thanks!

### **EMCal calibrations 24**

#### ≻1B a day

- >1.4 s budget Calorimeter processing
- >This takes 10s of seconds in the current setup
- Plan is to go from prdf -> raw towers only once
  - This is a necessity given computing resources
- It is the case that we did not collect enough statistics in 23 to rehearse the planned 24 calibration strategy although the 23 approach is of similar complexity.

#### Will we take 1B events in commissioning to "rehearse" EMCal Calibration?

### Useful

#### Group disk space /sphenix/tg/tg01/commissioning/CaloCalibWG/

#### Software twiki

### HCal 2024

>When watch shifts start, have shifters take cosmics once a day (for an hour or so) to establish and test a reliable routine.

>During beam time, take cosmics during down time.