

Calculation of Specific Heat in 3 GeV FXT Au+Au

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STAR Motivation

❖ For a system undergoing phase transition, C_{ν} is expected to diverge at the critical point.

❖ Thus the variation of thermal fluctuations with temperature can be effectively used to probe the critical point.

❖ In the 2D Ising Model we can see the divergence of C_v as we change the temperature of the system.

❖ Simulations from 2D Ising Model for C_{v} .

STAR Objective

- ❖ Calculate C_v as a function of $\sqrt{s_{NN}}$, locate the critical point, where the analytical crossover becomes a 1st Order Phase Transition.
- ❖ C_V = $\left(\frac{\partial E}{\partial T}\right)_V$ Definition from Thermodynamics
- ❖ Heat Capacity, C of a system in thermal equilibrium connected to a bath at T can be computed from the event- by-event fluctuations of Energy:

$$
C = \frac{(<\!E^2> - <\!E>^2)}{<\!T>^2}
$$

[arXiv:1601.05631](https://arxiv.org/abs/1601.05631)

 \triangleleft For a system in equilibrium, the event-by-event temperature fluctuation is controlled by the heat capacity:

$$
P(T) \propto exp[-\frac{C}{2} \frac{\left(\Delta T\right)^2}{<\!\! T\!\!>^2}]
$$

[arXiv:1601.05631](https://arxiv.org/abs/1601.05631)

❖ Specific Heat calculated from fluctuations in temperature:

$$
\frac{1}{C} = \frac{\langle T^2 \rangle - \langle T \rangle^2}{\langle T \rangle^2}
$$

❖ At kinetic freeze-out:

$$
\boxed{\frac{1}{C} = \frac{<\!T_{kin}^2> - <\!T_{kin}>^2}{<\!T_{kin}>^2}}
$$

[arXiv:1601.05631](https://arxiv.org/abs/1601.05631)

Caveat: Measuring the temperature of a single event based on the $\rho_{_{\mathsf{T}}}$ distribution does not yield a very good precision.

 $=\frac{(<\!\!T_{kin}^2\!\!>=<\!\!T_{kin}\!\!>^2)}{<\!\!T_{kin}\!\!>^2}\approx\frac{(<\!\!T_{eff}^2\!\!>=<\!\!T_{eff}\!\!>^2)}{<\!\!T_{kin}\!\!>^2}$ $\frac{1}{C}$ =

- ❖ Effective temperature (T_{eff}) .
- ❖ Calculated from $+p_T$ is a distribution of $−p_T$.
- $\begin{array}{cc} \text{\textbullet} & \text{\textbullet} \\ \text{\textbullet} & \text{\textbullet} \\ \text{\textcirc} & \text{\textcirc} \\ \beta_T & \end{array}$

Fig. from Sumit Basu *et al* 2016 *J. Phys.: Conf. Ser.* 668 012043

❖ The Kinetic freeze-out parameters are extracted using the SSH BW.

$$
\left(\tfrac{1}{2\pi p_T} \tfrac{d^2N}{dp_Tdy} \!\propto \int_0^R r dr m_T I_0(\tfrac{p_T sinh \rho}{T_{kin}}) K_1(\tfrac{m_Tcosh \rho}{T_{kin}})\right.
$$

- ❖ Using the blast wave we can find:
	- \triangleright Kinetic freeze-out temperature $\mathsf{T}_{\mathsf{kin}}$.
	- \triangleright Transverse radial flow
	- \triangleright Flow profile n.
- ❖ 3 Fit parameters.

$$
\left|\rho=tanh^{-1}\beta_T\right|\leq\beta_T>=\tfrac{2}{2+n}\beta_s
$$

PRC48 (1993) 2462

- $\star \rho$: Is the boost angle.
- $\alpha \beta_s$: Is the surface velocity of the thermal source and n is the flow profile.

Combined Blast Wave Fit Combined Blast Wave Fit Au + Au $\sqrt{S_{_{NN}}}$ = 3 GeV; Centrality (0-5%) Au + Au $\sqrt{S_{\text{max}}}$ = 3 GeV; Centrality (0-5%) 10 Midrapidity ([-0.05,0.05] CM frame, [-1.10,-1.00] lab frame Midrapidity ([-0.05,0.05] CM frame, [-1.10,-1.00] lab frame $\frac{1}{2}$ π 10 $\frac{1}{2}$ π 10 $\rightarrow \pi^+$ $\frac{1}{2\pi} \frac{1}{N_{\text{ext}}} \frac{1}{p_{\text{T}}} \frac{d^2 N}{dp_{\text{T}}} \frac{(\text{GeV/C})^2}{(S \text{eV/C})^2}$ $\rightarrow \pi^+$ $\frac{1}{2\pi N}\frac{1}{e_{\alpha\epsilon}}\frac{d^{\epsilon}N}{P_{T}}\frac{d^{\epsilon}N}{dp_{T}}$ (GeWc)² K K $\overline{+}$ K⁺ $+K$ ⁺ $= 65 \pm 2$ MeV $= 52 \pm 2$ MeV $\dddot{=}$ 0.49 ± 0.009 $10⁻$ 6^{10} = 0.414 ± 0.04 $10⁻$ FIXED $n = 1$ $n = 2.162 \pm 0.052$ χ^2 /NDF = 39.2 χ^2 /NDF = 16.3 10^{-2} _{0.2} 10^{-2} _{0.2} 0.3 0.4 0.5 0.6 0.7 0.8 0.9 0.3 0.5 0.6
 p_T (GeV/c) 0.7 0.8 0.9 0.4 p_r (GeV/c) Ratio Ratio \circ $-\frac{1}{2}$ π $\frac{1}{2}$ π 12 1.2 $\stackrel{1}{\rightarrow} \pi^*$ $\stackrel{<}{\leftrightarrow} \pi^*$ $+$ K K K^+ K^+ Data/Fit Data/Fit \circ \circ 0.8 0.8 $0.\overline{2}$ 0.6
 $p₋$ (GeV/c) 0.7 0.9 $0.\overline{2}$ 0.6
 $p₋$ (GeV/c) 0.7 0.9 0.8 0.3 0.4 0.8

*Systematics of 5%

Acceptance Cuts:

- ❖ Fixed Target, run 2018 data production, 3 GeV Au+Au collision, y_c.m. ≈1.045
- ❖ 198 < V_z < 202 cm
- ❖ V <1.5 cm about beam spot centered around [0,-2].
- ❖ Trigger ID 620052 and 620053 (Min.Bias)
- \bullet DCA < 3.0 cm
- ❖ NhitsFit/NHitsMax > 0.51
- ❖ 0.5 M Events analysed (Rest ongoing)

Centrality Definition:

❖ *Pile-up cut at 195

- ❖ All charged particles
- ❖ 0.15 < p_{T} < 2.0 (GeV/c)
- \div -2 < n < 0

It has been observed that the $\langle \mathsf{p}_{\mathsf{T}} \rangle$ distributions are nicely described by using the gamma (Γ) distribution.

$$
f(x)=\tfrac{x^{\alpha-1}e^{\tfrac{-x}{\beta}}}{\Gamma(\alpha)\beta^{\alpha}}
$$

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STAR Mixed Event Analysis

❖ In order to establish whether the observed fluctuations are partly dynamical in nature, we need to disentangle statistical effects i.e. effects due to the finite number of particles in the final state of the collision.

❖ The Mixed event reconstruction makes synthetic events with tracks from different events to remove any kind of correlations.

STAR Mixed Event Analysis

$$
f(x)=\tfrac{x^{\alpha-1}e^{\tfrac{-x}{\beta}}}{\Gamma(\alpha)\beta^{\alpha}}
$$

From the fit parameters α , β We can calculate μ and σ :

$$
\begin{array}{ll}\n\textbf{A} & \mu = \alpha \textbf{B} \\
\textbf{B} & \sigma^2 = \alpha \textbf{B}^2\n\end{array}
$$

Next Steps

Working on:

- 1. Running on RCF (all possible events)
- 2. Working on i-HRG to obtain baseline
- 3. Systematics

Phys. Rev. C 94, 044901 (2016)

Thank you for your attention