



Calculation of Specific Heat in 3 GeV FXT Au+Au

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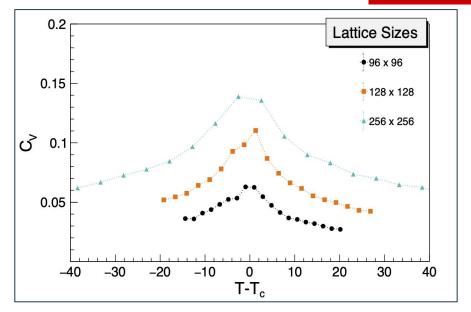


Motivation

 For a system undergoing phase transition, C_v is expected to diverge at the critical point.

 Thus the variation of thermal fluctuations with temperature can be effectively used to probe the critical point.

In the 2D Ising Model we can see the divergence of C_v as we change the temperature of the system.



Simulations from 2D Ising Model for C_v.





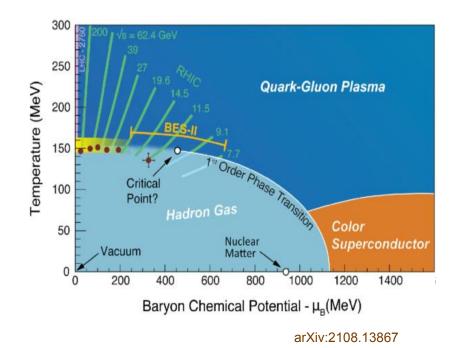
Objective

- ◆ Calculate C_v as a function of $\sqrt{s_{NN}}$, locate the critical point, where the analytical crossover becomes a 1st Order Phase Transition.
- $C_v = \left(\frac{\partial E}{\partial T}\right)_V$ Definition from Thermodynamics
- Heat Capacity, C of a system in thermal equilibrium connected to a bath at T can be computed from the event- by-event fluctuations of Energy:

$$C = rac{(<\!E^2>\!-\!<\!E>^2)}{<\!T>^2}$$

arXiv:1601.05631







STAR Objective

 For a system in equilibrium, the event-by-event temperature fluctuation is controlled by the heat capacity:

$$P(T) \propto exp[-rac{C}{2}rac{\left(\Delta T
ight)^2}{< T >^2}]$$

arXiv:1601.05631

 Specific Heat calculated from fluctuations in temperature:

$$\left(\frac{1}{C} = \frac{< T^2 > - < T >^2}{< T >^2} \right)$$



At kinetic freeze-out:

$$\overline{rac{1}{C} = rac{<\!T_{kin}^2 \!>\! -\! <\! T_{kin} \!>^2}{<\!T_{kin} \!>^2}}$$

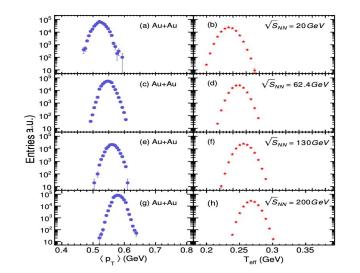
arXiv:1601.05631

<u>Caveat:</u> Measuring the temperature of a single event based on the p_{T} distribution does not yield a very good precision.





$$\left(rac{1}{C} \, = rac{(<\! T_{kin}^2 \! >\! - <\! T_{kin} \! >^2)}{<\! T_{kin} \! >^2} \, pprox rac{(<\! T_{eff}^2 \! >\! - <\! T_{eff} \! >^2)}{<\! T_{kin} \! >^2}
ight)$$



- Effective temperature (T_{eff}) . *
- Calculated from $< p_{T} >$ distributions. *
- $T_{eff} = T_{kin} + f()$ *

Fig. from Sumit Basu et al 2016 J. Phys.: Conf. Ser. 668 012043





- The Kinetic freeze-out parameters are extracted using the SSH BW.
- Using the blast wave we can find:
 - Kinetic freeze-out temperature
 T_{kin}.
 - Transverse radial flow
 - ➢ Flow profile n.
- 3 Fit parameters.

$$rac{1}{2\pi p_T}rac{d^2N}{dp_Tdy} \propto \int_0^R r dr m_T I_0(rac{p_T sinh
ho}{T_{kin}}) K_1(rac{m_T cosh
ho}{T_{kin}})$$

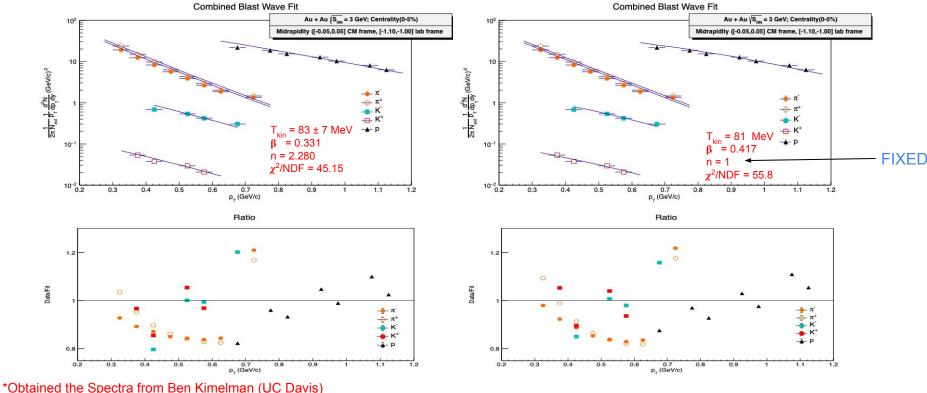
PRC48 (1993) 2462

$$\left[
ho = tanh^{-1}eta_T
ight] < eta_T > = rac{2}{2+n}eta_s$$

- $_\star
 ho$: Is the boost angle.
- $\star \beta_s$: Is the surface velocity of the thermal source and n is the flow profile.







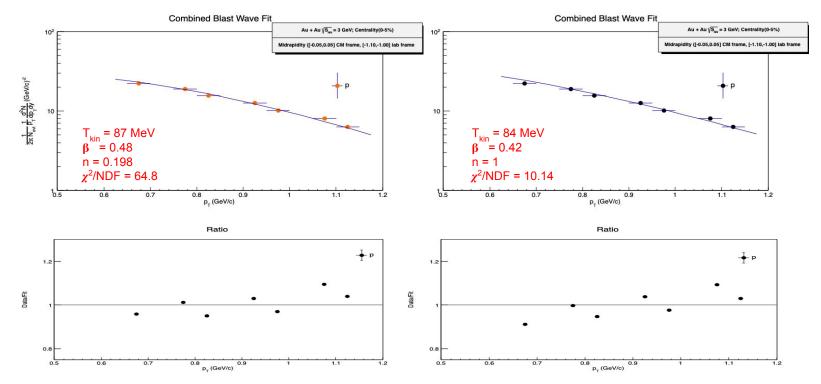




Combined Blast Wave Fit Combined Blast Wave Fit Au + Au VS_{NN} = 3 GeV; Centrality(0-5%) Au + Au VS_{NN} = 3 GeV; Centrality(0-5%) Midrapidity ([-0.05,0.05] CM frame, [-1.10,-1.00] lab frame Midrapidity ([-0.05.0.05] CM frame, [-1.10.-1.00] lab frame $\rightarrow \pi^+$ $rac{1}{2\pi}rac{1}{h_{evt}}rac{1}{p_T}rac{d^2N}{dp_Tdy}(\text{GeV/c})^2$ $\rightarrow \pi^+$ $\frac{1}{2\pi N_{evt}} \frac{1}{p_T} \frac{d^4N}{dp_T^2} (GeV/c)^2$ --------------------------------K ----- K - K+ - K+ $T_{kin} = 65 \pm 2 \text{ MeV}$ $T_{kin} = 52 \pm 2 \text{ MeV}$ $= 0.49 \pm 0.009$ 10 $\mathbf{B} = 0.414 \pm 0.04$ 10-**FIXED** n = 1 $n = 2.162 \pm 0.052$ γ^2 /NDF = 39.2 γ^2 /NDF = 16.3 10⁻² 10-2 0.5 0.6 0.7 0.8 0.9 0.3 0.4 0.3 0.4 0.5 0.6 0.7 0.8 0.9 p_T (GeV/c) p_ (GeV/c) Ratio Ratio - 🕂 π΄ 🕂 π΄ 1.2 1.2 $- \overline{\phi} \pi^+$ $-\frac{7}{2}\pi^{+}$ 🔶 К - K 📥 K+ Data/Fit Data/Fit 0 0.8 0.8 0.2 0.6 p_ (GeV/c) 0.7 0.8 0.9 0.2 0.4 0.6 p_T (GeV/c) 0.7 0.8 0.9











$\sqrt{s_{NN}}(GeV)$	Fit	n	$T_{kin}(MeV)$	$<\beta>$	$\frac{\chi^2}{NDF}$
200	$\pi^{\pm}, K^{\pm}, p \bar{p}$	0.817 ± 0.025	88 ± 5	0.594 ± 0.012	0.186
200	π^{\pm}, K^{\pm}	0.842 ± 0.167	84 ± 24	0.602 ± 0.078	0.06
62.4	$\pi^{\pm}, K^{\pm}, p\bar{p}$	0.61 ± 0.04	97 ± 1	0.554 ± 0.01	0.63
62.4	π^{\pm}, K^{\pm}	0.917 ± 0.094	88 ± 7	0.57 ± 0.031	0.16
14.5	$\pi^{\pm}, K^{\pm}, p\bar{p}$	0.991 ± 0.123	120 ± 2	0.431 ± 0.011	0.29
14.5	π^{\pm}, K^{\pm}	1.361 ± 0.191	125 ± 4	0.385 ± 0.013	0.25
3	π^{\pm}, K^{\pm}, p	1	81	0.415 ± 0.05	0.74
3	π^{\pm}, K^{\pm}	1	69 ± 5	0.465 ± 0.027	0.44
3	π^{\pm}, K^{\pm}, p	1.415 ± 0.447	81 ± 1	0.386 ± 0.022	0.71
3	π^{\pm}, K^{\pm}	1.9 ± 0.34	54 ± 12	0.431 ± 0.031	0.26

*Systematics of 5%

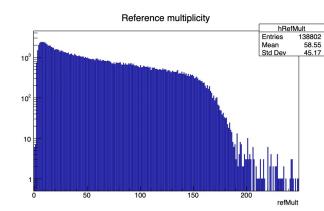


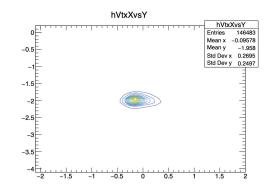
STAR $< p_{T} > Distributions$

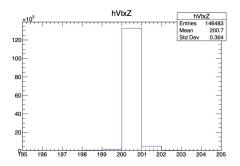


Acceptance Cuts:

- Fixed Target, run 2018 data production, 3 GeV Au+Au collision, y_c.m. ≈1.045
- ✤ 198 <V_z< 202 cm</p>
- V_r<1.5 cm about beam spot centered around [0,-2].
- Trigger ID 620052 and 620053 (Min.Bias)
- ✤ DCA < 3.0 cm</p>
- NhitsFit/NHitsMax > 0.51
- 0.5 M Events analysed (Rest ongoing)





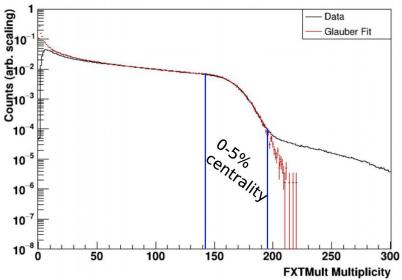




<p_> Distributions



Centrality Definition:



3 GeV FXTMult Distribution

Data Glauber Fit	Cen
	0
	5-

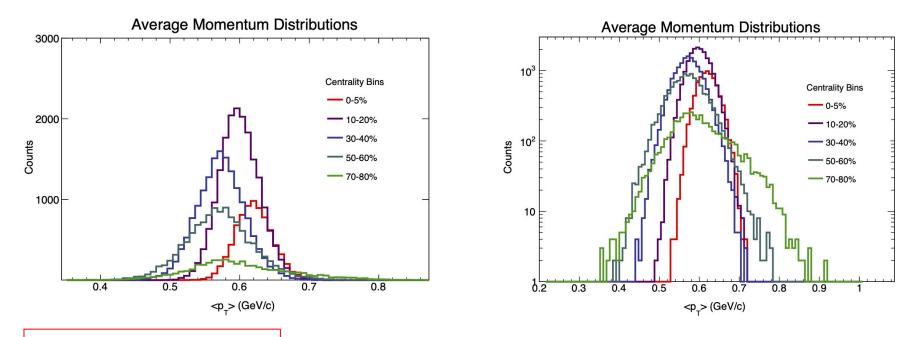
Centrality Bin	FXTMult Cuts (inclusive)		
0-5%	195-142		
5-10%	141-120		
10-20%	119-86		
20-30%	85-61		
30-40%	60-42		
40-50%	41-27		
50-60%	26-17		
60-70%	16-9		
70-80%	8-5		

*Pile-up cut at 195 *



STAR $< p_{T} > Distributions$



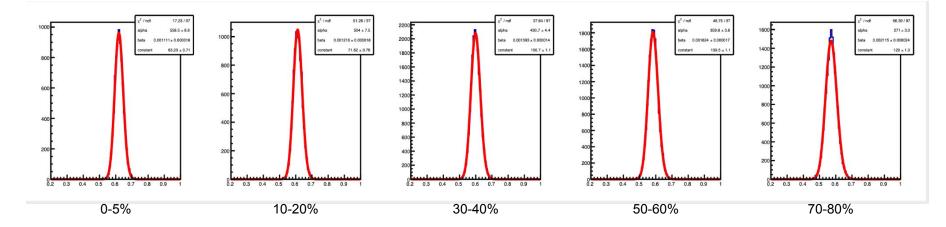


- All charged particles
- ✤ 0.15 < p_⊤ < 2.0 (GeV/c)</p>
- $\bullet \quad -2 < \eta < 0$



STAR $< p_{T} > Distributions$





It has been observed that the $\langle p_T \rangle$ distributions are nicely described by using the gamma (Γ) distribution.

```
f(x)=rac{x^{lpha-1}e^{rac{-x}{eta}}}{\Gamma(lpha)eta^{lpha}}
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Phys. Rev. C 94, 044901 (2016)

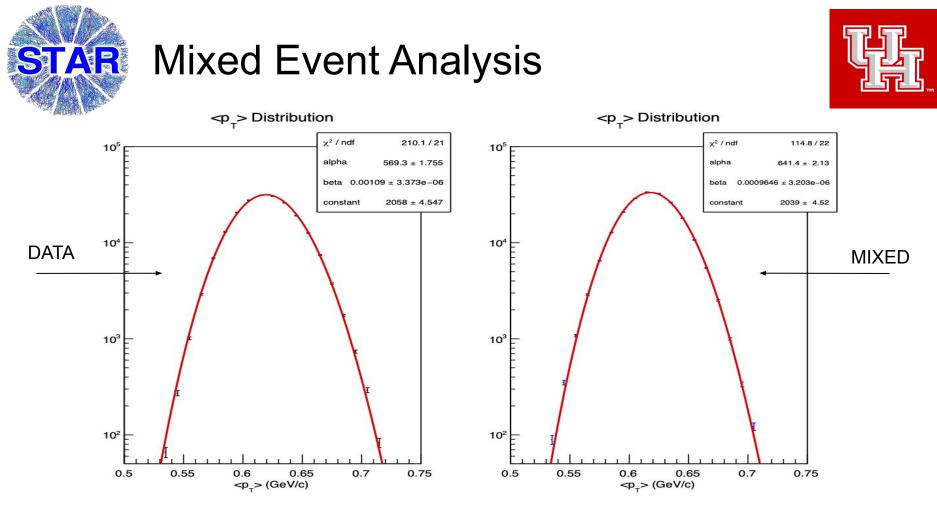


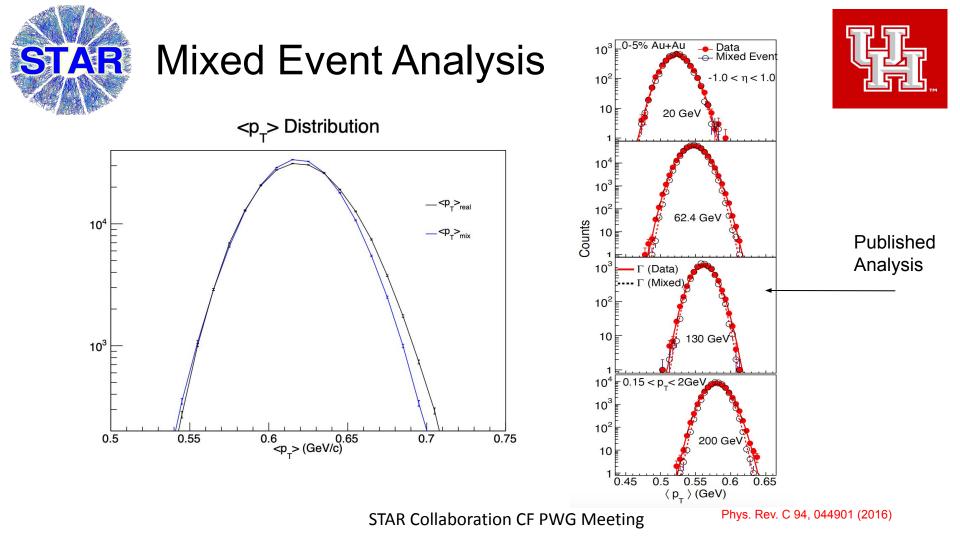
STAR Mixed Event Analysis



 In order to establish whether the observed fluctuations are partly dynamical in nature, we need to disentangle statistical effects i.e. effects due to the finite number of particles in the final state of the collision.

The Mixed event reconstruction makes synthetic events with tracks from different events to remove any kind of correlations.







STAR Mixed Event Analysis



$$\left(f(x)=rac{x^{lpha-1}e^{rac{-x}{eta}}}{\Gamma(lpha)eta^{lpha}}
ight)$$

From the fit parameters α , β We can calculate μ and σ :

Case	α	β (GeV)	μ (GeV)	σ (GeV)
Data	569.3	1.09 x 10 ⁻³	0.6205	0.0260
Mixed	641.4	9.64 x 10 ⁻⁴	0.6186	0.0244

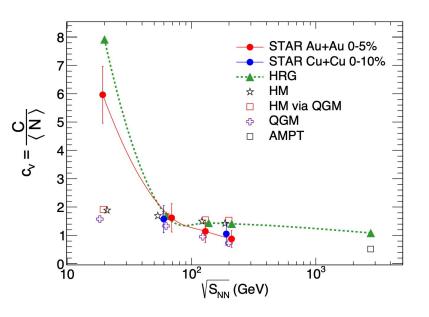


Next Steps



Working on:

- 1. Running on RCF (all possible events)
- 2. Working on i-HRG to obtain baseline
- 3. Systematics



Phys. Rev. C 94, 044901 (2016)





Thank you for your attention