

# The RHIC Cold QCD Program

## 1 Contents

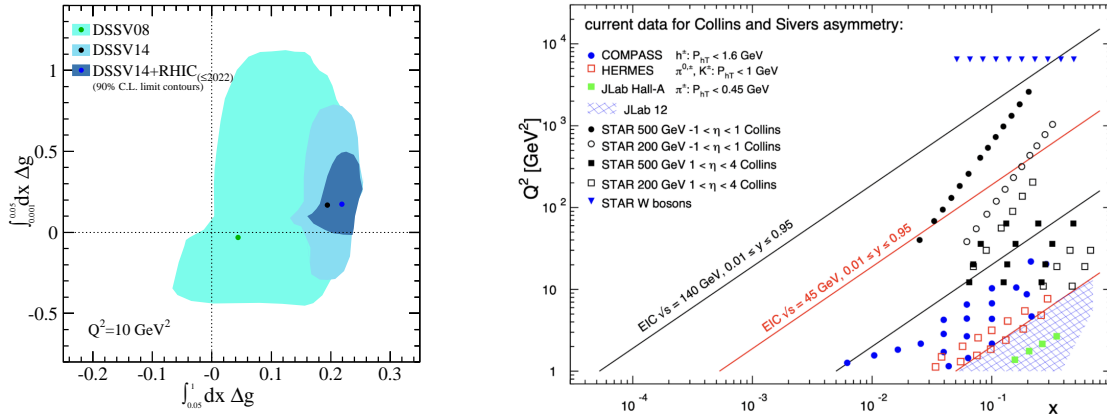
2	<b>1 Executive Summary</b>	<b>2</b>
3	1.1 Recommendations and Initiatives . . . . .	5
4	<b>2 Polarized High-Energy Proton Beams</b>	<b>6</b>
5	2.1 Preparation and Preservation of Hadron Beam Polarization . . . . .	6
6	2.2 Polarimetry of High-Energy Hadron Beams . . . . .	8
7	<b>3 Collinear Proton Structure</b>	<b>11</b>
8	3.1 $W$ $A_L$ and sea quark polarization . . . . .	12
9	3.2 Double helicity asymmetries $A_{LL}$ and gluon polarization . . . . .	13
10	3.3 Nonlinear QCD effects . . . . .	15
11	<b>4 Three-dimensional Structure</b>	<b>19</b>
12	4.1 Studies of initial and final state TMD effects with jets . . . . .	21
13	4.2 Transversity from di-hadron interference fragmentation functions . . . . .	24
14	4.3 Transverse single-spin asymmetry in the forward region . . . . .	25
15	4.4 Transverse single-spin asymmetry of weak bosons . . . . .	27
16	4.5 Transverse single-spin asymmetries of direct photons and heavy flavor decay	
17	leptons . . . . .	28
18	4.6 Nuclear dependence of single spin asymmetries . . . . .	29
19	4.7 Ultra-peripheral collisions . . . . .	30
20	<b>5 Appendix</b>	<b>33</b>
21	5.1 STAR Forward Upgrade . . . . .	33
22	5.2 sPHENIX Detector . . . . .	35

# 1 Executive Summary

In 2025 RHIC data taking will come to an end, concluding 25 years of innovation in accelerator science and advances in the experimental techniques necessary to collide highly polarized, high-energy proton beams. These achievements, discussed in detail in Section 2, include the design and construction of the the world’s highest luminosity polarized proton source, the use of Siberian snakes to reduce the depolarizing effects of the resonance field harmonics, the ability to set bunch-by-bunch polarization directions in order to minimize the systematic effects due to correlations between the spin direction and bunch intensity, the implementation of spin rotators that allow proton beams to be polarized in the longitudinal, transverse or radial direction and the development of techniques to maintain orbit and emittance stability from injection to full energy to allow maximal polarization lifetimes. In parallel, new techniques and tools were developed to monitor and evaluate the quality of the beams. As a result it is now possible to make precision measurements of the beam spin tune, to extract the transverse/radial polarization component in a longitudinal polarized beam and precisely measure the spin dependent relative luminosities. These advances, along with the design, construction, and operation of absolute and relative high precision hadron polarimetry have played an essential role in the success of the RHIC experimental cold QCD program and have laid the foundation for the design of the future Electron-Ion Collider’s (EIC) highly polarized high energy hadron beams.

RHIC has driven the exploration of the fundamental structure of strongly interacting matter into new territory and will continue to enable advances in the field for years to come. These explorations have always thrived on the complementary nature of lepton scattering and purely hadronic probes. This is demonstrated clearly in the flagship measurements of the gluon and sea-quark helicity distributions that are discussed in detail in Section 3. The sea-quark program exploited the advantages afforded by high energy hadron beams, using  $W^{+/-}$  production to reveal the flavor asymmetry of the  $\Delta\bar{u}$  and  $\Delta\bar{d}$  distributions without the complications of fragmentation effects. Similarly, reconstructed jet and pion asymmetries were used for the first time to directly probe gluon interactions in proton-proton collisions, discovering a sizable gluon helicity distribution in the region  $x > 0.05$ , as shown in the left panel of Fig. 1.

The RHIC cold QCD program leveraged the techniques and tools developed in the high profile helicity program to open new frontiers in the rapidly evolving field of transverse spin physics. For example, the reconstruction of  $W$  bosons in transversely polarized proton collisions is used to test the predicted sign change of the Sivers’ function and to provide the first constraints on the sea-quark Sivers functions. Hadron-in-jet asymmetries, measured for the first time at RHIC, and di-hadron asymmetries provide access to the collinear quark transversity distributions, as well as the transverse momentum dependent (TMD) Collins Fragmentation Function (hadron-in-jet) and collinear Interference Fragmentation Functions (di-hadron) in the final state. These new channels, and many more, are discussed in detail in Section 4. Again, the transverse spin program exploited the complementarity of the high energy hadron collider configuration by accessing distributions originally measured in lepton scattering experiments, but in a different kinematic regime, allowing for new insights into



**Figure 1:** Left: The impact of RHIC data to constrain gluon helicity [1, 2]. Right: The  $x$ - $Q^2$  probed with data from the future EIC and Jlab-12 GeV as well as the current SIDIS data and the jet and  $W$ -boson data from RHIC. All data are sensitive to the Sivers function and transversity times the Collins FF in the TMD formalism.

66 As the realization of a future EIC draws closer, there is a growing scientific imperative  
 67 to complete a set of “must-do” measurements in  $pp$  and  $pA$  collisions in the remaining RHIC  
 68 runs. The ongoing RHIC cold QCD program will build on the accelerator’s unique ability  
 69 to collide a variety of ion beams in addition to polarized protons, and a detector with wide  
 70 kinematic coverage that has been further enhanced through an upgrade at forward rapidities  
 71 consisting of electromagnetic and hadronic calorimetry as well as tracking. The forward  
 72 upgrade, including its forward tracking capabilities, will make possible charged hadron identifi-  
 73 cation and full jet reconstruction in the forward direction for the first time, allowing RHIC  
 74 to extend the full complement of the existing transverse spin program into new kinematic  
 75 regimes! This will expand the existing transverse spin program into both lower and higher  
 76  $x$  domains, as illustrated in the right panel of Fig. 1. In addition to the expanded trans-  
 77 verse spin program, RHIC will be able to further explore exciting new signatures of gluon  
 78 saturation and non-linear gluon dynamics (see Section 3). The ratio of forward Drell-Yan  
 79 and photon-jet yields in  $pp$  and  $pA/AA$  collisions are clean probes of nuclear modifications  
 80 to initial state parton distributions as well as gluon saturation effects. All of these measure-  
 81 ments rely critically on the successful completion of runs currently scheduled to be completed  
 82 before the 2025 RHIC shutdown.

83 While the remaining RHIC cold QCD program is unique and offers discovery potential  
 84 on its own, it is also essential to fully realize the scientific promise of the EIC. These data  
 85 will provide a comprehensive set of measurements in hadronic collisions that, when combined  
 86 with EIC data, will establish the validity and limits of factorization and universality. The  
 87 separation between the intrinsic properties of hadrons and interaction dependent dynamics,  
 88 formalized by the concept of factorization, is a cornerstone of QCD and largely responsible for

89 the predictive power of the theory in many contexts. While this concept and the associated  
90 notion of universality of the quantities that describe hadron structure has been successfully  
91 tested for unpolarized and - to a lesser extent - longitudinally polarized parton densities, its  
92 experimental validation remains an unfinished task for much of what the EIC is designed  
93 to study, namely the three-dimensional structure of the proton and the physics of dense  
94 partonic systems in heavy nuclei. To establish the validity and the limits of factorization  
95 and universality, it is essential to have data from both lepton-ion and proton-ion collisions,  
96 with an experimental accuracy that makes quantitative comparisons meaningful. The final  
97 experimental accuracy achieved with the data collected during this final RHIC campaign will  
98 enable quantitative tests of process dependence, factorization and universality by comparing  
99 lepton-proton with proton-proton collisions. When combined with data from the EIC, it will  
100 provide a broad foundation to a deeper understanding of Quantum Chromodynamics.

## 1.1 Recommendations and Initiatives

The RHIC cold QCD community proposes the following recommendations. These proposals were presented, discussed and received strong support at the QCD Town Hall Meeting in September of 2022.

1. Continued funding of RHIC operations to enable collection of the last  $pp$ ,  $pA$  and  $AA$  datasets that are required for completion of the RHIC Hot and Cold QCD missions.
2. Continued strong funding of the RHIC Hot and Cold QCD experimental analysis groups well beyond the final operation of RHIC. This will enable effective and timely analysis and publication of the wealth of data collected, and to be collected. The unprecedented originality of the RHIC data sets have discovery potential on their own and are critical to fully accomplish the scientific mission of the EIC. Continued support of the experimental groups is also critical to ensure the continued effective training of the next generation of experimentalists in preparation for EIC operations.
3. Continued strong funding of the RHIC Hot and Cold QCD theoretical groups and collaborations well beyond the final RHIC operations to ensure that the knowledge generated by analyses of the RHIC data are fully incorporated into the next-generation of theoretical interpretations.

## 2 Polarized High-Energy Proton Beams

The Relativistic Heavy Ion Collider, RHIC, has the unique capability to collide polarized protons at center-of-mass energies up to 510 GeV. High beam polarizations are an important prerequisite for the efficient and timely execution of the physics program at RHIC, since the figure of merit for any spin-dependent observable is directly proportional to the square of the polarization of the beam and the number of events. For double-spin observables, it is the product of the two beam polarizations which enters into the figure of merit, so losses that affect both beams similarly are potentially even more severe.

The BNL collider complex is using a set of accelerators between the proton (ion) source and the main RHIC collider. The proton beams are injected in up to 120 bunches which can collide in several different interaction points. Proton-proton collisions can reach center-of-mass energies up to 510 GeV. Other collision energies in the past have included 64 and 200 GeV with longitudinal or transverse beam polarizations. Proton bunch intensities reach  $2.7 \cdot 10^{11}$ , resulting in peak luminosities of about  $5 \cdot 10^{32} \text{ cm}^{-2}\text{s}^{-1}$  with average beam polarizations of  $\langle P \rangle \approx 55\%$ . The recent RHIC Run-22 delivered about  $800 \text{ pb}^{-1}$  with an average  $0.68 \text{ pb}^{-1}$  per week (the luminosity was limited to  $1.27 \cdot 10^{32} \text{ cm}^{-2}\text{s}^{-1}$  by request from the experiment).

### 2.1 Preparation and Preservation of Hadron Beam Polarization

Unlike electron beams, hadron beams at the RHIC energy scale lack any significant synchrotron radiation mechanism for self-polarizing or natural emittance damping. This means that the production of high-energy, high-brightness, highly polarized beams consists largely of creating intense, low emittance polarized beams at the source and carefully preserving the beam quality during every stage of acceleration. Developments in both source technology and accelerator physics are therefore both of key importance in supporting the RHIC spin physics program.

The RHIC polarized proton beams are created in an optically pumped polarized ion source (OPPIS). The source has undergone continuous development during the RHIC era, including a major upgrade in 2013. Source technology upgrades include changing from an electron cyclotron resonance source to a fast atomic beam source (ABS), addition of a superconducting solenoid for enhanced polarization preservation and installation of a novel pulsed electromagnetic vacuum valve. The output beam of the source is actually an  $H^-$  beam where the protons have been polarized. The upgraded source reliably produces  $H^-$  beam of up to  $10^{12}$  particles in a  $300 \mu\text{s}$  pulse. Beam polarization of 82-84% has been achieved out of the source as measured by high precision carbon target polarimeter at the end of the 200 MeV Linac. The source can change spin orientation (between vertically up and down) on a pulse to pulse basis, allowing for an arbitrary pattern of spins along the RHIC bunch train. The high per pulse intensity (a factor of 3 higher than what is required in RHIC) enables various *scraping* schemes in the downstream accelerators that reduce the intensity by scraping away high amplitude particles (transversely and longitudinally) but preserve a high brightness core with an intensity and emittance optimized for RHIC operations.

After creation in the source and acceleration in the 200 MeV Linac (to a total energy of 1.1

159 GeV), the  $H^-$  beam is injected into the Booster synchrotron via a charge-exchange injection  
160 process which strips the electrons with a carbon foil and produces the proton beam. From  
161 there the beam is accelerated in three successive synchrotrons. First they are accelerated  
162 in the Booster to 2.3 GeV, then in the Alternating Gradient Synchrotron (AGS) to 23 GeV  
163 and finally in the RHIC rings themselves to a top energy of 255 GeV. The magnetic fields  
164 that produce the steering and focusing necessary to confine the beam during acceleration in  
165 these machines will also induce motion of the spin vectors of the particles as they precess  
166 about the local field lines. This combination of particle motion and spin motion in a periodic  
167 system like a synchrotron ring opens the possibility of depolarization via resonances. These  
168 resonances occur whenever the natural frequency of the spin motion (characterized as the  
169 number of spin precessions per turn, or spin tune) is equal to some frequency with which  
170 the spin perturbing magnetic fields are sampled. A magnetic field that deflects a proton  
171 by an angle  $\theta$  will precess the spin vector by an angle  $G\gamma\theta$ , where  $G$  is the anomalous  
172 gyromagnetic ratio ( $G = 1.79$  for protons) and  $\gamma$  is the relativistic factor. The spin motion  
173 in magnetic fields is therefore highly energy dependent, so during acceleration from rest to  
174 top energy in RHIC these resonance conditions are met very often.

175 Avoiding these resonances consists largely of manipulating the spin motion to prevent the  
176 resonance conditions from being satisfied. In RHIC this is accomplished with helical dipoles  
177 (so-called Siberian snakes). Two helical dipole magnets in each ring, located diametrically  
178 opposite one another, each provide a full spin-flip (from up to down or vice versa) in a single  
179 passage. The result is that the spin tune is made energy independent (fixed at  $1/2$ ) and  
180 there is complete cancellation of all

181 first order depolarizing resonance terms. Even in the presence of full snakes, however,  
182 depolarization is still possible due to higher order resonances called snake resonances. Min-  
183 imizing the effects of these resonances requires careful control of the betatron tune (the  
184 natural frequency of the transverse particle oscillations) during acceleration. In RHIC this  
185 is done using a fast phase-locked loop tune feedback system, in operations since **Run-XX**  
186 which enables acceleration very close to a low order betatron resonance without loss of in-  
187 tensity or emittance dilution, and which maximizes the distance to the nearest depolarizing  
188 snake resonance. This is an excellent example of developments in accelerator physics and  
189 beam control being driven by the stringent demands of a polarized collider physics program.

190 In the AGS, resonance avoidance is somewhat more complicated because there is not  
191 sufficient space in the magnetic lattice for snakes that provide full spin flip. Instead the  
192 AGS uses two weaker *partial* snakes, which rotate the spin through an angle less than  $180^\circ$   
193 ( $18^\circ$  and  $10^\circ$ , respectively). In this case the spin tune is still energy dependent, but is pre-  
194 vented from taking a small range of values. By careful control of the betatron tune, this  
195 allows the strongest resonance to be avoided. A fast tune jump minimizes the polarization  
196 loss from the many weak residual resonances that are driven by the partial snakes them-  
197 selves. The Booster, by contrast, only has two resonance crossings, which are each handled  
198 by an orbit harmonic correction scheme, rather than specialized snake magnets.

199 The strongest resonances encountered during acceleration are dependent on the beam  
200 emittance, with the larger amplitude particle experiencing greater depolarization. This

201 makes emittance control for a polarized beam particularly important since it affects two com-  
202 ponents of the collider performance: luminosity and polarization. A variety of techniques are  
203 employed in the RHIC accelerator complex to prevent emittance growth. Bunch lengthening  
204 manipulations, like dual harmonic RF schemes (in AGS and Booster) and the addition of the  
205 low frequency 9 MHz RF system in RHIC lower the peak current of the bunches and prevent  
206 emittance increases due to space charge forces and electron cloud buildup. A strong driver  
207 of emittance growth during RHIC polarized proton stores is the nonlinear defocusing force of  
208 each beam upon the other (the beam-beam effect) at the point of collision, which can cause  
209 fast emittance blowup immediately upon steering the beams into one another. Starting in  
210 Run-15, this fast blowup was suppressed by a pair of electron lenses, which provide an equal  
211 but opposite nonlinear force.

212 In addition to preserving the beam polarization, the spin physics program requires spe-  
213 cialized instrumentation and methods to measure aspects of the spin dynamics like the spin  
214 tune and the spin direction at key locations in the ring. RHIC is the first polarized beam  
215 facility to measure the spin tune non-destructively using coherent excitation of the spin mo-  
216 tion. An interleaved sequence of 5 RF dipoles and 4 DC dipoles produces a driven coherent  
217 precession of the particle spin vectors. The pC polarimeters deliver a measurement of the  
218 asymmetry synchronized to the phase of the driving field. The spin tune is then calculated  
219 from the relationship between the driven and measured amplitudes. Since this is an adi-  
220 abatic excitation, it can be reversed without loss of polarization, which allows scans and  
221 repeated measurements at top energy without costly refill times. This integration of accel-  
222 erator physics methods with polarimetry allowed optimization of the accelerator lattice to  
223 optimize the spin tune and its spread.

224 These efforts will continue to be important in the EIC era and a host of further develop-  
225 ments will continue through the end of the RHIC program. The hadron spin dynamics and  
226 equipment in the EIC are more complicated than those of RHIC. The altered IP geometry  
227 produces additional complications for the spin rotation manipulation, including change of  
228 spin direction in the arcs, which will have to be measured, controlled and verified. A more  
229 complete partial snake resonance compensation for the AGS is planned for commissioning  
230 in Run-24. Furthermore, the production and acceleration of polarized Helium-3 requires  
231 extensive source development and use of the full toolkit of resonance avoidance including  
232 re-introduction of methods previously used for polarized protons, like use of an AC dipole  
233 resonance crossing, currently under development in the Booster.

## 234 2.2 Polarimetry of High-Energy Hadron Beams

235 The experiments at RHIC require that the beam polarizations are known with good ac-  
236 curacy. Double helicity asymmetries are typically small ( $O(10^{-3})$ ) and knowledge of the  
237 relative luminosity is the leading uncertainty for these observables. Early estimates have  
238 set a requirement of less than  $\sigma(P)/P = 4\%$  for the relative polarization uncertainty. With  
239 increased interest in large transverse single-spin asymmetries and high luminosity data sets,  
240 the achieved benchmark has been lowered to 1.4% in recent years.

241 The beam polarizations at RHIC are measured through a combination of absolute and



242 fast polarimeters. Both are based on the detection of recoil particles from elastic scattering  
243 at low energies, where a spin-dependent asymmetry is introduced through a spin-flip in the  
244 Coulomb-Nuclear interference region. The determination of the absolute beam polarization  
245 makes use of a polarized atomic hydrogen gas jet target, HJET, resulting in an uncertainty of  
246 a 3–4% over the course a typical RHIC fill (8 hours long). The target polarization is prepared  
247 from a state-of-art atomic beam source in a sextupole magnet system in combination with an  
248 RF-transition unit, which optimizes the atomic target density and polarization,  $P \approx 96\%$ .  
249 The remaining molecular fraction of  $H_2$  in the target has been the major source of uncertainty  
250 in the determination of the beam polarization. The recent significant improvement is a  
251 major step towards the applicability of the existing system at the EIC, where a polarization  
252 uncertainty of 1% or better is required.

253 The HJET is complemented by fast measurements with Carbon fiber targets every few  
254 hours which allow for the tracking of polarization decay from injection to the end of each  
255 RHIC fill. The time-dependent knowledge of the beam polarization is important for the cor-  
256 rect use in the experiments, where the luminosity-weighted polarization can be significantly  
257 different from the time-average value.

258 The ultra-thin Carbon targets scan transversely through the beam bunches during each  
259 set of measurements, thereby providing a picture of the transverse beam polarization profile  
260 itself. While it was originally assumed that the transverse polarization profile is flat, the  
261 measurements show that the polarization indeed peaks in the center of the bunches, in both  
262 the horizontal and vertical directions. This information, again, is essential for the proper  
263 determination of the beam polarizations in collision at the experiments. The convolution of  
264 bunch intensities with the polarization profile results in a higher polarization value than that  
265 measured with the HJET. Maybe even more surprising, a longitudinal polarization profile  
266 of the proton bunches has been measured with the HJET and the Carbon polarimeters  
267 independently. The longitudinal polarization dependence is smaller than the transverse  
268 profile; in addition, it is smaller in the center of the bunch.

269 Polarimetry at RHIC relies on a good understanding of the spin dynamics and the stable  
270 spin direction of the accelerator. The transverse direction of the polarization vector (with  
271 respect to the beam momentum) at the collision points in the experiments is confirmed and  
272 monitored through local polarimetry. The concept is based on neutron production at very  
273 forward directions, measured in the Zero Degree Calorimeters. This method was discov-  
274 ered in the first polarized proton collisions at RHIC and it was essential for the successful  
275 commissioning of the spin rotators for longitudinally polarized experiments.

276 Control and verification of the spin direction is important at the experimental collision  
277 points as well as at the location of the polarimeters. In a lattice like RHIC with full snakes,  
278 the design stable spin direction (the direction about which misaligned spin vectors will precess  
279 from turn to turn), is vertical. Pairs of additional helical dipoles, called spin rotators, that  
280 flank the interaction points of the large experiments in the RHIC ring can rotate the spin  
281 locally, providing longitudinally or radially directed polarization at those specific locations,  
282 while leaving the polarization in the rest of the ring unperturbed. The spin direction can  
283 deviate from the intended orientations for a number of reasons, including errors in the helical

284 dipole fields and misalignment of quadrupoles in the lattice. Over the years, systematic  
285 scans of the helical dipole settings and the beam energy have helped to understand and  
286 characterize the source of these deviations. These efforts were particularly vital in shaping  
287 the response to a failure of two of the four magnets that make up one of the snakes in the  
288 Blue ring in Run-22. Meeting the physics requirements with only the remaining two magnets  
289 required compensation with the other functioning snake, a change in the store energy to  
290 minimize the resulting (and now strongly energy dependent) deviation of the spin direction,  
291 and development of new methods of spin direction measurement. Since the local polarimetry  
292 at STAR and the pC polarimeters are only sensitive to the transverse components of the  
293 stable spin direction, measurements were developed using the spin rotators (at STAR) and  
294 a horizontal orbit angle (at the pC) to rotate the hidden longitudinal direction into the  
295 transverse plane.

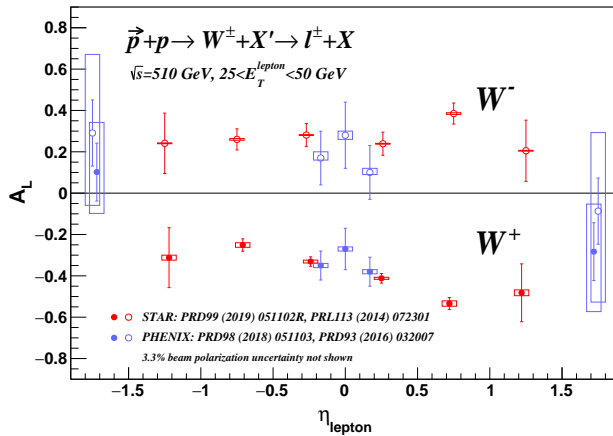
296 Proton polarimetry at RHIC does not only provide vital input for the experiments and  
297 fast feedback to the collider during beam development and regular operations. It has also  
298 delivered surprising results, which further our general understanding of spin-dynamics in  
299 particle accelerators and storage rings. The intricate correlation of the spin tune, stable spin-  
300 direction, and the polarization lifetime has been studied over many years, but it proved to  
301 be of special importance during RHIC Run 22 when one of the Siberian snakes had a partial  
302 failure. Based on past experiences and through ingenious combination of the remaining  
303 snake parts with spin rotators and the polarimeters, it was possible to determine a setup  
304 that showed no significant loss of beam polarization at flattop energy. All of this information  
305 will directly benefit the future Electron-Ion Collider and any other polarized accelerator for  
306 medical or other applications.

### 3 Collinear Proton Structure

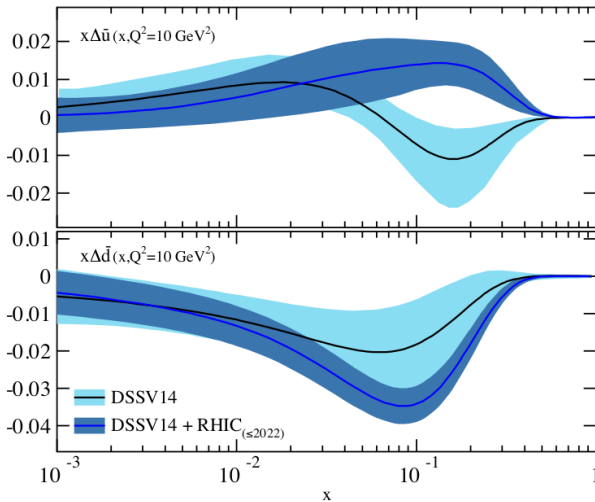
- RHIC high precision longitudinally polarized proton-proton data for a variety of probes and center of mass energies have played a decisive role in constraining the sea-quark and gluon helicity distributions in the proton.
- $W$  production in longitudinally polarized  $pp$  collisions revealed the existence of a flavor asymmetry in the polarization in the sea of light anti-quarks with  $\Delta\bar{u}$  being positive, while  $\Delta\bar{d}$  is negative.
- Collisions at center of mass energies of 200 GeV provided the first evidence that the gluons inside a proton are polarized. Data from the RHIC run in 2009, when included in global analysis, showed that gluons carry approximately 40% of the proton spin in the region where the gluon carries more than 5% of the proton momentum ( $x > 0.05$ ) at  $Q^2 = 10 \text{ GeV}^2$  [1, 3].
- The published and preliminary results based on data collected in 2012, 2013 and 2015 at center of mass energies of 200 and 510 GeV reduce the present uncertainties on gluon helicity  $\Delta g$  even further, providing more insights in the region of momentum fraction  $x$  between about 0.01 to 0.5 of the momentum of a polarized proton.
- STAR  $pp$  and  $pA$  forward di-hadron correlation results pioneered the observation of nonlinear gluon dynamics dependence on the nuclear mass number  $A$  [4]. Higher-precision measurements will be performed with the STAR forward upgrade to further explore nonlinear gluon dynamics. All the studies provide the baseline for searching for gluon saturation at the future EIC.

### 3.1 $W A_L$ and sea quark polarization

The STAR and PHENIX Collaborations have concluded the measurements of the parity-violating spin asymmetry in the production of weak bosons from collisions with one of the proton beams polarized longitudinally [5–10]. In 510 GeV center-of-mass proton-proton collisions at RHIC,  $W^+$  bosons are produced primarily in the interactions of  $u$  quarks and  $\bar{d}$  antiquarks, whereas  $W^-$  bosons originate from  $d$  quarks and  $\bar{u}$  antiquarks. The longitudinal single spin-asymmetry ( $A_L$ ) measurements of the decay positrons provide sensitivity to the  $u$  quark and  $\bar{d}$  helicities in the proton, whereas the decay electrons do so for the  $d$  and  $\bar{u}$  helicities. Combined, they make it possible to delineate the light quark and antiquark polarizations in the proton by flavor.



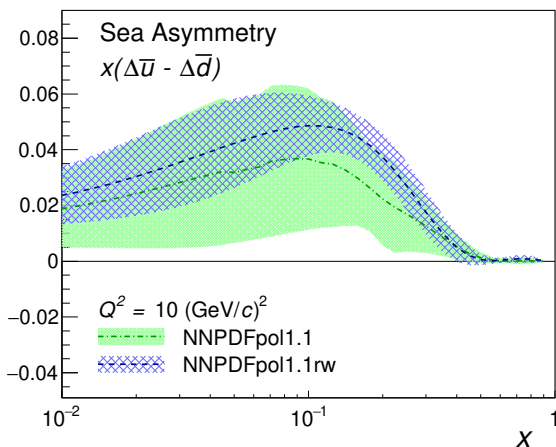
**Figure 2:** Longitudinal single-spin asymmetries,  $A_L$ , for  $W$  production as a function of the lepton pseudorapidity,  $\eta_{\text{lepton}}$ , for the combined STAR and PHENIX data samples [7–10].



**Figure 3:** The impact of the RHIC  $W A_L$  results on  $\bar{u}$  (top) and  $\bar{d}$  (bottom) polarizations as a function of  $x$  at a scale of  $Q^2 = 10 \text{ GeV}^2$ . The black curves with the  $1\sigma$  uncertainty bands marked in light blue show the results from the DSSV14 global fit [11] and the blue curves with  $1\sigma$  uncertainty bands in dark blue show the results for the new preliminary DSSV fit [2] including the RHIC  $W$  data [7, 8, 10].

These measurements shed light on understanding of the light quark polarizations – one of the two initial motivations for the spin-physics program at RHIC. The data, shown in Fig. 2, are the final results from STAR and PHENIX on this topic [7–10] that combine all the published data obtained in 2011, 2012, and 2013. The impact of the RHIC  $W$

342 data on the sea quark helicity distributions  $\Delta\bar{u}$  and  $\Delta\bar{d}$  is presented in Fig. 3. The plot  
 343 shows the impact of the RHIC  $W$  data [7, 8, 10] from the new global fit by the DSSV group  
 344 including also the recent jet, dijet, and pion data [12–18] (that constrain mostly the gluon  
 345 helicity). The sea quark  $\bar{u}$  helicity  $\Delta\bar{u}$  is now known to be positive and  $\Delta\bar{d}$  is negative.  
 346 The STAR 2013 data [10] were also used in the reweighting procedure with the publicly  
 347 available NNPDFpol1.1 PDFs [19]. The results from this reweighting, taking into account  
 348 the total uncertainties of the STAR 2013 data and their correlations, are shown in Fig. 4  
 349 as the blue hatched bands. The NNPDFpol1.1 uncertainties are shown as the green bands  
 350 for comparison. As seen from the plot, the data have now reached a level of precision that  
 351 makes it possible, for the first time, to conclude that there is a clear asymmetry between the  
 352 helicity distribution of  $\bar{u}$  and  $\bar{d}$ , and it has the opposite sign from the  $\bar{d}/\bar{u}$  flavor asymmetry  
 353 in the unpolarized sea.

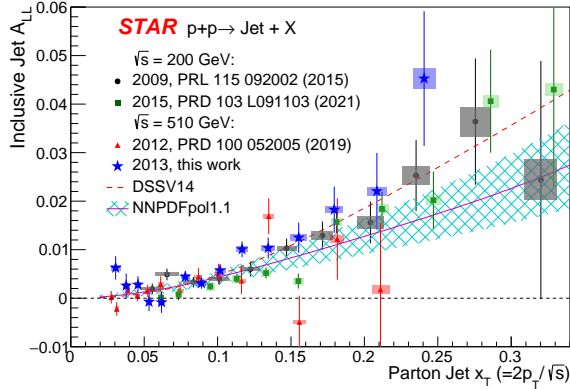


**Figure 4:** The difference of  $\bar{u}$  and  $\bar{d}$  polarizations as a function of  $x$  at a scale of  $Q^2 = 10 \text{ GeV}^2$  before and after NNPDFpol1.1 [19] reweighting with STAR 2013  $W A_L$  [10]. The green band shows the NNPDFpol1.1 results [19] and the blue hatched band shows the corresponding distribution after the STAR 2013  $W$  data are included by reweighting.

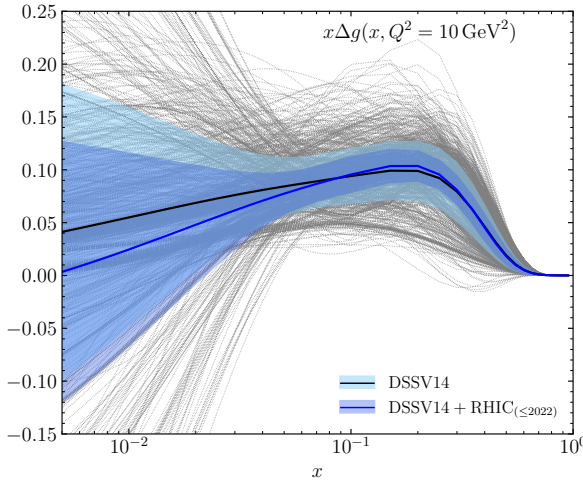
### 354 3.2 Double helicity asymmetries $A_{LL}$ and gluon polarization

355 The measurement of the gluon polarization inside protons has been a major emphasis of the  
 356 longitudinally polarized RHIC program. At RHIC, gluon polarization can be accessed by  
 357 measurements of the spin-dependent rates of production of jets [13–15, 20–23], dijets [12–15,  
 358 24],  $\pi^0$ s and charged pions, [17, 18, 25–31], and direct photons [32]. Data from the RHIC run  
 359 in 2009 have for the first time shown that gluons inside a proton are polarized with a strong  
 360 constraint from the jet data at a center-of-mass energy of  $\sqrt{s} = 200 \text{ GeV}$  [11, 19]. Perturbative  
 361 QCD analyses [11, 19] of the world data, including 2009 inclusive jet and  $\pi^0$  results, at next-  
 362 to-leading order (NLO) precision, suggest that gluon spins contribute  $\simeq 40\%$  to the spin of  
 363 the proton for gluon fractional momenta  $x > 0.05$  at a scale of  $Q^2 = 10 (\text{GeV}/c)^2$ . Results for  
 364 dijet production provide a better determination of the functional form of  $\Delta g(x)$ , compared  
 365 to inclusive observables, because of better constraints on the underlying kinematics [33].

366 Recent STAR results [13–15] and preliminary results [16, 34] on longitudinal double-spin  
 367 asymmetries of inclusive jet and dijet production at center-of-mass energies of 200 GeV (run  
 368 2015) and 510 GeV (runs 2012 and 2013) at mid and intermediate rapidity complement



**Figure 5:** STAR results on inclusive jet  $A_{LL}$  versus  $x_T$  at  $\sqrt{s} = 200$  GeV [14, 23] and 510 GeV [13, 15] at mid-rapidity from data collected in years 2009-2015, and evaluations from DSSV14 [11] and NNPDFpol1.1 (with its uncertainty) [19] global analyses. The vertical lines are statistical uncertainties. The boxes show the size of the estimated systematic uncertainties. Scale uncertainties from polarization (not shown) are  $\pm 6.5\%$ ,  $\pm 6.6\%$ ,  $\pm 6.4\%$  and  $\pm 6.1\%$  from 2009 to 2015, respectively. Source: [15].

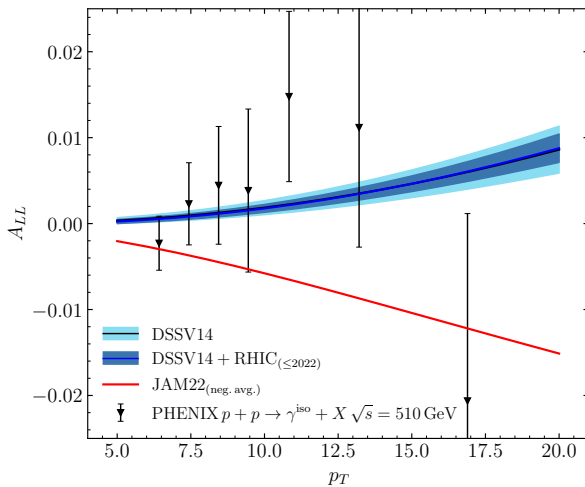


**Figure 6:** The impact of the recent jet and dijet [12–16], pion [17, 18] and  $W$  [7, 8, 10] data on the  $x$ -dependence of the gluon helicity distribution at  $Q^2 = 10 \text{ GeV}^2$  based on the global fit by the DSSV group. The black curve with the  $1\sigma$  uncertainty light blue band illustrates the DSSV14 results [11], while the blue curve with  $1\sigma$  uncertainty band in dark blue [2] shows the results after the inclusion of the new data.

369 and improve the precision of previous STAR measurements. Figure 5 shows recent STAR  
 370 results on inclusive jet  $A_{LL}$  versus  $x_T = 2p_T/\sqrt{s}$  at  $\sqrt{s} = 200$  GeV and 510 GeV at mid-  
 371 rapidity from data collected in years 2009-2015, and evaluations from the DSSV14 [11] and  
 372 NNPDFpol1.1 [19] global analyses. The overall impact of the recent jet and dijet [12–16],  
 373 pion [17, 18] and  $W$  [8, 10] data on the  $x$ -dependence of the gluon helicity distribution at  
 374  $Q^2 = 10 \text{ GeV}^2$  based on the global fit by the DSSV group is presented in Fig. 6. The truncated  
 375 moment of the gluon helicity from the new DSSV evaluations [2] at  $Q^2 = 10 (\text{GeV}/c)^2$   
 376 integrated with the range of  $x \in (0.001, 0.05)$  is 0.173(156) and in the range of  $x \in (0.05, 1)$   
 377 is 0.218(27) (at 68% C.L.), which can be seen in the left panel of Fig. 1.

378 The truncated moment of the gluon helicity integrated from  $x = 0.0071$  to 1 at  $Q^2 =$   
 379  $10 (\text{GeV}/c)^2$  from the recent JAM global QCD analysis [35] including a subset of RHIC data,  
 380 i.e., STAR inclusive jet results, and assuming the SU(3) flavor symmetry and PDF positivity  
 381 is 0.39(9). Authors of [35] also discuss the possibility of the solution with negative gluon  
 382 contribution if the PDF positivity constraint is removed from the global fit. They argue that  
 383 there is no fundamental theoretical requirement for PDF to be positive at all values of  $x$ , and

384 therefore it would be highly desirable to have an observable which is linearly sensitive to gluon  
 385 helicity distribution. Direct photons coming mainly from the quark-gluon Compton process  
 386 and dijets narrowing down the parton kinematics are ideal probes to distinguish between  
 387 positive and negative gluon helicity solutions. Figure 7 demonstrates the preference of posi-  
 388 tive solution with the PHENIX direct photon  $A_{LL}$  data [32]. Figure 8 shows that the STAR  
 389 dijet data [15] also strongly disfavors distributions with large and negative gluon helicities.  
 390 In the plot the asymmetries  $A_{LL}$  are presented for four dijet event topologies, namely, with  
 391 forward-forward jets (top left), forward-central jets (top right), central-central jets (bottom  
 392 left), and forward-backward jets (bottom right), where forward jet rapidity is  $0.3 < \eta < 0.9$ ,  
 393 central jet rapidity is  $|\eta| < 0.3$ , and backward jet rapidity is  $-0.9 < \eta < -0.3$ . The forward-  
 394 forward and forward-central configurations probe the most asymmetric collisions down to  
 395  $x \simeq 0.015$ . The forward-forward and central-central events probe collisions with  $|\cos \theta^*|$   
 396 near zero, whereas forward-central and forward-backward events are more sensitive to larger  
 397  $|\cos \theta^*|$ , where  $\theta^*$  is the scattering angle in the center-of-mass frame of scattering partons.  
 398 In both Figs. 7 and 8, the DSSV14 calculations are plotted as the black curves with the  $1\sigma$   
 399 uncertainty bands marked in light blue. The blue curves with  $1\sigma$  uncertainty bands in dark  
 400 blue show the impact of all the data sets included in the new preliminary DSSV fit [2] as in  
 401 Fig. 6. The curves for JAM  $\Delta g < 0$  solution [35] are presented in red.

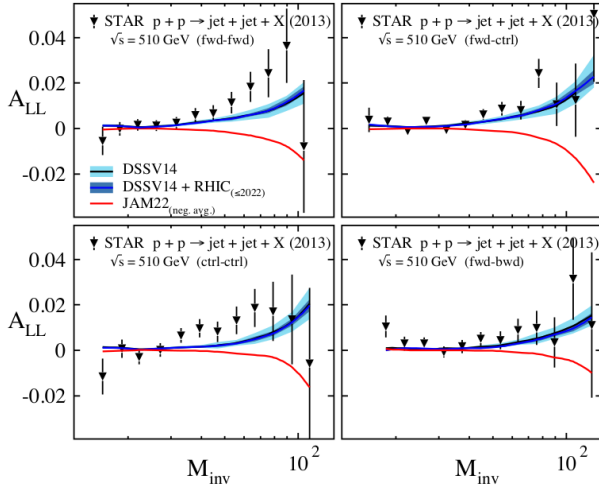


**Figure 7:** PHENIX double-helicity asymmetry  $A_{LL}$  vs  $p_T$  for isolated direct-photon production in polarized  $pp$  collisions at  $\sqrt{s}=510$  GeV at midrapidity [32]. DSSV14 calculation is plotted as the black curve with the  $1\sigma$  uncertainty band marked in light blue. The blue curve with  $1\sigma$  uncertainty band in dark blue shows the impact of all the data sets included in the new preliminary DSSV fit [2] as in Fig. 6. The curve for JAM  $\Delta g < 0$  solution [35] was calculated by W.Vogelsang.

### 3.3 Nonlinear QCD effects

403 To understand where the saturation of gluon densities sets in, whether there is a simple  
 404 boundary that separates this region from that of more dilute quark-gluon matter, is one of  
 405 the most important physics cases of the RHIC Cold QCD program and future EIC.

406 It is well known that PDFs grow rapidly at small- $x$ . The power-law growth of the  
 407 gluon density can be explained by gluon splitting, which leads to a linear evolution of gluon  
 408 dynamics. But if one imagines how such a high number of small- $x$  partons would fit in the



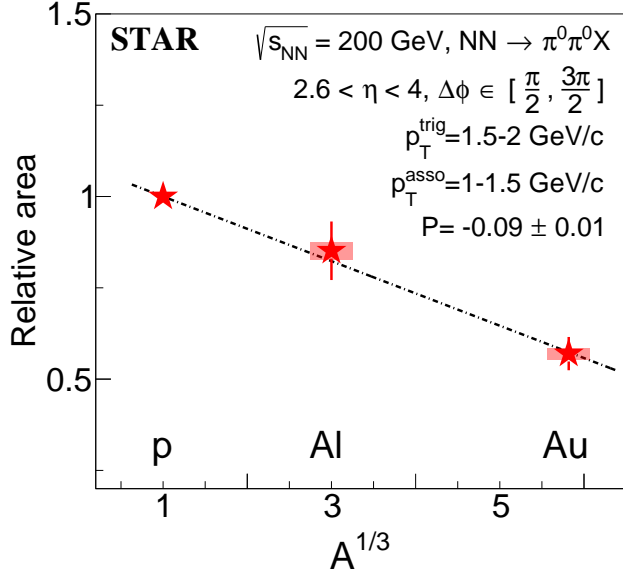
**Figure 8:** STAR double-helicity asymmetries  $A_{LL}$  for dijet production vs dijet invariant mass  $M_{inv}$  in polarized  $pp$  collisions at  $\sqrt{s}=510$  GeV at midrapidity from 2013 data set [15]. DSSV14 evaluation [11] is plotted as the black curve with the  $1\sigma$  uncertainty band marked in light blue. The blue curve with  $1\sigma$  uncertainty band in dark blue shows the impact of all the data sets included in the new preliminary DSSV fit [2] as in Fig. 6. The red curves show the JAM  $\Delta g < 0$  solution [35] calculated by the DSSV group.

409 (almost) unchanged proton radius, one arrives at the picture that the gluons and quarks  
 410 are packed very tightly in the transverse plane. The typical distance between the partons  
 411 decreases as the number of partons increases, and can get small at low- $x$  (or for a large  
 412 nucleus instead of the proton). In QCD, the black disk limit states that the total hadronic  
 413 cross section cannot grow forever. Thus, the growth of gluon density which can be described  
 414 by BFKL evolution, has to be tamped at some point. At very high density, partons may  
 415 start to recombine with each other on top of splitting. The recombination of two partons  
 416 into one is proportional to the number of pairs of the partons. Therefore the BFKL function  
 417 needs to be modified by adding a nonlinear term of recombination. Saturation is a new  
 418 regime of QCD, where gluon splitting and recombination reach a balance. One can define  
 419 the saturation scale as the inverse of this typical transverse interparton distance. Hence  $Q_s$   
 420 indeed grows with  $A$  and decreasing  $x$ .

421 Collisions between hadronic systems, *i.e.*,  $pA$  and  $dA$  at RHIC provide a window to the  
 422 parton distributions of nuclei at small momentum fraction  $x$  (down to  $10^{-3}$ ). Several RHIC  
 423 measurements have shown that, at forward pseudorapidities (deuteron going direction), the  
 424 hadron yields are suppressed in  $dAu$  collisions relative to  $pp$  collisions in inclusive produc-  
 425 tions [36–39] and di-hadron correlations [39, 40]. However, for the inclusive channel, it was  
 426 indicated that the nuclear modified fragmentation can serve as another interpretation be-  
 427 yond gluon saturation to explain the suppression. The di-hadron correlation measurement  
 428 can provide future test for the saturation physics. For di-hadron correlation in  $dA$ , the  
 429 contributions from double-parton scatterings (DPS) to the  $d+A \rightarrow \pi^0\pi^0X$  cross section are  
 430 suggested as an alternative explanation for the suppression [41] beyond gluon saturation.  
 431 Therefore, it is important to make the same measurements in the theoretically and experi-  
 432 mentally cleaner  $pA$  collisions. Under the color glass condensate (CGC) framework [42–44],  
 433 at a given  $x$ , gluons from different nucleons are predicted to amplify the total transverse  
 434 gluon density by a factor of  $A^{1/3}$  for a nucleus with mass number  $A$ . RHIC 2015  $pp$ ,  $pA$ ,  
 435 and  $pAu$  datasets are ideal to study the  $A$ -dependence by varying the nuclei species.

436 The recent published forward di- $\pi^0$  correlation measured by the STAR detector pioneered



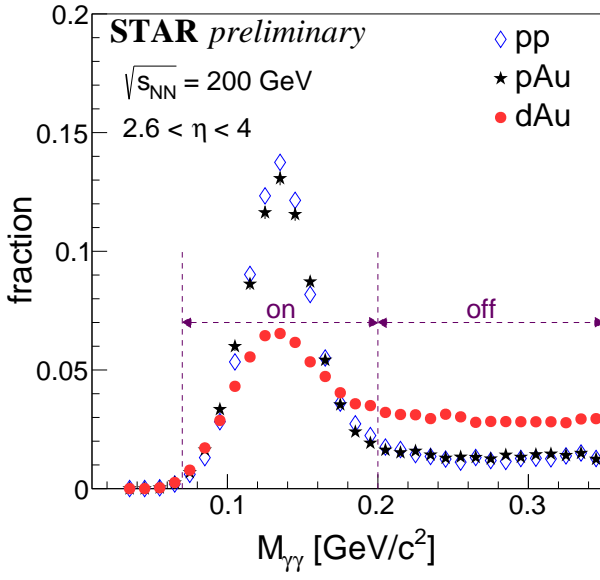


**Figure 9:** At  $p_T^{trig} = 1.5-2$  GeV/ $c$  and  $p_T^{asso} = 1.5-2$  GeV/ $c$ , a linear dependence of the suppression of back-to-back  $\pi^0$  correlation as a function of  $A^{1/3}$  is observed within the uncertainties, the slope ( $P$ ) is found to be  $-0.09 \pm 0.01$ . The plot is from [4].

437 the observation of the dependence of nonlinear gluon dynamics on the nuclear mass number  
 438  $A$  [4], see Fig. 9. The area is extracted by a Gaussian fit of the back-to-back correlation  
 439 measured from each collision system. The area ratio of  $pA/pp$  presents the relative yields of  
 440 back-to-back di- $\pi^0$ s in  $pA$  with respect to  $pp$  collisions. The area ratio in  $pAu$  over  $pp$  is about  
 441 50% indicating a clear suppression of back-to-back di- $\pi^0$  correlation in  $pAu$  compared to  $pp$   
 442 collisions. The same trend but smaller amount of suppression is observed in  $pAl$  collisions.  
 443 The suppression is found to scale with  $A$  and linearly dependent on  $A^{1/3}$ . The extracted slope  
 444 from the linear dependence will be critical input for the gluon saturation model in CGC.  
 445 Meanwhile, STAR revisited the same measurement for  $dAu$  collisions. It was predicted by  
 446 comparing the forward di- $\pi^0$  correlation in  $pp$ ,  $pAu$ , and  $dAu$  collisions, one can access the  
 447 contribution from DPS [41].

448 For RHIC 2016 data, a large background of  $\pi^0$  identification is found in  $dAu$  collisions,  
 449 in comparison with the  $pp$  and  $pAu$  collisions from 2015 in Fig. 10. The generated combi-  
 450 natoric correlation dominates in  $dAu$  collisions, which makes it very challenging to identify  
 451 the signal correlation. The forward di- $\pi^0$  correlation measurement favors the cleaner  $pA$  col-  
 452 lisions rather than  $dA$  collisions. It emphasizes the importance of measuring the di-hadron  
 453 correlation in  $pA$  collisions with the STAR Forward Upgrade in the future run 2024. The  
 454 higher delivered integrated luminosity for this run together with the Forward Upgrade will  
 455 enable one to study more luminosity-hungry processes and/or complementary probes to the  
 456 di- $\pi^0$  correlations, i.e. di-hadron correlations for charged hadrons, photon-jet, photon-hadron  
 457 and di-jet correlations. Utilizing the forward tracking systems, the background for particle  
 458 identification will be much suppressed with respect to the current di- $\pi^0$  studies.

459 These results are crucial for the equivalent measurements at the EIC, which are planned  
 460 at close to identical kinematics, because only if non-linear effects are seen with different  
 461 complementary probes, i.e.,  $ep$  and  $pA$ , one can claim a discovery of saturation effects and  
 462 their universality. Therefore it is imperative that analysis activities related to the unpolarized



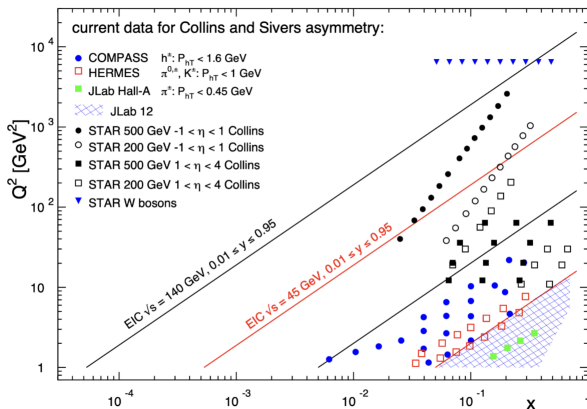
**Figure 10:** The reconstructed invariant mass of two photons for  $pp$ ,  $pAu$ , and  $dAu$  collisions. The background is higher in  $dAu$  collisions in comparison with the  $pp$  and  $pAu$  collisions.  $pp$  and  $pAu$  collisions are similar.

463 Cold-QCD program continue to be supported throughout the upcoming years.

## 4 Three-dimensional Structure

STAR opened new territory in studying the 3D structure of the proton in the region of momentum fractions down to  $x \sim 0.01$  and high  $Q^2$ , a region not probed by prior experiments. See Fig. 11.

- The collected unique sets of transversely polarized data in  $pp$  and  $pA$  collisions, including the most recent campaign with the forward upgrade, will be finalized with the 2024 RHIC run.
- To accomplish the scientific mission of the transverse spin program, it is imperative that analysis activities continue to be supported throughout the upcoming years. These activities offer discovery potential of their own, and they are critical for properly interpreting data from the future Electron-Ion Collider.
- STAR pioneered the novel use of jets and their substructure to study initial and final state transverse momentum dependent (TMD) effects in polarized  $pp$  collisions. For example, the measured single-spin asymmetries of hadrons in jets probe the quark transversity distribution and Collins TMD fragmentation function, and the single-spin asymmetry of dijet opening angle is sensitive to the Sivers TMD parton distribution.
- STAR has also measured quark transversities via dihadron interference fragmentation functions. The results from early measurements have been included in a global analysis, and found to provide significant constraints. Ongoing analysis of more recent STAR data, together with the data that STAR will record during 2024, will provide far more stringent constraints.
- Substantial progress on the large forward transverse single-spin asymmetry puzzle has been made. The  $A_N$  of the isolated  $\pi^0$ s was found to be significantly larger than that for non-isolated ones both in  $pp$  and  $p+A$  collisions at STAR. The  $A_N$  for  $\pi^0$ s at large  $x_F$ , far forward pseudorapidity ( $\eta > 6$ ), and  $p_T < 1$  GeV/ $c$  at RHICf was found to be comparable to that at the same  $x_F$ , but with  $2.5 < \eta < 4$  and  $p_T > 2$  GeV/ $c$  at STAR.



**Figure 11:** The  $x$ - $Q^2$  probed with data from the future EIC and Jlab-12 GeV as well as the current SIDIS data and the jet and  $W$ -boson data from RHIC. All data are sensitive to the Sivers function and transversity times the Collins FF in the TMD formalism.

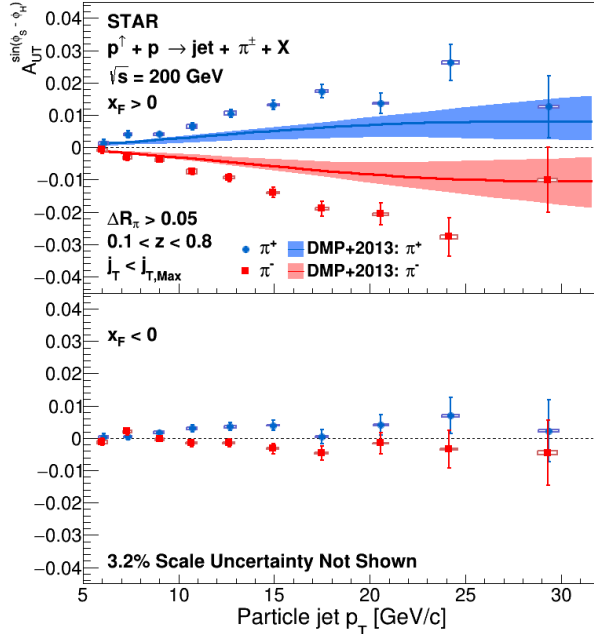
490  
491  
492  
493  
  
494  
495  
496  
497  
498  
499  
  
500  
501  
502  
  
503  
504  
505  
506  
507  
  
508  
509  
510  
511  
512

The  $A_N$  for electromagnetic jets was found to be small but non-zero, which provided significant constraints to the quark Sivers function. The  $A_N$  for forward diffractive EM-jets has been measured and found not to be the source of the large  $A_N$ . In fact, it favors a negative contribution.

- Transverse single-spin asymmetry  $A_N$  of weak bosons, sensitive to the Sivers TMD function, has been probed at STAR. With the increased precision provided by 2017 data, STAR found smaller asymmetries than were suggested by 2011 data. As a result, the increased statistics of the 2022 dataset are critical to improve the precision of our asymmetry measurements in order to provide a conclusive test of the Sivers' function sign change.
- PHENIX has measured transverse single-spin asymmetries at mid-rapidity that provide constraints on the twist-3 correlation functions, including the first RHIC result of direct photon  $A_N$  and high precision neutral meson  $A_N$ .
- PHENIX and STAR have both measured the nuclear dependence of the forward inclusive hadron single-spin asymmetries. PHENIX finds a strong nuclear dependence for positive hadrons at  $1.2 < \eta < 2.4$ , whereas STAR finds a weak nuclear dependence for  $\pi^0$  at  $2.7 < \eta < 3.8$ . Neither the origin of the nuclear dependence, nor the difference between the PHENIX and STAR results is well understood at this time.
- Transverse single-spin asymmetry of exclusive  $J/\psi$  photoproduction in ultra-peripheral collisions is expected to directly probe the generalized parton density (GPD) distribution. The STAR forward detector and data beyond 2022 can measure unique kinematic phase space, e.g., close to the threshold production energy of  $J/\psi$ , where a large asymmetry signal is expected.

## 513 4.1 Studies of initial and final state TMD effects with jets

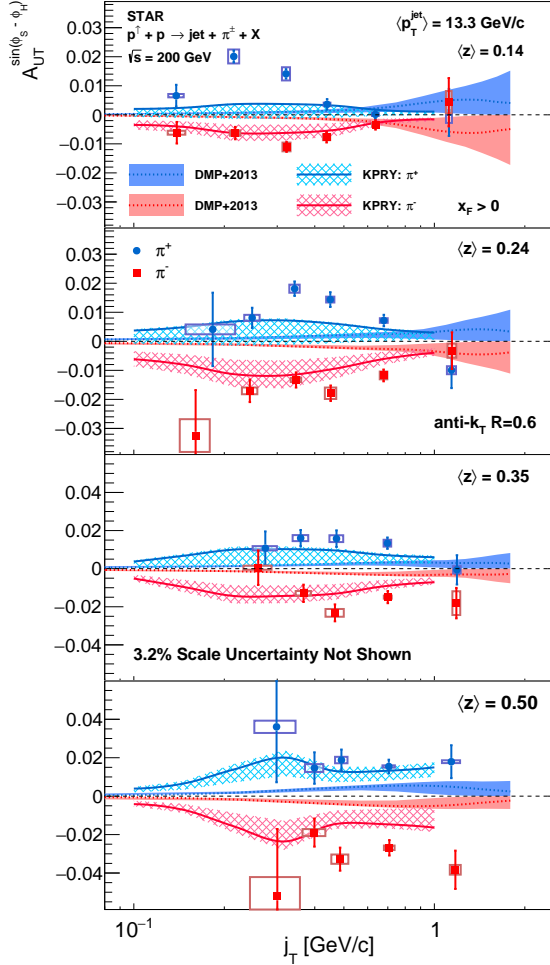
514 STAR has pioneered the novel use of jets and their substructure to study initial state and  
 515 final state TMD effects in polarized  $pp$  collisions.



**Figure 12:** Collins asymmetry plotted for identified  $\pi^+$  (blue) and  $\pi^-$  (red) particles as a function of jet  $p_T$  for jets that scatter forward relative to the polarized beam ( $x_F > 0$ ) in the top panel and those that scatter backward ( $x_F < 0$ ) in the lower panel, extracted from data collected in 2012 and 2015 [45]. The full ranges of both  $z$  and  $j_T$  are integrated over. Theoretical evaluations from [46] with their uncertainties are presented for  $\pi^+$  (blue) and  $\pi^-$  (red). Source: [45].

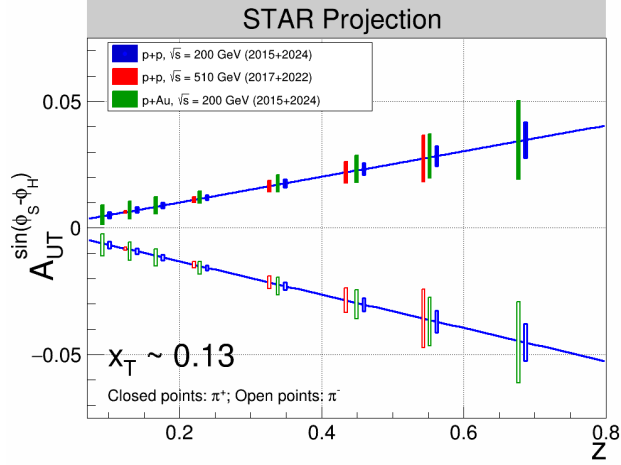
516 The single-spin asymmetries of the azimuthal distribution of identified pions, kaons, and  
 517 protons in high-energy jets measured at STAR probe the *collinear* quark transversity in  
 518 the proton, coupled to the transverse momentum dependent Collins fragmentation function  
 519 [47–49]. This makes  $pp$  collisions a more direct probe of the Collins fragmentation function  
 520 than SIDIS, where a convolution with the TMD transversity distribution enters. The Collins  
 521 asymmetry in  $pp$  collisions is an ideal tool to explore the fundamental QCD questions of TMD  
 522 factorization, universality, and evolution. Figure 12 shows the recent results on combined  
 523 2012 and 2015 Collins asymmetries for charged pions within jets as a function of jet  $p_T$  [45].  
 524 By integrating over the hadron longitudinal and transverse momenta within the jets, Fig. 12  
 525 is sensitive primarily to the quark transversity. The measured asymmetries for jets that  
 526 scatter forward relative to the polarized beam are larger than theoretical predictions [46],  
 527 which are based on the transversity and Collins fragmentation function from SIDIS and  $e^+e^-$   
 528 processes within the TMD approach. Alternatively, the asymmetries can be investigated as  
 529 functions of  $z$ , the fraction of jet momentum carried by the hadron, and  $j_T$ , the momentum of  
 530 the pion transverse to the jet axis, as shown in Fig. 13. This provides a direct measurement  
 531 of the kinematic dependence of the Collins fragmentation function. The  $j_T$  dependence  
 532 appears to vary with  $z$ , contrary to the assumptions of most current phenomenological  
 533 models [47–49]. STAR has also published Collins asymmetry measurements from a smaller  
 534 500 GeV data set collected in 2011 [50]. While statistics are limited, the results are consistent  
 535 with those at 200 GeV for overlapping  $x_T$ , despite sampling  $Q^2$  that is larger by a factor of 6.

536 Analysis of the higher statistics 510 GeV data collected in 2017 is underway and will provide  
 537 unique insight into the  $Q^2$  evolution of the Collins TMD fragmentation function. Concurrent  
 538 with the Collins effect measurements, STAR has also measured azimuthal modulations that  
 539 are sensitive to the twist-3 analogs of the quark and gluon Sivers functions and to linear  
 540 polarization of gluons in transversely polarized protons [45, 50]. Analysis is also underway  
 541 to determine the unpolarized TMD fragmentation functions.



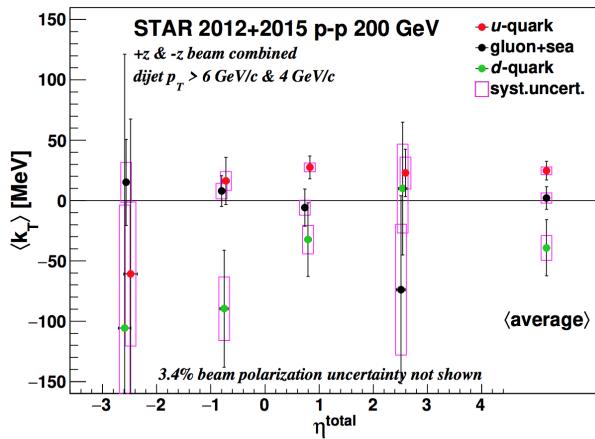
**Figure 13:** Collins asymmetry plotted for identified  $\pi^+$  (blue) and  $\pi^-$  (red) particles as a function of  $j_T$  for four separate bins of hadron  $z$ , in jets with  $p_T > 9.9$  GeV/ $c$  and  $0 < \eta < 0.9$ . Theoretical evaluations from [48] and [46] are also shown. Source: [45]

542 As shown in Fig. 11, data from 200 GeV  $pp$  collisions from the upcoming run 2024 with  
 543 the STAR Forward Upgrade will interpolate between the coverage that we will achieve with  
 544 the forward data collected at 510 GeV in 2022 at high- $x$  and the data at low- $x$  from the  
 545 STAR mid-rapidity detectors. Overall, all STAR data will provide valuable information  
 546 about evolution effects and, with the projected statistical precision presented in Fig. 14, will  
 547 establish the most precise benchmark for future comparisons to  $ep$  data from the EIC. It is  
 548 also important to recognize that the hadron-in-jet measurements with the STAR Forward  
 549 Upgrade will provide a very valuable experience detecting jets close to beam rapidity that  
 550 will inform the planning for future jet measurements in similar kinematics at the EIC.



**Figure 14:** Projected statistical uncertainties for STAR Collins asymmetry measurements at  $0 < \eta < 0.9$  in  $pp$  at  $\sqrt{s} = 200$  and 510 GeV and  $p+Au$  at  $\sqrt{s_{NN}} = 200$  GeV. The points have arbitrarily been drawn on the solid lines, which represent simple linear fits to the STAR preliminary 200 GeV  $pp$  Collins asymmetry measurements from 2015. (Note that only one bin is shown spanning  $0.1 < z < 0.2$  for 510 GeV  $pp$ , whereas three bins are shown covering the same  $z$  range for the 200 GeV measurements.)

551 STAR also has the unique opportunity to extend the Collins effect measurements to  
 552 nuclei. This will provide an alternative look at the universality of the Collins effect in  
 553 hadron production (by dramatically increasing the color flow options of the sort that have  
 554 been predicted to break factorization for TMD PDFs like the Sivers effect [51,52]) and explore  
 555 the spin dependence of the hadronization process in cold nuclear matter. STAR collected a  
 556 proof-of-principle dataset during the 2015  $p+Au$  run that is currently under analysis. Those  
 557 data will provide the first estimate of medium-induced effects. However, the small nuclear  
 558 effects seen by STAR for forward inclusive  $\pi^0$   $A_N$  [53] indicate that greater precision will  
 559 likely be needed. Figure 14 shows the projected statistical uncertainties for the  $p+Au$  Collins  
 560 asymmetry measurement at  $\sqrt{s_{NN}} = 200$  GeV from 2015 and 2024 data, compared to those  
 561 for  $pp$  at the same energy.



**Figure 15:** Preliminary results of the average transverse momentum  $\langle k_T \rangle$  for individual partons, inverted using parton fractions from simulation and tagged  $\langle k_T \rangle$  in data, plotted as a function of summed pseudorapidities of the outgoing jets  $\eta_{total} \sim \log(x_1/x_2)$ . (Positive  $\eta_{total}$  represents dijets emitted in the direction of the polarized beam.) The rightmost points represent the average of all the  $\eta_{total}$  bins. The systematic uncertainty in  $\eta_{total}$  is set to be non-zero to improve the visibility of the error bars. Source: [54].

562 Another example of utilizing jets to unravel the internal TMD structure of the proton is  
 563 the measurement of the asymmetry of the spin-dependent ‘tilt’ of the dijet opening angle,  
 564 which is sensitive to the Sivers TMD PDF. For transversely polarized protons, the Sivers  
 565 effect probes whether the transverse momentum  $\vec{k}_T$  of the constituent quarks is preferentially

566 oriented in a direction perpendicular to both the proton momentum and its spin. Figure 15  
567 shows the first-ever observation of the Sivers effect in dijet production from the 200 GeV  
568 transverse spin data that STAR recorded in 2012 and 2015 [54]. The jets are sorted accord-  
569 ing to their net charge  $Q$ , yielding jet samples with enhanced contributions from  $u$  quarks  
570 (positive  $Q$ ) and  $d$  quarks (negative  $Q$ ), with a large set near  $Q = 0$  dominated by gluons.  
571 Simple kinematics allow for conversion from the spin-dependent ‘tilt’ of the dijet pair to a  
572 value of  $k_T$  on an event-by-event basis. Finally, the results are unfolded for the  $k_T$  of indi-  
573 vidual partons. Such measurements are crucial to explore questions regarding factorization  
574 of the Sivers function in dijet hadroproduction [51, 52, 55, 56]. New data to be taken in 2024  
575 will reduce the uncertainties for the region of summed pseudorapidities of the outgoing jets  
576  $|\eta_3 + \eta_4| < 1$  by about a factor of two. The increased acceptance from the iTPC will reduce  
577 the uncertainties at  $|\eta_3 + \eta_4| \approx 2.5$  by a much larger factor, while the Forward Upgrade will  
578 enable the measurements to be extended to even larger values of  $|\eta_3 + \eta_4|$ . When combined  
579 with the 510 GeV data from Run-17 and Run-22, the results will provide a detailed mapping  
580 *vs. x* for comparison to results for Sivers functions extracted from SIDIS, Drell-Yan, and  
581 vector boson production.

## 582 4.2 Transversity from di-hadron interference fragmentation func- 583 tions

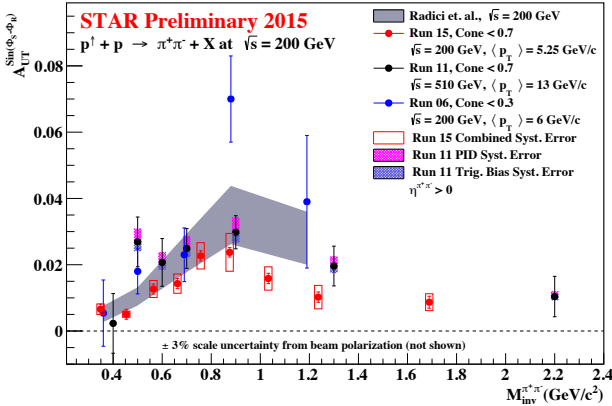
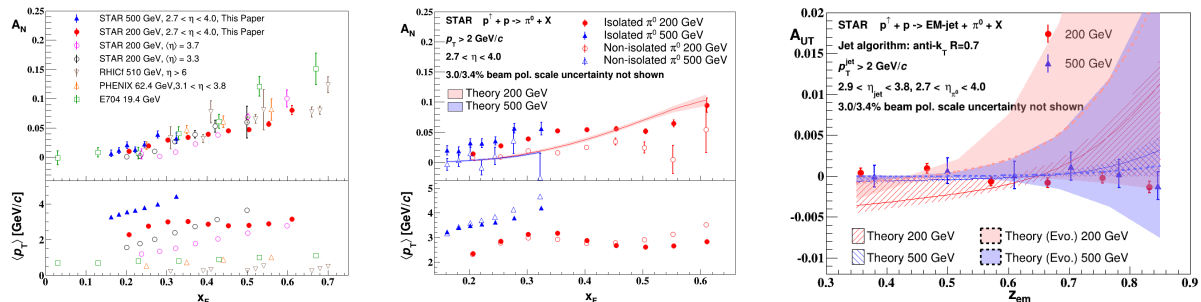


Figure 16: A comparison of STAR published [57, 58] and preliminary [59] IFF asymmetries vs. dipion invariant mass to predictions from the global analysis of [60], which only included the 200 GeV data from 2006 in the fit. The  $p_T$  bins at 200 and 500 GeV have been chosen to sample similar values of  $x_T = 2p_T/\sqrt{s}$ . Source: [59].

584 STAR has also measured quark transversity via dihadron Interference Fragmentation  
585 Functions (IFF) in 200 and 500 GeV  $pp$  collisions [57, 58], as shown in Fig. 16. The IFF is a  
586 collinear observable, so these measurements provide a complementary probe of transversity  
587 relative to the Collins asymmetry measurements that obeys different evolution equations.  
588 The results from the first measurements at 200 GeV, which were based on data recorded  
589 during 2006 [57], have been included together with IFF measurements from SIDIS in a  
590 global analysis [60] that is also shown in Fig. 16. The STAR IFF measurements were found to  
591 provide significant additional constraints on the  $u$ - and  $d$ -quark transversities. The dominant  
592 systematic uncertainties in the global analysis arose from the current lack of knowledge  
593 regarding the unpolarized gluon dihadron fragmentation functions. Analysis on unpolarized



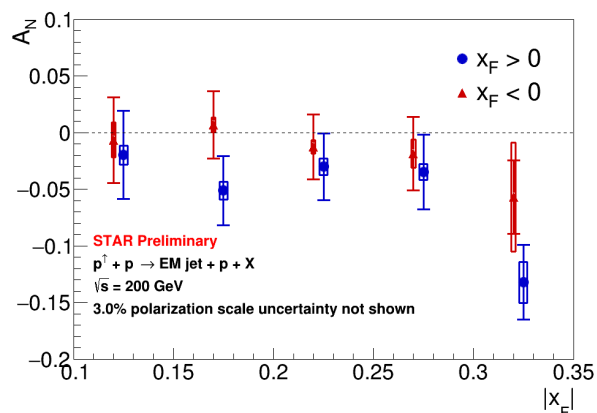


**Figure 17:** Left: Transverse single-spin asymmetry  $A_N$  as a function of  $x_F$  for inclusive  $\pi^0$  in  $pp$  collisions up to RHIC energies of 200 and 510 GeV. Middle:  $A_N$  asymmetries for the isolated and non-isolated  $\pi^0$  in  $pp$  collisions at 200 and 500 GeV. Right: The Collins asymmetry for  $\pi^0$  in an electromagnetic jet for  $pp$  collisions at  $\sqrt{s} = 200$  and 500 GeV. The plots are from [61].

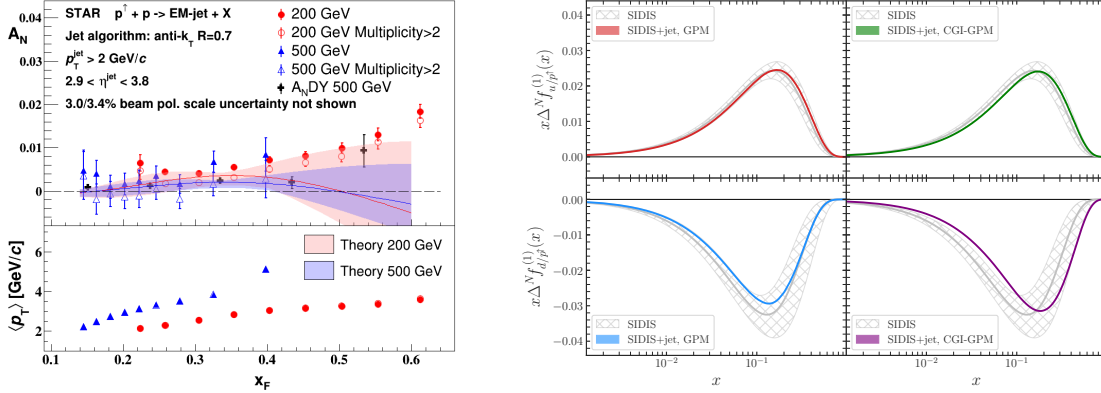
594 IFF function is underway at STAR that will help to reduce these uncertainties. The analysis  
 595 of IFF asymmetries with more recent STAR data taken in 2017 and 2022 at 510 GeV, together  
 596 with the data that STAR will record during 2024 at 200 GeV, will provide far more stringent  
 597 constraints on quark transversities than have been obtained to date when they are included  
 598 in future global analyses.

### 599 4.3 Transverse single-spin asymmetry in the forward region

600 STAR measurements have demonstrated the persistence of sizeable transverse single-spin  
 601 asymmetries  $A_N$  for forward  $\pi^0$  production at RHIC energies up to 510 GeV with a weak  
 602 energy dependence (see left panel of Fig. 17), where different QCD mechanisms including  
 603 the high twist effect, TMD effects like the Sivers or Collins effects, and diffractive processes  
 604 could all contribute. It is thus important to study different effects separately for a full  
 605 understanding of the underlying mechanism, and a series of measurements were performed  
 606 in  $pp$  collisions at both 200 and 500 GeV and in  $p+A$  collisions at STAR [53, 61, 62].



**Figure 18:** Transverse single-spin asymmetry for diffractive EM-jet as a function of  $x_F$  in transversely polarized proton-proton collisions at  $\sqrt{s} = 200$  GeV [62]. The blue points are for  $x_F > 0$ . The red points are for  $x_F < 0$  with a constant shift of -0.005 along x-axis for clarity. The rightmost points are for  $0.3 < |x_F| < 0.45$ .



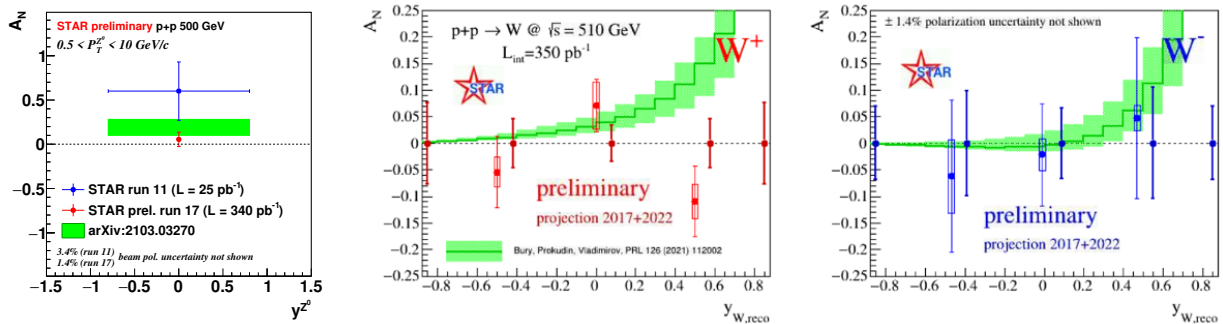
**Figure 19:** Left: Transverse single-spin asymmetry as a function of  $x_F$  for electromagnetic jets in transversely polarized proton-proton collisions at  $\sqrt{s} = 200$  and 500 GeV [61]. Right: Comparison between the Sivers first  $k_{\perp}$ -moments from SIDIS data and their reweighted SIDIS+jet (data from STAR) in two frameworks: GPM and CGI-GPM [63].

607 Firstly, the topological dependence of the  $\pi^0$   $A_N$  was studied, and the  $A_N$  of the isolated  
608  $\pi^0$ 's (meaning no other particles around) are significantly larger than the non-isolated ones,  
609 as shown in the middle panel of Fig. 17. Consistent results were obtained in both  $pp$   
610 and  $pA$  collisions with very weak  $A$  dependence in  $pA$  [53, 61]. This triggered discussions  
611 on the possible contribution from the diffractive process, which motivated a measurement  
612 of  $A_N$  for singly and double diffractive events, utilizing the STAR Roman Pot detectors  
613 to tag diffractive processes with scattered protons close to the beamline. Figure 18 shows  
614 the preliminary results for forward diffractive EM-jet  $A_N$  as a function of  $x_F$  at  $\sqrt{s} =$   
615 200 GeV [62]. The results favor a non-zero negative  $A_N$  with  $3.3\sigma$  significance, so these  
616 diffractive processes are most probably not the source of the large positive  $A_N$  of  $\pi^0$ . The  
617 negative contribution from diffractive jets is not currently described by theory.

618 In studying the contribution from the final-state effect, STAR also measured the Collins  
619 asymmetry of  $\pi^0$  in an electromagnetic jet, which is shown in the right panel of Fig. 17.  
620 The measured Collins asymmetry was consistent with zero, in agreement with a theoretical  
621 prediction based on collinear twist-3 factorization, resulting from significant cancellation  
622 between Collins effects of different quark flavors [48].

623 In a closely related study, RHICf has measured  $A_N$  for neutral pions in 510 GeV  $pp$   
624 collisions at very large pseudorapidity ( $\eta > 6$ ), very large  $x_F$  (up to 0.8), and  $p_T < 1$   
625 GeV/c [64]. The asymmetries that they found are similar to those at comparable  $x_F$  and  
626 much higher  $p_T$ , as shown in the left panel of Fig. 17. A very recent calculation [65] based on  
627 diffractive triple Regge exchange provides a very good description of the RHICf  $A_N$  results.

628 Another study is the measurement of the  $A_N$  for inclusive electromagnetic jets, which is  
629 considered only related to the initial-state effect. The results of electromagnetic jet  $A_N$  in  
630 both 200 and 500 GeV  $pp$  collisions are shown in the left panel of Fig. 19. The electromagnetic  
631 jet  $A_N$  was found to increase with  $x_F$ , but the magnitude is much smaller than the  $\pi^0$   $A_N$ .  
632 These data have been included in the recent global fit of the Sivers function [63], and showed



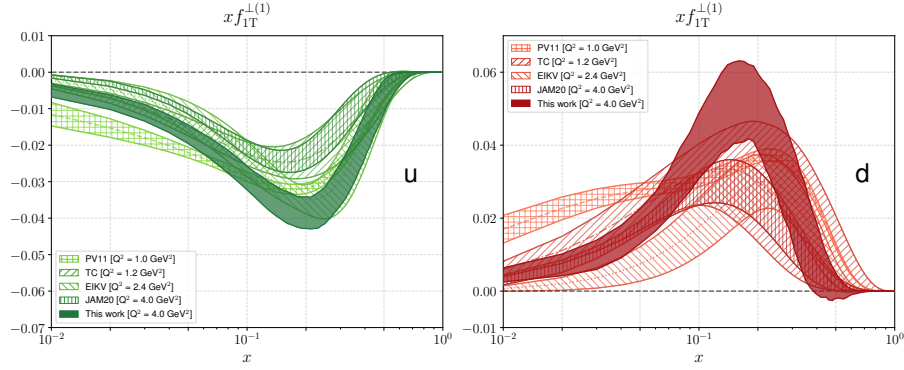
**Figure 20:** Left: Transverse single-spin asymmetry of  $Z^0$  from STAR 2011 and 2017 data. The results are compared with the calculation from [66]. Middle and Right: Transverse single-spin asymmetry of  $W^\pm$  from STAR 2017, and the projected statistical uncertainties from 2017 and 2022 data. The results are compared with calculation from [66] based on the next-to-next-to-next-to leading log ( $N^3LL$ ) accuracy TMD evolution from [67].

633 a significant impact in constraining the Sivers function, as shown in the right panel of Fig. 19.  
 634 With the implementation of the Forward Upgrade at STAR in the  $pp$  and  $pA$  running  
 635 in 2022 and 2024, we will be able to perform measurements with full jets in the forward  
 636 rapidity region and also the Collins asymmetry with charge separated hadrons, which will  
 637 allow a complete understanding of the underlying QCD mechanism.

#### 638 4.4 Transverse single-spin asymmetry of weak bosons

639 Proton-proton collisions at  $\sqrt{s} = 510$  GeV allow STAR to study the evolution and sign  
 640 change of the Sivers function with weak bosons at mid-rapidity ( $-1 < y^{W^\pm/Z^0} < 1$ ). By  
 641 focusing on interactions in which the final state involves only leptons, and hence the trans-  
 642 verse partonic motion must be in the initial state, one can test the predicted sign change  
 643 in  $A_N$  relative to interactions in which these terms must appear in the final state, such as  
 644 SIDIS measurements. Following the low statistics proof-of-principle measurement using the  
 645 2011 data, STAR measured the transverse single-spin asymmetry  $A_N$  for  $W$  and  $Z$  with 2017  
 646 data, which had about 14 times more integrated luminosity.

647 In Fig. 20, the recent preliminary results on  $A_N$  of  $W/Z$  are compared with predictions  
 648 from [66,67] that include STAR 2011 data. The recent global QCD extraction of the Sivers  
 649 function including STAR 2011  $W$  and  $Z$   $A_N$  data from [68] can be found in Fig. 21. With the  
 650 increased precision provided by run 2017, we find smaller asymmetries than were suggested  
 651 by run 2011. As a result, the increased statistics of the 2022 dataset are critical to improve  
 652 the precision of our asymmetry measurements in order to provide a conclusive test of the  
 653 Sivers' function sign change. Projected statistical uncertainties of  $W$   $A_N$  from combined  
 654 2017 and 2022 data can be found in Fig. 20. The figure also illustrates that the improved  
 655 tracking capabilities provided by the STAR iTPC upgrade will allow us to push our mid-  
 656 rapidity  $W^\pm$  and  $Z$  measurements to larger rapidity  $y_{W/Z}$ , a regime where the asymmetries  
 657 are expected to increase in magnitude and the anti-quark Sivers' functions remain largely

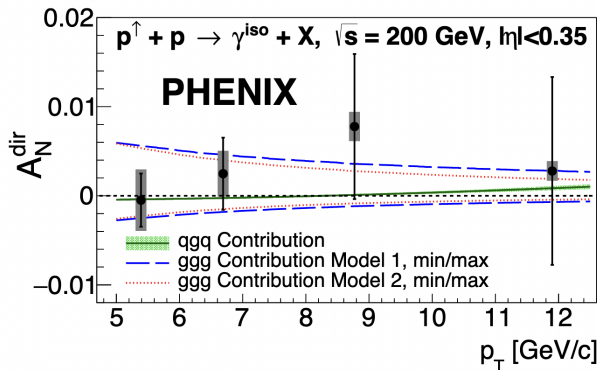


**Figure 21:** The first transverse moment  $x f_{1T}^{\perp(1)}$  of the Siverts TMD as a function of  $x$  for the up (left panel) and down quark (right panel) extracted from world data including STAR 2011  $W/Z$  data. Solid band: the 68% confidence interval obtained in this work at  $Q^2 = 4 \text{ GeV}^2$ . The plot is from [68].

658 unconstrained.

## 659 4.5 Transverse single-spin asymmetries of direct photons and heavy 660 flavor decay leptons

661 PHENIX has reported the first direct photon transverse single-spin asymmetry result at  
662 RHIC [69]. The asymmetry was measured at midrapidity  $|\eta| < 0.35$  in  $pp$  collisions at  
663  $\sqrt{s} = 200 \text{ GeV}$ . Photons do not interact via the strong force, and at this kinematics they are  
664 produced dominantly by the quark-gluon Compton process. Therefore, the measurement  
665 offers a clean probe of gluon dynamics that is only sensitive to initial-state effects. The  
666 asymmetry is shown in Fig. 22 and is consistent with zero to within 1% across the measured  
667  $p_T$  range. The result is also compared with predictions from collinear twist-3 correlation  
668 functions. The solid green curve shows the contribution from  $qqq$  correlation function [70]  
669 while the dashed (blue) and dotted (red) curves are from  $ggg$  correlation functions [71].  
670 Given the small predicted contributions from  $qqq$  correlation functions to the asymmetry,  
671 the result can provide a constraint on the  $ggg$  correlation function.



**Figure 22:** Transverse single-spin asymmetry of isolated direct photons at  $\sqrt{s} = 200 \text{ GeV}$  compared with calculations from  $qqq$  and  $ggg$  correlation functions. Source: [69].

672 Similarly, the production of open heavy flavor at RHIC energies is dominated by gluon-  
673 gluon hard interactions. As such, also in single-spin asymmetries of heavy flavor decay  
674 leptons no final-state effect contributions are expected, and one is almost entirely sensitive  
675 to the initial state effects of the gluon correlators. The recent heavy flavor decay electron  
676 single-spin asymmetries at central rapidities obtained at PHENIX [72] are the first that  
677 quantify the gluon correlator contributions in two theoretical models [73,74], as can be seen  
678 in Fig. 23. While each decay lepton asymmetry is only sensitive to a linear combination of  
679 the two model parameters, the combination of both charges enables the determination of  
680 both.

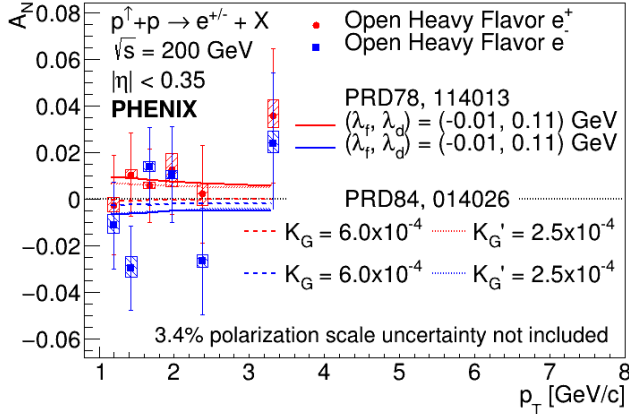


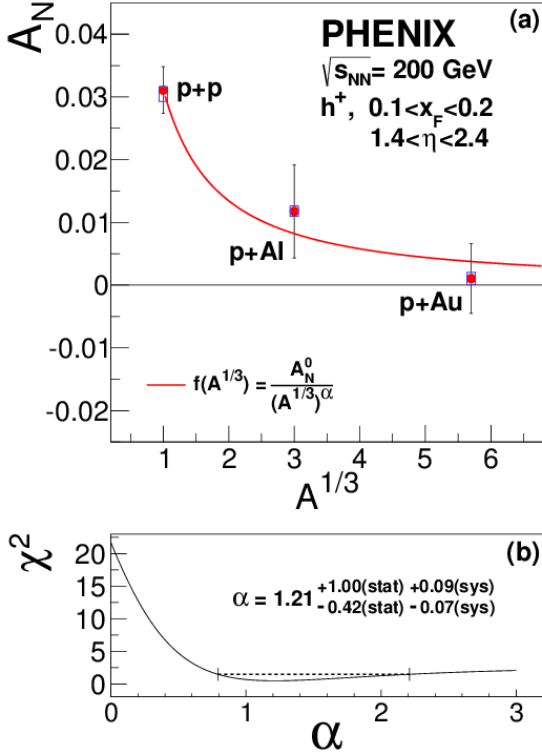
Figure 23: Transverse single-spin asymmetries of heavy flavor decay electrons at  $\sqrt{s} = 200$  GeV [72] including parameterizations of the tri-gluon correlator in two theoretical models and the best values fitting the data [73,74].

## 681 4.6 Nuclear dependence of single spin asymmetries

682 In 2015, RHIC also investigated polarized proton-nucleus collisions with either Al or Au  
683 beams. These have been utilized to study the  $A$  dependence of the nonzero single-spin asym-  
684 metries that were observed for hadrons in the forward region. In PHENIX the asymmetries  
685 for charged hadrons at rapidities of 1.2 to 2.4 were studied. A strong nuclear dependence was  
686 observed that was consistent with an  $A^{-1/3}$  suppression for positive hadrons [75], as shown  
687 in Fig. 24. A similar suppression is also seen as a function of the centrality of the collisions.

688 STAR has also published the  $A$  dependence for neutral pions at forward rapidities of  
689 2.7 to 3.8 and higher  $x_F$  that also show a suppression of the asymmetries [53]. However, in  
690 that rapidity region the suppression appears much smaller than seen by PHENIX, as seen  
691 in Fig. 25. The initial motivation for studying the nuclear dependence of the single-spin  
692 asymmetries originates from possible saturation effects on these asymmetries, but it has  
693 since been realized that the presented measurements neither reach  $x$  nor scales that are low  
694 enough for such effects to be relevant [76]. As such, there is at present no clear understanding  
695 of the mechanism that produces the suppression of these asymmetries.

696 In the far forward region also the nuclear dependence of neutron asymmetries was ex-  
697 tracted as a function of transverse momentum and the longitudinal momentum fraction  
698 [77,78]. Neutron asymmetries in proton-proton collisions can be described by the inter-  
699 ference of pion and other meson interactions between the two colliding nucleons [79] and

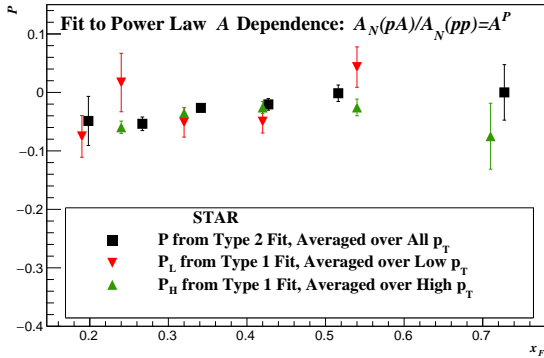


**Figure 24:**  $A$  dependence of transverse single-spin asymmetries of positively charged hadrons at  $\sqrt{s} = 200$  GeV at rapidities of 1.2 to 2.4 measured at PHENIX [75].

700 are found to be negative. In contrast, the pAl asymmetries are on average close to zero,  
 701 while the pAu asymmetries change sign and have a significantly larger magnitude. It was  
 702 found that the origin of this nuclear dependence originates from the additional contribution  
 703 of ultra-peripheral collisions that increase quadratically with the charge of the nucleus [80].  
 704 When correlating the asymmetries with event activity related to hadronic activity, one in-  
 705 deed sees that the asymmetries remain negative while the events more likely to originate from  
 706 ultra-peripheral collisions show even larger, positive asymmetries already for pAl collisions.

## 707 4.7 Ultra-peripheral collisions

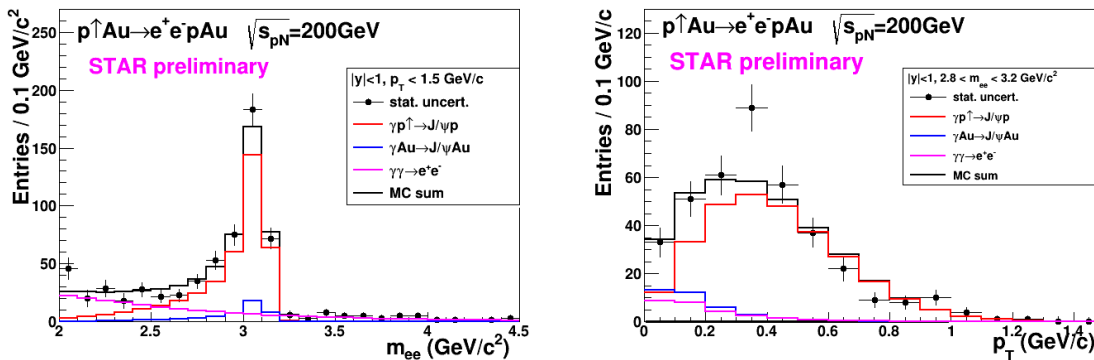
708 Constraints on GPDs have mainly been provided by exclusive reactions in DIS, e.g. deeply  
 709 virtual Compton scattering. RHIC, with its unique capability to collide transversely polar-  
 710 ized protons at high energies, has the opportunity to measure  $A_N$  for exclusive  $J/\psi$  produc-  
 711 tion in ultra-peripheral collisions (UPCs) [81]. In such a UPC process, a photon emitted by  
 712 the opposing beam particle ( $p$  or  $A$ ) collides with the polarized proton. The measurement is  
 713 at a fixed  $Q^2 \sim M_{J/\psi}^2 \approx 10$  GeV<sup>2</sup> and  $10^{-4} < x < 10^{-1}$ . A nonzero asymmetry would be the  
 714 first signature of a nonzero GPD  $E_g$  for gluons, which is sensitive to spin-orbit correlations  
 715 and is intimately connected with the orbital angular momentum carried by partons in the



**Figure 25:** The exponent,  $P$ , for nuclear  $A$  dependence of the  $\pi^0$  transverse single-spin asymmetry ratio of  $pA$  to  $pp$  as a function of  $x_F$  at  $\sqrt{s} = 200$  GeV at  $2.7 < \eta < 3.8$  at STAR [53]. The main difference of two types of fits is with and without correlated uncertainties.

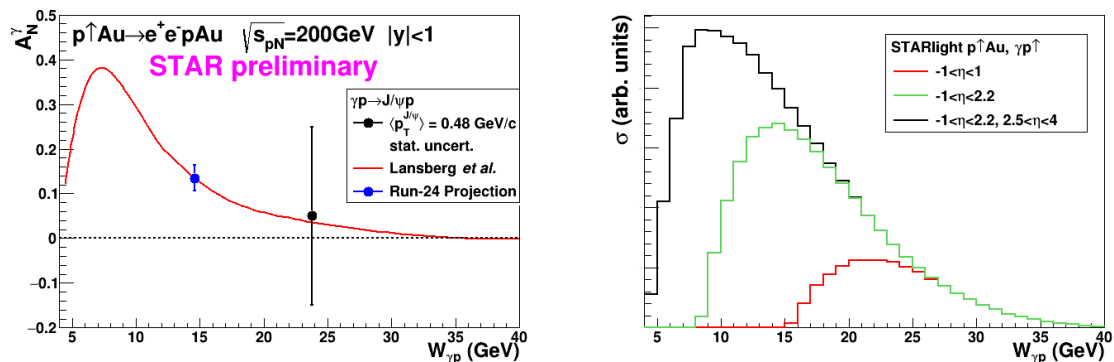
716 nucleon and thus with the proton spin puzzle.

717 The Run-15  $p^\uparrow Au$  data allowed a proof-of-principle of such a measurement. A trigger  
 718 requiring back-to-back energy deposits in the Barrel Electromagnetic Calorimeter selected  
 719  $J/\psi$  candidates. The  $e^+e^-$  mass distribution after selection cuts is shown in the left of  
 720 Fig. 26, and the pair  $p_T$  distribution of the  $J/\psi$  mass peak is shown on the right of that  
 721 figure. The data are well described by the STARlight model [82] (colored histograms in  
 722 the figure), including the dominant  $\gamma+p^\uparrow \rightarrow J/\psi$  signal process and the  $\gamma+Au \rightarrow J/\psi$  and  
 723  $\gamma+\gamma \rightarrow e^+e^-$  background processes. The left of Fig. 27 shows the STAR preliminary mea-  
 724 surement (solid circle marker) of the transverse asymmetry  $A_N^\gamma$  for the  $J/\psi$  signal, which has  
 725 a mean photon-proton center-of-mass energy  $W_{\gamma p} \approx 24$  GeV. The result is consistent with  
 726 zero. Also shown is a prediction based on a parameterization of  $E_g$  [83]; the present data  
 727 provide no discrimination of this prediction.



**Figure 26:** Mass distribution of selected  $e^+e^-$  pairs (left), and  $p_T$  distribution of the  $J/\psi$  mass peak (right). The colored histograms are the indicated processes modelled by STARlight and the sum fit to the data.

728 This measurement can be greatly improved with a high statistics transversely polarized  
 729  $p^\uparrow Au$  Run-24. The integrated luminosity for the Run-15 measurement was  $140 \text{ nb}^{-1}$ ; Run-24  
 730 will provide about  $1.2 \text{ pb}^{-1}$ , allowing a sizeable reduction of statistical uncertainty in the



**Figure 27:** Left: The measured  $J/\psi$  transverse asymmetry  $A_N^\gamma$  and a prediction based on a parameterization of  $E_g$ . Right: The accepted cross section for  $\gamma + p^\uparrow \rightarrow J/\psi$  for various detector pseudorapidity  $\eta$  ranges; the black curve shows the result for the full STAR detector with the Forward Upgrade and the iTPC.

731 same  $W_{\gamma p}$  range. In addition, the Forward Upgrade and iTPC will provide a significant  
 732 extension of the  $W_{\gamma p}$  range of the measurement. The right panel of Fig. 27 shows the  
 733 accepted cross section for  $\gamma + p^\uparrow \rightarrow J/\psi$  for various detector pseudorapidity ranges. With  
 734 the full detector, the sensitive cross section is a factor of five times the central barrel alone  
 735 and the expected asymmetry is substantially larger. The projected statistical uncertainty  
 736 on  $A_N^\gamma$  is shown in the left of Fig. 27 (blue square marker), offering a powerful test of a  
 737 non-vanishing  $E_g$ . Also, the accepted region has a lower mean  $W_{\gamma p} \approx 14$  GeV. Predictions  
 738 based on  $E_g$  parameterizations such as shown in the figure have a larger asymmetry at lower  
 739  $W_{\gamma p}$ , with increased possibility of a nonzero result. Alternatively, the increased statistics  
 740 will allow a measurement of  $A_N^\gamma$  in bins of  $W_{\gamma p}$ .

741 The UPC cross section scales with  $\sim Z^2$  of the the nucleus emitting the photon; for pro-  
 742 tons this is  $1/79^2$  relative to Au nuclei, which makes analogous measurements in  $pp$  collisions  
 743 extremely luminosity-hungry. Therefore, the  $pAu$  run is important for this measurement.



## 744 5 Appendix

### 745 5.1 STAR Forward Upgrade

746 The forward upgrade consists of four major new subsystems, an electromagnetic calorimeter,  
747 a hadronic calorimeter and a tracking system, formed from a silicon detector and a small-  
748 strip Thin Gap Chambers tracking detector. It has superior detection capabilities for neutral  
749 pions, photons, electrons, jets, and leading hadrons within the pseudorapidity range  $2.5 <$   
750  $\eta < 4$ , see Fig. 28. The construction of the electromagnetic and hadronic calorimeters had  
751 been successfully completed by the end of 2020. They were fully installed, instrumented, and  
752 commissioned during the 2021 RHIC running period. The tracking detectors were installed  
753 in summer and fall 2021, on schedule and ready for the start of Run-22. Note that the entire  
754 construction, installation, and commissioning of the four systems were completed in the  
755 pandemic period. Enormous efforts were made to keep the forward upgrades on schedule.  
756 During Run-22, despite all the difficulties from the machine side, the forward upgrades  
757 performed exceptionally well and took data smoothly throughout the run. The forward  
758 upgrades will continue taking data in parallel with sPHENIX through Run-25.

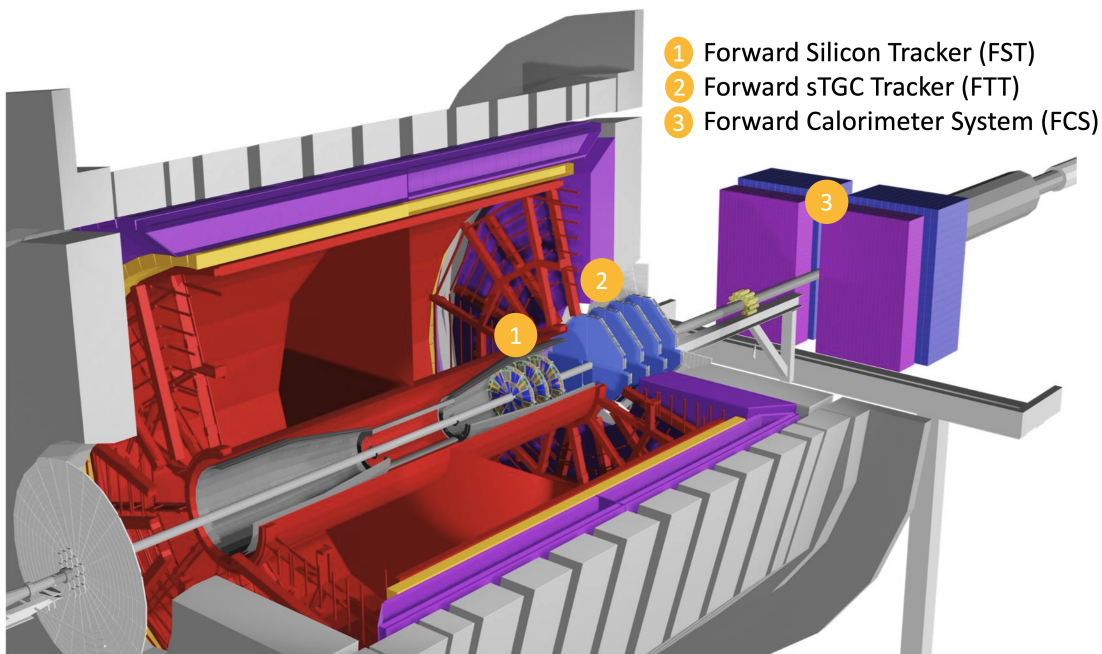


Figure 28: STAR detector with Forward Upgrades

#### 759 • Forward Calorimeter System:

760 The Forward Calorimeter System (FCS) consists of an Electro-Magnetic Calorimeter  
761 (Ecal) with 1486 towers, and a Hadronic Calorimeter (Hcal) with 520 towers. All SiPM  
762 sensors, front-end electronics boards and readout & triggering boards called DEP were

763 installed, commissioned and calibrated during Run-21. Signal splitter boards for the  
764 west EPD detector were installed before Run-22, and the west EPD was used as pre-  
765 shower detector in the electron triggers. FPGA code for FCS triggers was developed  
766 in fall 2021, and total of 29 triggers, including triggers for di-electrons, jets, di-jets,  
767 hadrons, and photons were commissioned and verified within a few days of RHIC start-  
768 ing to deliver stable  $pp$  collisions, and then used for data taking throughout Run-22  
769 successfully. FCS operations during Run-22 were successful and smooth. The only  
770 minor exceptions were 3 low-voltage power supply modules needing to be replaced,  
771 and occasional power cycling of electronics being needed due to beam related radiation  
772 upsets in the electronics. All 1486 channels of Ecal worked with no bad channels, and  
773 the Hcal had only a couple of dead channels. Radiation damage to the SiPM sensors  
774 due to beam was within expectations. There was an unexpected loss of signal ampli-  
775 tudes of  $\sim 20\%$  per week in the Ecal near the beam, which turned out to be radiation  
776 damage in the front-end electronics boards. The loss of signal was compensated during  
777 Run-22 by changing the gain factors on the DEP boards, attenuator settings in the  
778 front-end electronics, and raising the voltage settings tower by tower based on LED  
779 signals. Details of the radiation damage on the front-end electronics are currently  
780 under investigation.

781 • Small-strip Thin Gap Chambers:

782 The sTGC has four identical planes, each plane has four identical pentagonal shaped  
783 gas chambers. These gas chambers are made of double-sided and diagonal strips that  
784 give  $x, y, u$  in each plane. Sixteen chambers and about 5 spare chambers were built  
785 at Shandong University in China. A custom designed and fabricated aluminum frame  
786 allowed to fit the detector inside the pole-tip of the STAR magnet and around the beam-  
787 pipe on the west side of STAR. The sTGC chambers are operated with a quenching  
788 gas mixture of  $n$ -Pentane and  $\text{CO}_2$  at a ratio of 45%:55% by volume at a typical  
789 high voltage of 2900 V. This gas mixture allowed the chambers to operate in a high  
790 amplification mode. The sTGC was fully installed prior to the start of Run-22, and the  
791 detector was fully commissioned during the first few weeks of the run. The operating  
792 point of the high voltage was scanned for optimum efficiency. The gas chambers were  
793 stable at the desired operational high voltage and at the high luminosity, also the  
794 leakage current was well within the operational limits. In-house, a newly designed and  
795 built gas system for mixing, and supplying the gas along a long-heated path to deliver  
796 to the chambers, met the above requirements, and performed exceptionally well during  
797 Run-22.

798 • Forward Silicon Tracker:

799 The Forward Silicon Tracker (FST) consists of three identical disks, and each disk  
800 contains 12 modules. Each module has 3 single-sided double-metal Silicon mini-strips  
801 sensors which are readout by 8 APV chips. The module production was done by NCKU,  
802 UIC, and SDU. The readout was done by BNL and IU. The cooling was provided by  
803 NCKU and BNL. The installation of the FST was completed on August 13th, 2021,

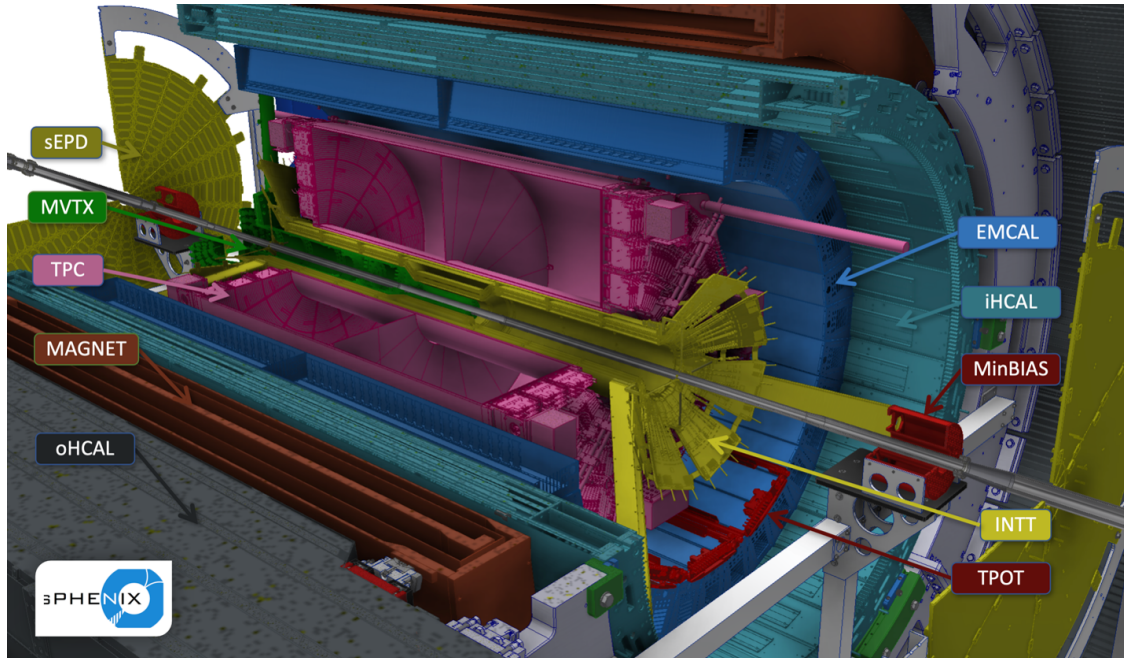


Figure 29: sPHENIX detector layout.

804 and the first  $pp$  510 GeV collision data were recorded on December 15, 2021. The FST  
 805 ran smoothly through the whole Run-22 and the detector operation via slow control  
 806 software was minimal to the shift crew.

## 807 5.2 sPHENIX Detector

808 sPHENIX is a major upgrade to the PHENIX experiment at RHIC capable of measuring  
 809 jets, photons, charged hadrons, and heavy flavor probes. sPHENIX will play a critical role in  
 810 the completion of the RHIC science mission, focused on the studies of the microscopic nature  
 811 of Quark-Gluon Plasma. Polarized proton collisions as well as proton-nucleus collisions will  
 812 also provide key opportunities for cold QCD measurements.

813 sPHENIX is a central rapidity detector ( $|\eta| < 1.1$ ) built around the Babar solenoid with  
 814 magnetic field up to 1.5T. The major systems are a high precision tracking system, and  
 815 electromagnetic and hadronic calorimeters, see Fig. 29.

816 The electromagnetic calorimeter is a compact tungsten-scintillating fiber design located  
 817 inside the solenoid. The outer hadronic calorimeter consists of steel in which scintillator tiles  
 818 with light collected by wavelength shifting fibers are sandwiched between tapered absorber  
 819 plates that project nearly radially from the interaction point. It also serves as a flux return  
 820 of the 1.5 T superconducting solenoid. The inner HCal is instrumented with scintillating  
 821 tiles similar to the tiles used in the Outer HCal, and serves as a support structure of the  
 822 electromagnetic calorimeter. The calorimeters use a common set of silicon photomultiplier  
 823 photodetectors and amplifier and digitizer electronics. Based on test beam data, such a

824 calorimeter system is expected to provide the energy resolution of  $\sigma_E/E = 13\%/\sqrt{E[\text{GeV}]} \oplus$   
825  $3\%$  for electromagnetic showers, and  $\sigma_E/E = 65\%/\sqrt{E[\text{GeV}]} \oplus 14\%$  for hadrons.

826 The central tracking system consists of a small Time Projection Chamber (TPC), micro  
827 vertex detector (MVTX) with three layers of Monolithic Active Pixel Sensors (MAPS), and  
828 two layers of the intermediate silicon strip tracker within the inner radius (INTT). Such a  
829 system provides momentum resolution  $\sigma_{p_T}/p_T < 0.2\% \cdot p_T \oplus 1\%$  for  $p_T = 0.2\text{--}40$  GeV/c, and  
830 Distance of Closest Approach (DCA) resolved at  $10 \mu\text{m}$  for  $p_T > 2$  GeV/c. The INTT with  
831 its fast integration time resolves beam crossings and provides pileup suppression.

832 The other sPHENIX subsystems are the Minimum Bias Detector (MBD) consisting of  
833 the refurbished PHENIX Beam-Beam Counter, Event Plane Detector (sEPD) consisting  
834 of two wheels of scintillator tiles positioned at  $2 < |\eta| < 4.9$  and serving for event plane  
835 measurements, and Micromegas-based TPC Outer Tracker (TPOT), offering calibration of  
836 beam-induced space charge distortions in TPC.

837 High speed data acquisition system is designed to be capable of taking minimum bias  
838  $AuAu$  collisions at 15 kHz with greater than 90% live time, and jet and photon triggers for  $pp$   
839 and  $pA$  operation. The DAQ system is design to be capable to work in hybrid mode: along  
840 with triggered data it will collect a significant fraction ( $\sim 10\%$ ) of all collision data from  
841 tracking detectors in streaming readout regime, which will greatly extend physics program  
842 in  $pp$  and  $pAu$  running.

## References

- [1] D. de Florian, R. Sassot, M. Stratmann, and W. Vogelsang, Phys. Rev. Lett. **113**, 012001 (2014).
- [2] DSSV preliminary.
- [3] NNPDF, E. R. Nocera, R. D. Ball, S. Forte, G. Ridolfi, and J. Rojo, Nucl. Phys. B **887**, 276 (2014).
- [4] STAR Collaboration, M. S. Abdallah *et al.*, Phys. Rev. Lett. **129**, 092501 (2022), 2111.10396.
- [5] STAR Collaboration, M. M. Aggarwal *et al.*, Phys. Rev. Lett. **106**, 062002 (2011), 1009.0326.
- [6] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. Lett. **106**, 062001 (2011), 1009.0505.
- [7] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. Lett. **113**, 072301 (2014), 1404.6880.
- [8] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. D **93**, 051103 (2016), 1504.07451.
- [9] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. D **98**, 032007 (2018), 1804.04181.
- [10] STAR Collaboration, J. Adam *et al.*, Phys. Rev. D **99**, 051102 (2019), 1812.04817.
- [11] D. de Florian, R. Sassot, M. Stratmann, and W. Vogelsang, Phys. Rev. Lett. **113**, 012001 (2014).
- [12] STAR Collaboration, J. Adam *et al.*, Phys. Rev. D **98**, 032011 (2018).
- [13] STAR Collaboration, J. Adam *et al.*, Phys. Rev. D **100**, 052005 (2019).
- [14] STAR Collaboration, M. S. Abdallah *et al.*, Phys. Rev. D **103**, L091103 (2021), 2103.05571.
- [15] STAR Collaboration, M. S. Abdallah *et al.*, Phys. Rev. D **105**, 092011 (2022), 2110.11020.
- [16] STAR Collaboration, N. Lukow, Constraining the Polarized Gluon Distribution Function of the Proton with Recent STAR Measurements, in *24th International Spin Symposium*, 2021.
- [17] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. D **93**, 011501 (2016).
- [18] PHENIX Collaboration, U. A. Acharya *et al.*, Phys. Rev. D **102**, 032001 (2020), 2004.02681.

- 874 [19] NNPDF, E. R. Nocera, R. D. Ball, S. Forte, G. Ridolfi, and J. Rojo, Nucl. Phys. B  
875 **887**, 276 (2014), 1406.5539.
- 876 [20] STAR Collaboration, B. I. Abelev *et al.*, Phys. Rev. Lett. **97**, 252001 (2006).
- 877 [21] STAR Collaboration, B. I. Abelev *et al.*, Phys. Rev. Lett. **100**, 232003 (2008).
- 878 [22] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. D **86**, 032006 (2012).
- 879 [23] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. Lett. **115**, 092002 (2015).
- 880 [24] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. D **95**, 071103 (2017).
- 881 [25] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. D **76**, 051106 (2007).
- 882 [26] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. Lett. **103**, 012003 (2009).
- 883 [27] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. D **79**, 012003 (2009).
- 884 [28] STAR Collaboration, B. Abelev *et al.*, Phys. Rev. D **80**, 111108 (2009).
- 885 [29] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. D **89**, 012001 (2014).
- 886 [30] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. D **90**, 012007 (2014).
- 887 [31] STAR Collaboration, J. Adam *et al.*, Phys. Rev. D **98**, 032013 (2018).
- 888 [32] PHENIX Collaboration, U. A. Acharya *et al.*, (2022), 2202.08158.
- 889 [33] D. de Florian, G. A. Lucero, R. Sassot, M. Stratmann, and W. Vogelsang, Phys. Rev.  
890 D **100**, 114027 (2019).
- 891 [34] STAR, J. Kwasizur, Longitudinal Double-Spin Asymmetries for Dijet Production at  
892 Intermediate Pseudorapidity in Polarized Proton-Proton Collisions at  $\sqrt{s} = 510$  GeV,  
893 in *APS Division of Nuclear Physics Meeting*, 2020.
- 894 [35] Jefferson Lab Angular Momentum (JAM), Y. Zhou, N. Sato, and W. Melnitchouk,  
895 Phys. Rev. D **105**, 074022 (2022), 2201.02075.
- 896 [36] BRAHMS, I. Arsene *et al.*, Phys. Rev. Lett. **91**, 072305 (2003), nucl-ex/0307003.
- 897 [37] BRAHMS, I. Arsene *et al.*, Phys. Rev. Lett. **93**, 242303 (2004), nucl-ex/0403005.
- 898 [38] PHENIX Collaboration, S. S. Adler *et al.*, Phys. Rev. Lett. **94**, 082302 (2005), nucl-  
899 ex/0411054.
- 900 [39] STAR Collaboration, J. Adams *et al.*, Phys. Rev. Lett. **97**, 152302 (2006), nucl-  
901 ex/0602011.

- 902 [40] PHENIX Collaboration, A. Adare *et al.*, Phys. Rev. Lett. **107**, 172301 (2011),  
903 1105.5112.
- 904 [41] M. Strikman and W. Vogelsang, Phys. Rev. D **83**, 034029 (2011), 1009.6123.
- 905 [42] L. D. McLerran and R. Venugopalan, Phys. Rev. D **49**, 2233 (1994), hep-ph/9309289.
- 906 [43] L. D. McLerran and R. Venugopalan, Phys. Rev. D **49**, 3352 (1994), hep-ph/9311205.
- 907 [44] E. Iancu, A. Leonidov, and L. McLerran, The colour glass condensate: An introduction,  
908 pp. 73–145, Springer, 2002.
- 909 [45] STAR Collaboration, M. Abdallah *et al.*, Phys. Rev. D **106**, 072010 (2022), 2205.11800.
- 910 [46] U. D’Alesio, F. Murgia, and C. Pisano, Phys. Lett. B **773**, 300 (2017), 1707.00914.
- 911 [47] Z.-B. Kang, X. Liu, F. Ringer, and H. Xing, JHEP **11**, 068 (2017), 1705.08443.
- 912 [48] Z.-B. Kang, A. Prokudin, F. Ringer, and F. Yuan, Phys. Lett. B **774**, 635 (2017),  
913 1707.00913.
- 914 [49] U. D’Alesio, F. Murgia, and C. Pisano, Phys. Rev. D **83**, 034021 (2011), 1011.2692.
- 915 [50] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. D **97**, 032004 (2018), 1708.07080.
- 916 [51] J. Collins and J.-W. Qiu, Phys. Rev. D **75**, 114014 (2007).
- 917 [52] T. C. Rogers and P. J. Mulders, Phys. Rev. D **81**, 094006 (2010).
- 918 [53] STAR Collaboration, J. Adam *et al.*, Phys. Rev. D **103**, 072005 (2021), 2012.07146.
- 919 [54] STAR, H. Liu, Measurement of transverse single-spin asymmetries for dijet production  
920 in polarized p+p collisions at  $\sqrt{s} = 200$  GeV at STAR, in *Nuclear Physics Seminar*,  
921 *Brookhaven National Laboratory*, 2020, <https://indico.bnl.gov/event/8633/>.
- 922 [55] X. Liu, F. Ringer, W. Vogelsang, and F. Yuan, Factorization and its breaking in dijet  
923 single transverse spin asymmetries in  $pp$  collisions, 2020, 2008.03666.
- 924 [56] Z.-B. Kang, K. Lee, D. Y. Shao, and J. Terry, The sivers asymmetry in hadronic dijet  
925 production, 2020, 2008.05470.
- 926 [57] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. Lett. **115**, 242501 (2015),  
927 1504.00415.
- 928 [58] STAR Collaboration, L. Adamczyk *et al.*, Phys. Lett. B **780**, 332 (2018), 1710.10215.
- 929 [59] STAR, B. Pokhrel, Transverse Spin Dependent Azimuthal Correlations of  $\pi^+\pi^-$  pair in  
930  $p^\uparrow p$  Collisions at  $\sqrt{s} = 200$  GeV at STAR, in *24th International Spin Symposium*, 2021.

- 931 [60] M. Radici and A. Bacchetta, Phys. Rev. Lett. **120**, 192001 (2018), 1802.05212.
- 932 [61] STAR Collaboration, J. Adam *et al.*, Phys. Rev. D **103**, 092009 (2021), 2012.11428.
- 933 [62] *Transverse Single Spin Asymmetry for Inclusive and Diffractive Electromagnetic Jets at*  
934 *Forward Rapidities in  $p\uparrow+p$  Collisions at  $s = 200$  GeV and  $510$  GeV at STAR*, Zenodo,  
935 2022.
- 936 [63] M. Boggione *et al.*, Phys. Lett. B **815**, 136135 (2021), 2101.03955.
- 937 [64] RHIC-f, M. H. Kim *et al.*, Phys. Rev. Lett. **124**, 252501 (2020), 2003.04283.
- 938 [65] H.-J. Kim, S. Clymton, and H.-C. Kim, Phys. Rev. D **106**, 054001 (2022), 2206.02184.
- 939 [66] M. Bury, A. Prokudin, and A. Vladimirov, JHEP **05**, 151 (2021), 2103.03270.
- 940 [67] M. Bury, A. Prokudin, and A. Vladimirov, Phys. Rev. Lett. **126**, 112002 (2021),  
941 2012.05135.
- 942 [68] A. Bacchetta, F. Delcarro, C. Pisano, and M. Radici, Phys. Lett. B **827**, 136961 (2022),  
943 2004.14278.
- 944 [69] PHENIX Collaboration, U. A. Acharya *et al.*, Phys. Rev. Lett. **127**, 162001 (2021).
- 945 [70] K. Kanazawa, Y. Koike, A. Metz, and D. Pitonyak, Phys. Rev. D **91**, 014013 (2015).
- 946 [71] Y. Koike and S. Yoshida, Phys. Rev. D **85**, 034030 (2012).
- 947 [72] PHENIX Collaboration, U. A. Acharya *et al.*, (2022), 2204.12899.
- 948 [73] Z.-B. Kang, J.-W. Qiu, W. Vogelsang, and F. Yuan, Phys. Rev. D **78**, 114013 (2008),  
949 0810.3333.
- 950 [74] Y. Koike and S. Yoshida, Phys. Rev. D **84**, 014026 (2011), 1104.3943.
- 951 [75] PHENIX Collaboration, C. Aidala *et al.*, Phys. Rev. Lett. **123**, 122001 (2019),  
952 1903.07422.
- 953 [76] S. Benić and Y. Hatta, Phys. Rev. D **99**, 094012 (2019), 1811.10589.
- 954 [77] PHENIX Collaboration, C. Aidala *et al.*, Phys. Rev. Lett. **120**, 022001 (2018),  
955 1703.10941.
- 956 [78] PHENIX Collaboration, U. A. Acharya *et al.*, Phys. Rev. D **105**, 032004 (2022),  
957 2110.07504.
- 958 [79] B. Z. Kopeliovich, I. K. Potashnikova, I. Schmidt, and J. Soffer, Phys. Rev. D **84**,  
959 114012 (2011), 1109.2500.



- 960 [80] G. Mitsuka, Phys. Rev. C **95**, 044908 (2017), 1702.03834.
- 961 [81] S. Klein and J. Nystrand, Photoproduction of J / psi and Upsilon in pp and anti-p p  
962 collisions, in *5th Workshop on Small x and Diffractive Physics*, 2003, hep-ph/0310223.
- 963 [82] S. R. Klein, J. Nystrand, J. Seger, Y. Gorbunov, and J. Butterworth, Comput. Phys.  
964 Commun. **212**, 258 (2017), 1607.03838.
- 965 [83] J. Lansberg, L. Massacrier, L. Szymanowski, and J. Wagner, Phys. Lett. B **793**, 33  
966 (2019), 1812.04553.