

# Updates on $\langle p_T \rangle$ fluctuations in $\sqrt{s_{NN}} = 3.0$ GeV FXT Au+Au

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# Brief Overview

- ❖  $C_v$  used to quantify dynamical fluctuations in temperature of the system.
- ❖ Effective temperature ( $T_{\text{eff}}$ ).
- ❖ Calculated from  $\langle p_T \rangle$  distributions.
- ❖  $T_{\text{eff}} = T_{\text{kin}} + f(\beta_T)$
- ❖  $T_{\text{kin}}$  obtained from Spectra

$$\frac{1}{C} = \frac{(\langle T_{\text{kin}}^2 \rangle - \langle T_{\text{kin}} \rangle^2)}{\langle T_{\text{kin}} \rangle^2} \approx \frac{(\langle T_{\text{eff}}^2 \rangle - \langle T_{\text{eff}} \rangle^2)}{\langle T_{\text{kin}} \rangle^2}$$

L. Stodolsky PRL.75.1044

$$C_v = \frac{C}{\Delta}$$

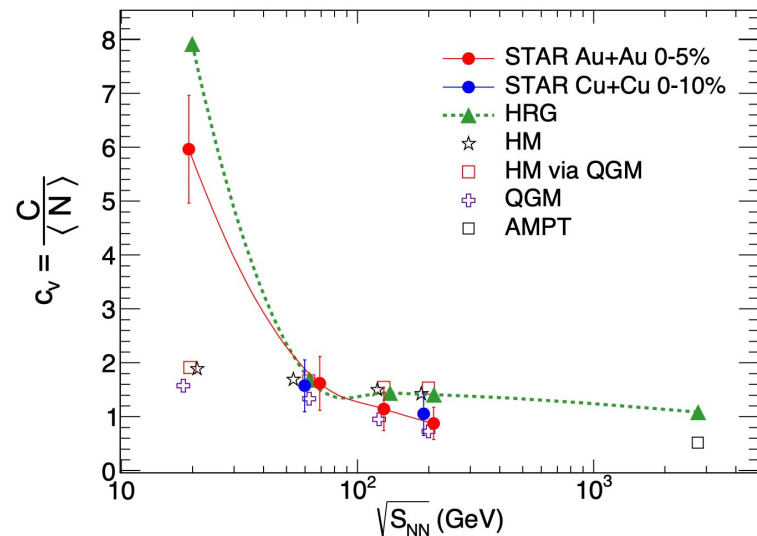
T.K.Nayak et al. PRC.94, 044901 (2016)

- ❖ In lattice calculations,  $\Delta = VT^3$ , although in experiments it's simpler to measure  $C/N$  where  $N$  is the charged particle multiplicity ( $N_{\text{ch}}$ ).

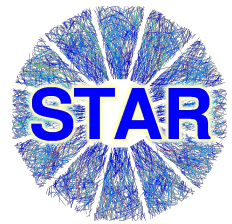


# Brief Overview

- ❖  $C_V$  has been calculated with STAR published data at energies 19.6, 62.4, 130 and 200 GeV.
- ❖ The experimental values match the prediction from the **HRG Model**.
- ❖ Monotonic increase coming into effect and the analysis at 3 GeV is critical to understand this trend.
- ❖ Previous CF talks :
  - [https://drupal.star.bnl.gov/STAR/system/files/CF-08\\_11\\_22.pdf](https://drupal.star.bnl.gov/STAR/system/files/CF-08_11_22.pdf)
  - [https://drupal.star.bnl.gov/STAR/system/files/Copy%20of%20CF-09\\_13\\_22.pdf](https://drupal.star.bnl.gov/STAR/system/files/Copy%20of%20CF-09_13_22.pdf)
  - <https://drupal.star.bnl.gov/STAR/system/files/CF%20-%20Dec822-Rutik.pdf>
  - <https://drupal.star.bnl.gov/STAR/system/files/Collaboration%20Meeting%20@%20BNL%20Feb23.pdf>



T.K.Nayak et al. PRC.94, 044901 (2016)



# Dataset and Event Cuts

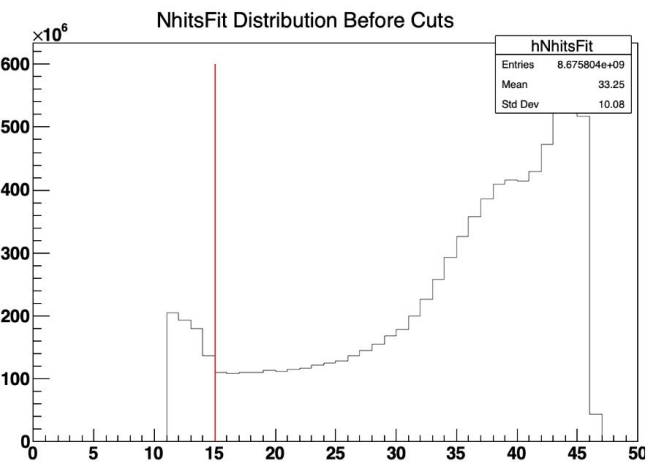
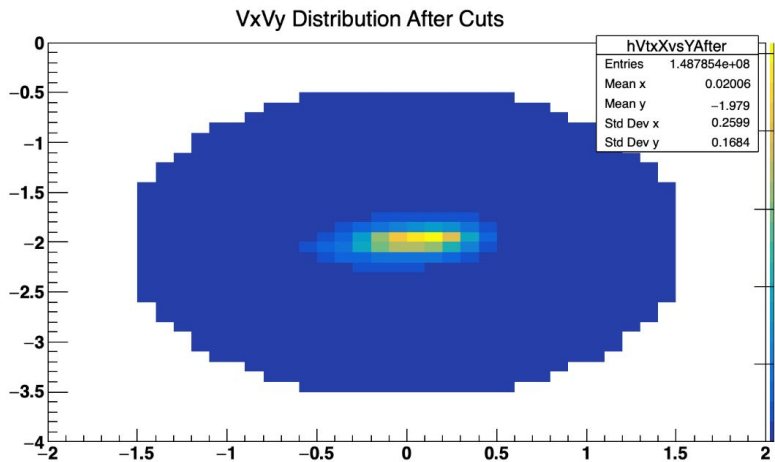
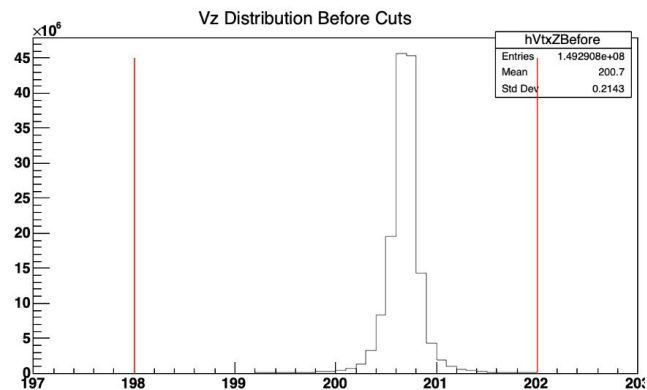
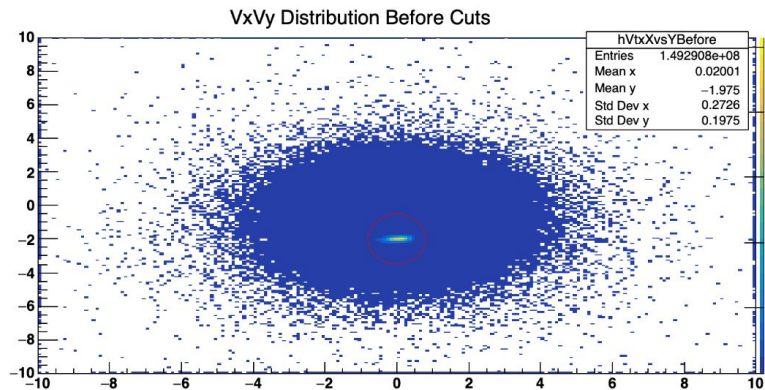
## Acceptance Cuts:

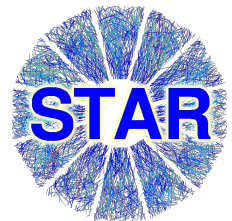
- ❖ Fixed Target, run 2018 data production, 3 GeV Au+Au collision,  $y_{\text{c.m.}} \approx 1.05$
- ❖ Events Cuts:
  - $198 < V_z < 202$  cm
  - $V_r < 1.5$  cm about beam spot centered around  $[0, -2]$ .
  - Trigger ID 620052 and 620053 (Min.Bias)
- ❖ Track Cuts:
  - $\text{DCA} < 3.0$  cm
  - $\text{NhitsFit}/\text{NHitsMax} > 0.51$
  - $\text{NhitsFit} > 15$

- ❖ Good Run List for Fluctuation analyses (149 M HLT Good):  
19153029, 19153031, 19153033, 19153034, 19153035, 19153036, 19153037, 19153042, 19153043, 19153044, 19153050, 19153051, 19153052, 19153053, 19153054, 19153055, 19153056, 19153057, 19153058, 19153059, 19153061, 19153062, 19153063, 19153064, 19153066, 19154001, 19154002, 19154005, 19154007, 19154027, 19154028, 19154029, 19154030, 19154031, 19154032, 19154036, 19154037, 19154038, 19154039, 19154040, 19154041, 19154044, 19154045, 19154046, 19154047, 19154048, 19154049, 19154052, 19154053, 19154054, 19154055, 19154056, 19154057, 19154058, 19154061, 19154063, 19154064, 19154065, 19154066, 19154067, 19155001, 19155003, 19155004, 19155005, 19155006, 19155008, 19155009, 19155010, 19155011, 19155016, 19155017, 19155018, 19155019, 19155020, 19155021, 19155022



# QA Plots

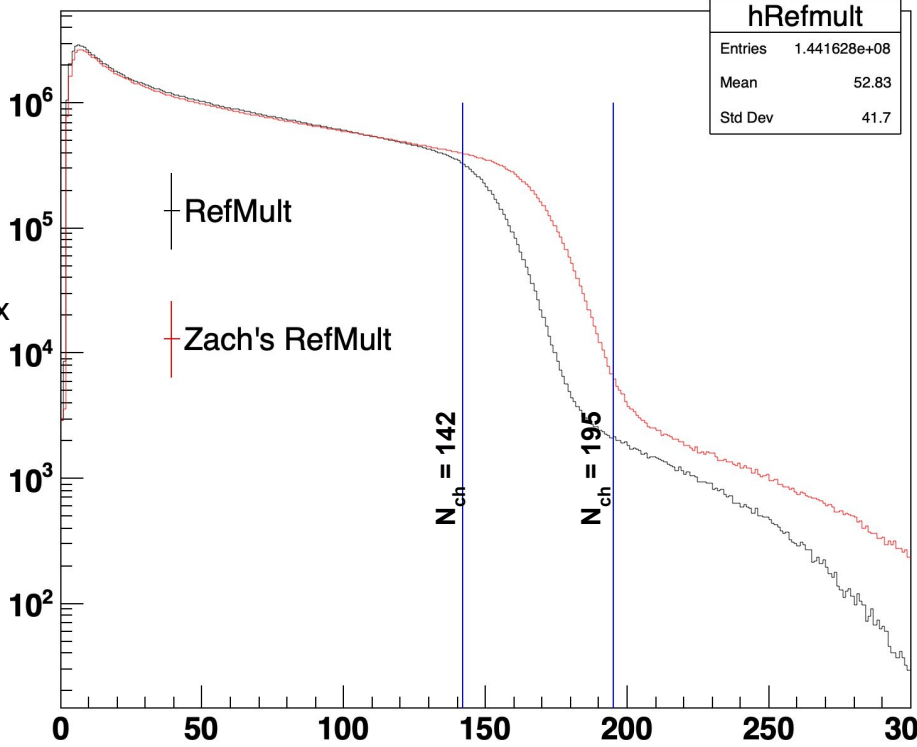




# QA Plots

hRefMult

hRefmult	
Entries	1.441628e+08
Mean	52.83
Std Dev	41.7



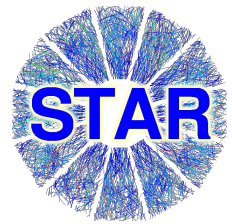
Centrality Low	Centrality High	Number of Events
0	5	1.04E+07
5	10	1.06E+07
10	15	1.01E+07
15	20	9.82E+06
20	25	1.05E+07
25	30	1.02E+07
30	35	9.61E+06
35	40	9.75E+06
40	45	9.85E+06
45	50	9.92E+06
50	55	8.17E+06
55	60	9.51E+06
60	65	9.00E+06
65	70	7.74E+06
70	75	5.42E+06
75	80	4.87E+06

Zach's RefMult:

- All primary tracks

My RefMult:

- All primary tracks
- DCA < 3.0
- NhitsFit/NHitsMax > 0.51



# Closure Test for Unfolding Method

- ❖ A Closure test was performed on UrQMD Data using Unfolding
- ❖ For the Closure test, the detector like events had to be made:

Step 1

Calculate efficiency as a function of  $P_{id}$  and  $p_T$ , for a track. ( $\epsilon$ )

Step 2

Generate a random number between  $[0,1]$ . ( $r$ )

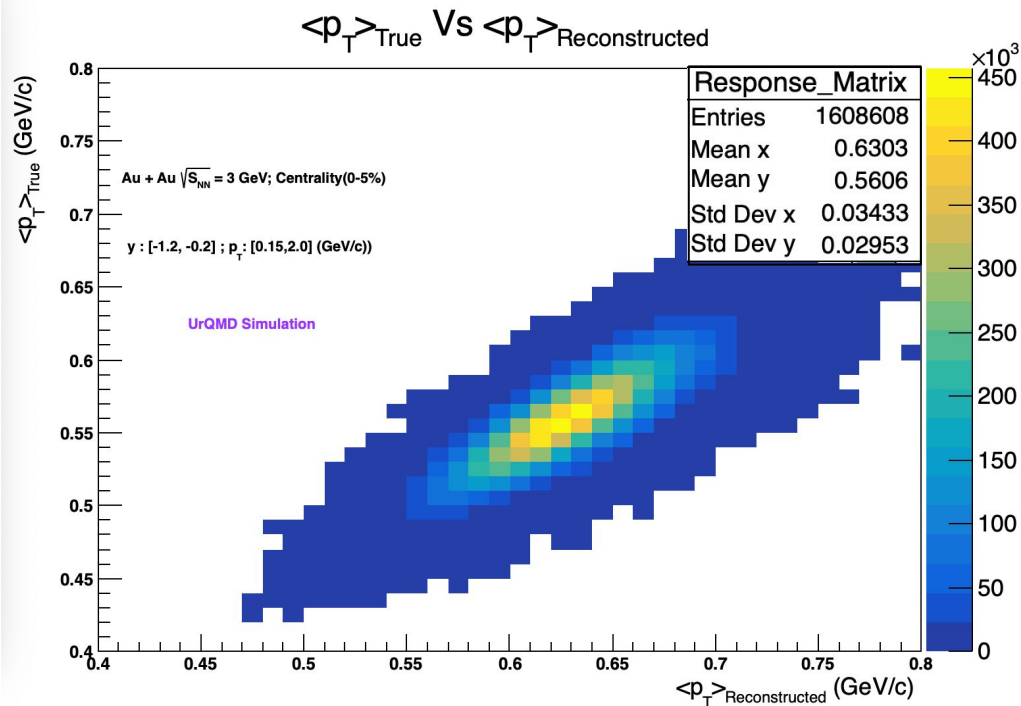
Step 3

Keep the particle if efficiency is greater than the random number. ( $r < \epsilon$ )

**Detector Like Events from UrQMD**



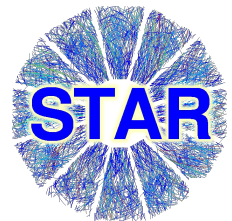
# Closure Test for Unfolding Method



The response matrix was calculated from UrQMD at 3 GeV, using the same cuts as the analysis:

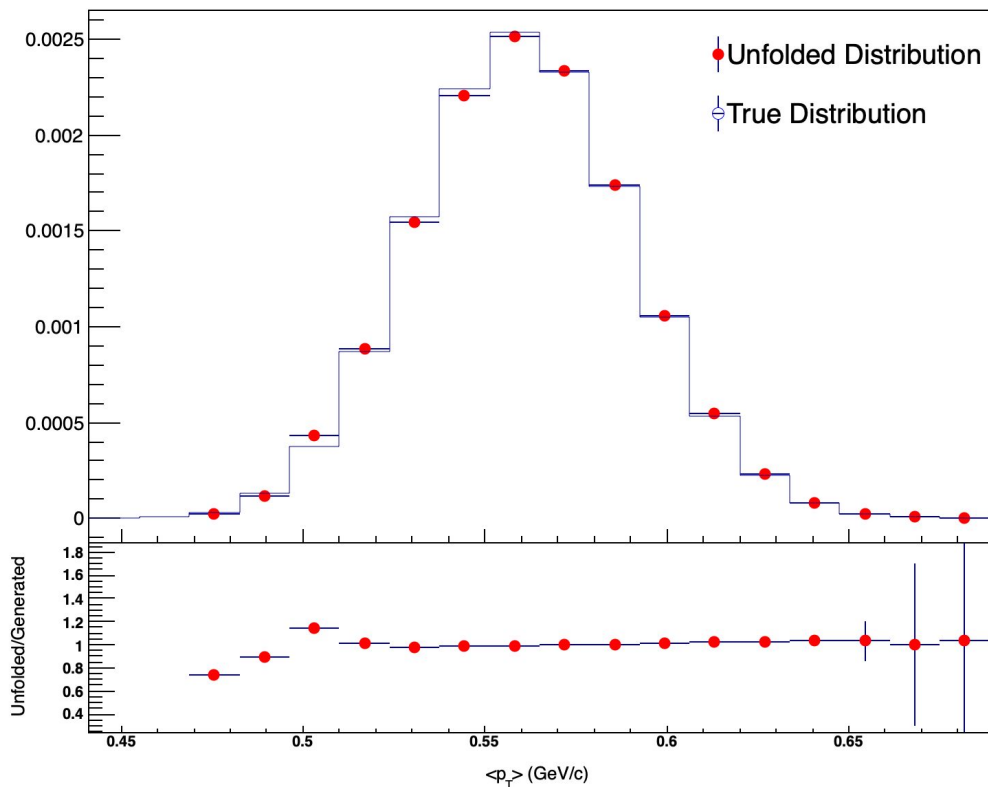
- ❖  $0.15 < p_T < 2.0$  (GeV/c)
- ❖ All primary Charged particles.
- ❖  $-1.2 < y < -0.2$
- ❖ 56% of total data (from UrQMD)
- ❖ 40 bins between 0.4-0.8 (GeV/c)
- ❖ With the remaining 44% Closure test was performed.





# Closure Test for Unfolding Method

Bayesian Unfolding

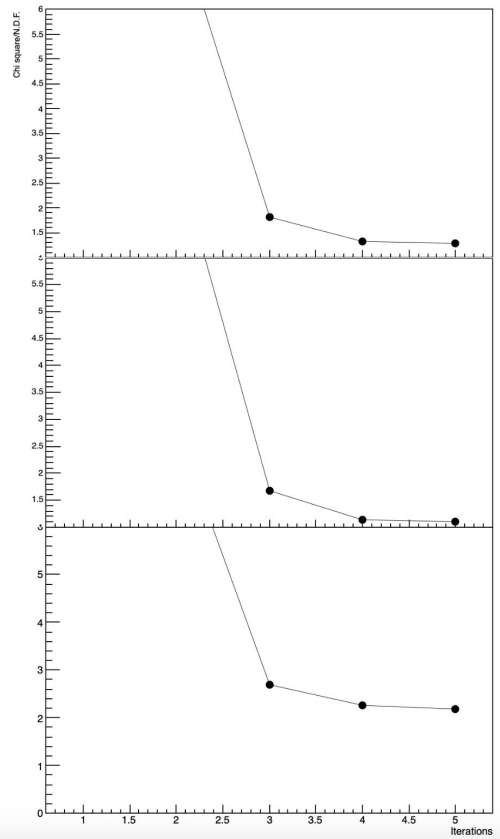
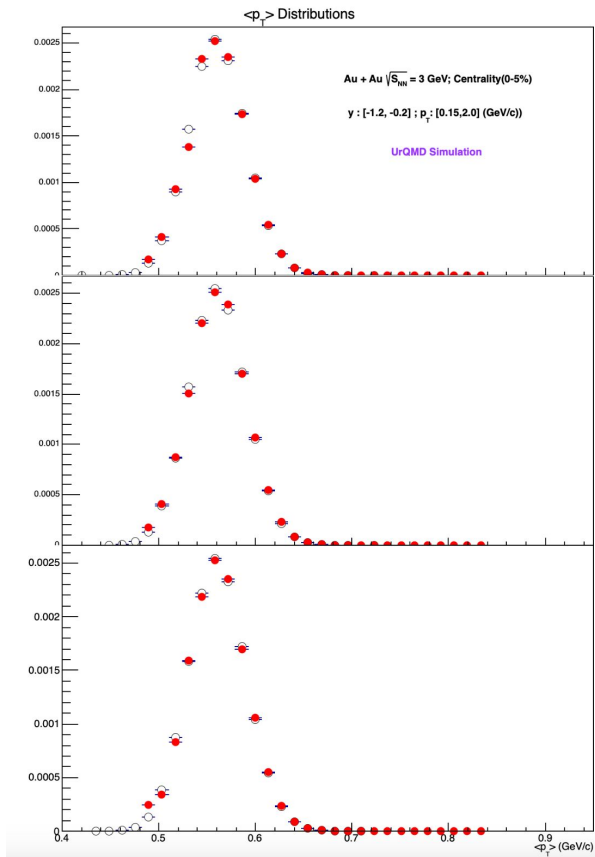


Statistical Bootstrap for the Closure Test, divided the testing set into 10 samples.

	Generated	Unfolded
Mean	$0.5606 \pm 2e-4$	$0.5608 \pm 3e-4$
Sigma	$0.02949 \pm 2e-4$	$0.02984 \pm 3e-4$

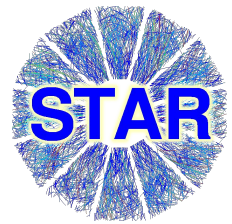


# Closure Test for Unfolding Method



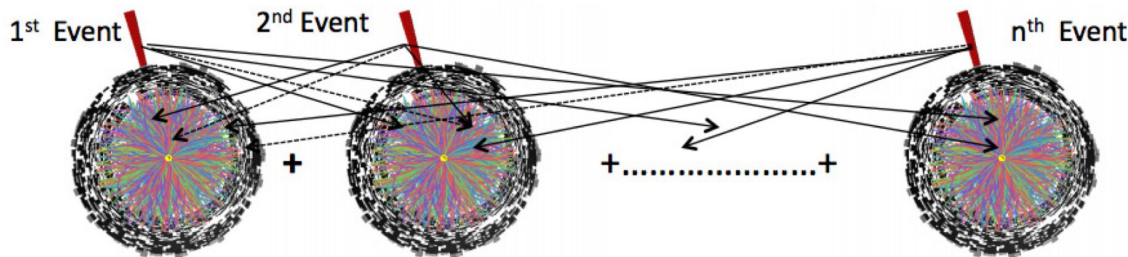
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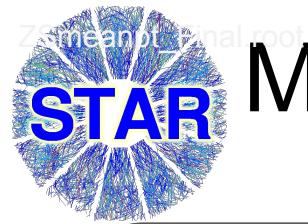


# Mixed Event Analysis for $\langle p_T \rangle$

- ❖ In order to establish whether the observed fluctuations are partly dynamical in nature, we need to disentangle statistical effects i.e. effects due to the finite number of particles in the final state of the collision.
- ❖ The Mixed event construction makes synthetic events with tracks from different events to remove any kind of correlations.



The schematic representation of mixed events

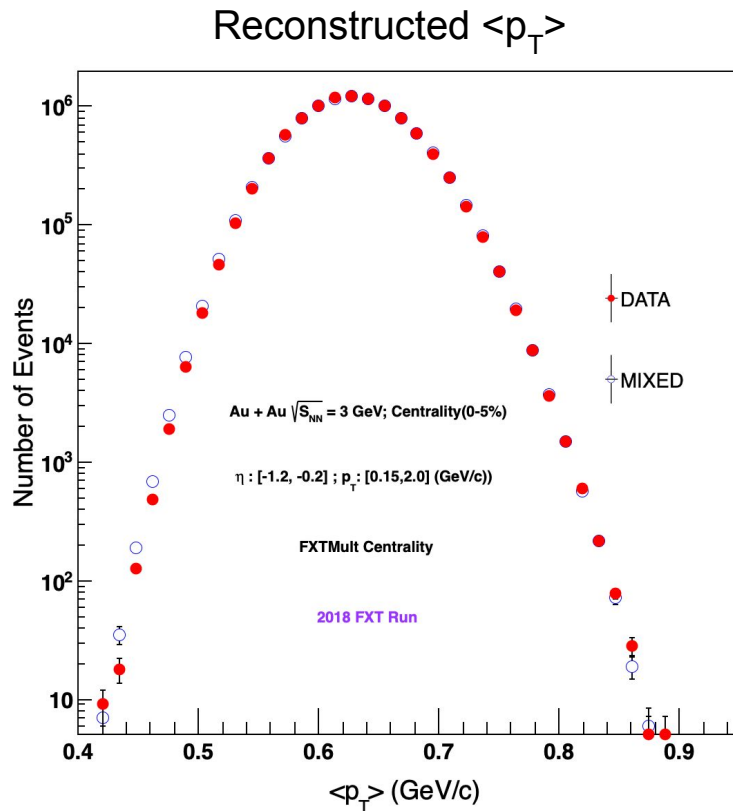


# Mixed Event Analysis for $\langle p_T \rangle$

**FXTMult**  $\langle p_T \rangle$  Distributions,

- ❖ All primary Charged particles.
- ❖  $0.15 < p_T < 2.0$  (GeV/c)
- ❖  $-1.2 < \eta < -0.2$

\*10 M Events





# Dynamical Fluctuations from $\langle p_T \rangle$

We fit the  $\langle p_T \rangle$  distributions with the gamma function to obtain the mean and sigma

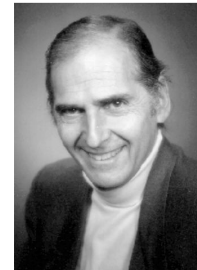
From the fit parameters  $\alpha, \beta$   
We can calculate  $\mu$  and  $\sigma$ :

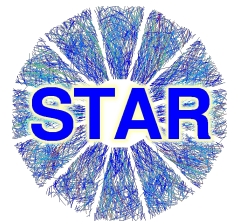
- ❖  $\mu = \alpha\beta$
- ❖  $\sigma^2 = \alpha\beta^2$

$$\sigma_{dyn} = \sqrt{\left(\frac{\sigma_{data}}{\mu_{data}}\right)^2 - \left(\frac{\sigma_{mix}}{\mu_{mix}}\right)^2}$$

$$f(x) = \frac{x^{\alpha-1} e^{-\frac{x}{\beta}}}{\Gamma(\alpha)\beta^\alpha}$$

It's not a  
Gaussian...it's a  
Gamma distribution!





# Dynamical Fluctuations from $\langle p_T \rangle$

Data\_Mean =  $0.6289 \pm 0.001$   
 Data\_Sigma =  $0.0449 \pm 7e-5$

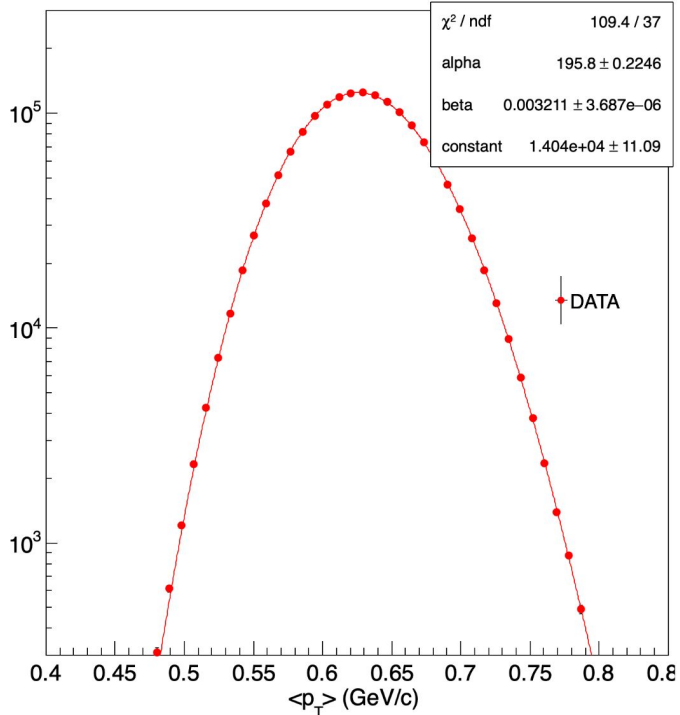
Mixed\_Mean =  $0.6287 \pm 0.001$   
 Mixed\_Sigma =  $0.0454 \pm 7e-5$

## FXTMult $\langle p_T \rangle$ Distributions,

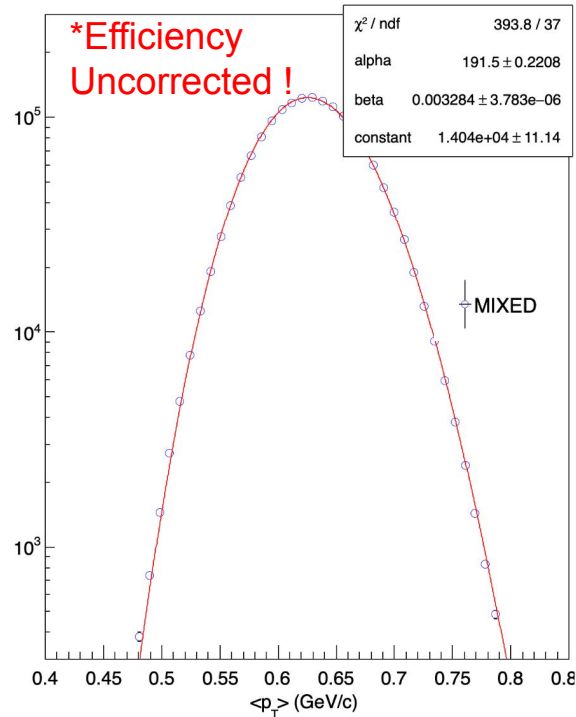
All charged particles.

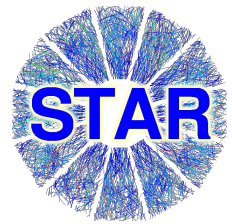
- ❖ Primary Charged particles.
- ❖  $0.15 < p_T < 2.0$  (GeV/c)
- ❖  $-1.2 < \eta < -0.2$

Reconstructed  $\langle p_T \rangle$



Reconstructed  $\langle p_T \rangle$





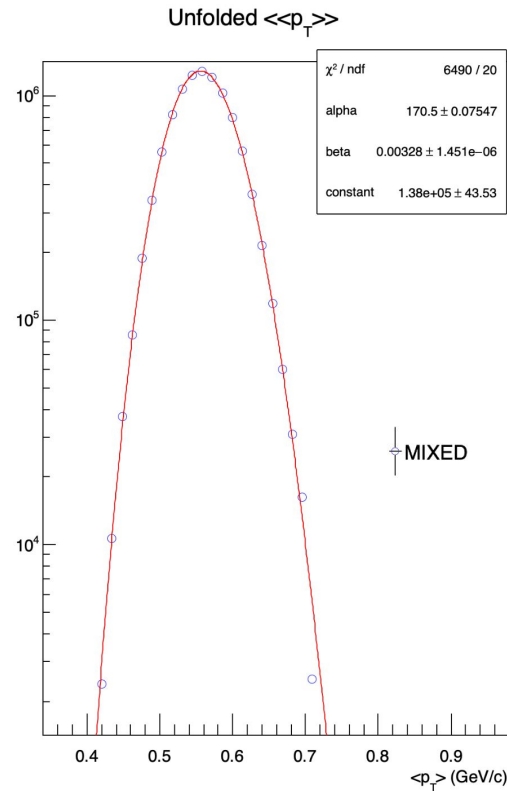
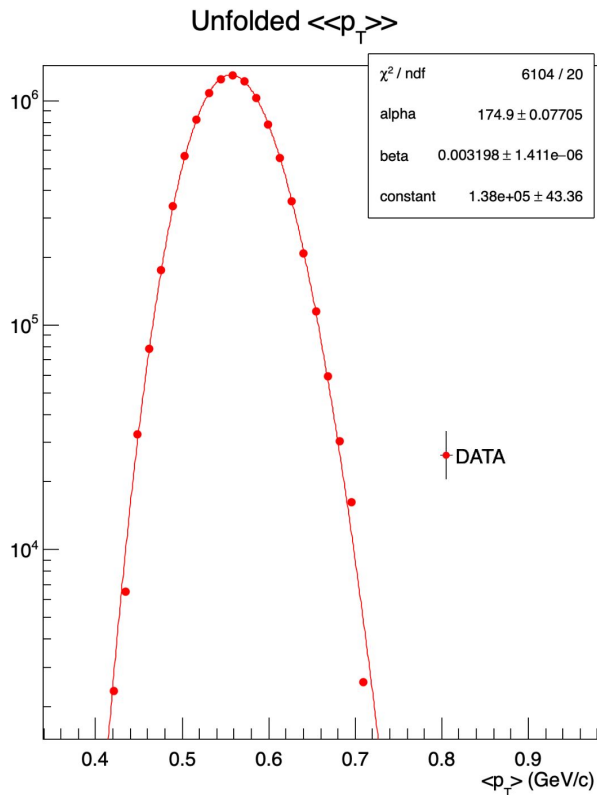
# Dynamical Fluctuations from $\langle p_T \rangle$

Data\_Mean =  $0.5593 \pm 0.004$   
Data\_Sigma =  $0.0423 \pm 2e-5$

Mixed\_Mean =  $0.5593 \pm 0.004$   
Mixed\_Sigma =  $0.0428 \pm 3e-5$

**FXTMult**  $\langle p_T \rangle$  Distributions,  
All charged particles.

- ❖ Primary Charged particles.
- ❖  $0.15 < p_T < 2.0$  (GeV/c)
- ❖  $-1.2 < \eta < -0.2$



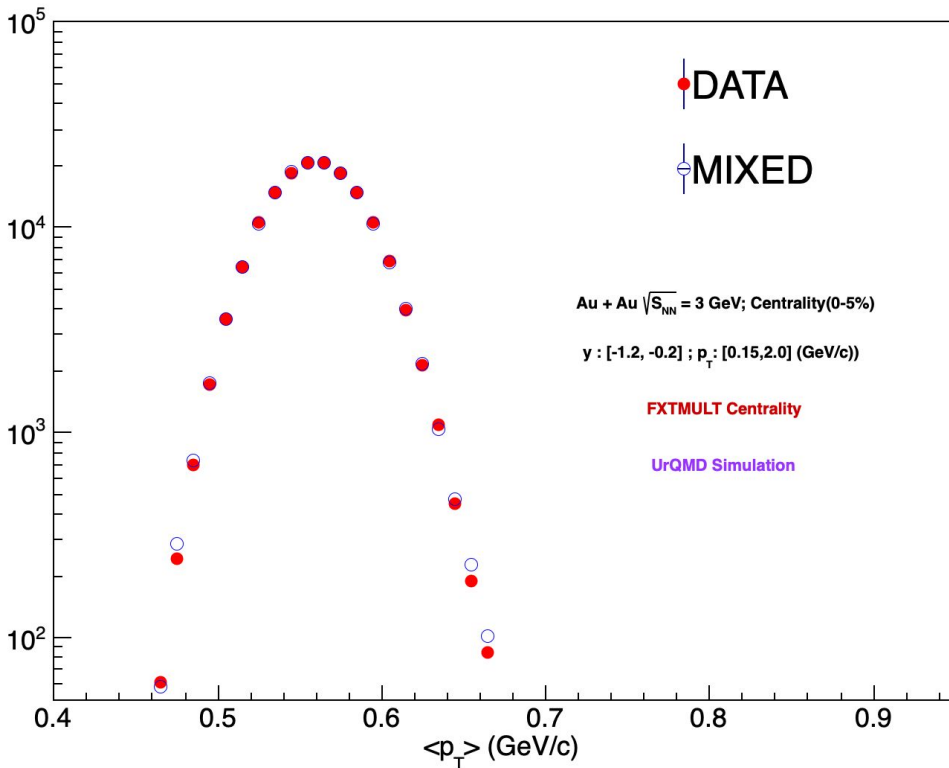


# Dynamical Fluctuations from $\langle p_T \rangle$

$\langle\langle p_T \rangle\rangle$  Distributions

- ❖ **UrQMD** Mixing of events, 150K events
- ❖ 0-5%  $\rightarrow$  [0.0, 3.45] fm (b)
- ❖ **No detector effects taken into account**

	Data	Mixed
<b>Mean</b>	0.5608	0.5607
<b>Sigma</b>	0.02961	0.02975







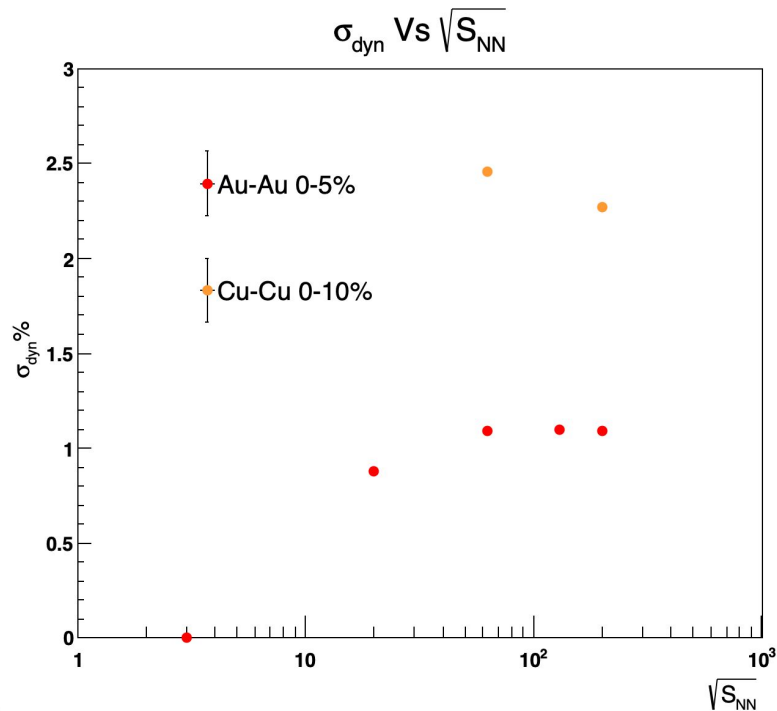
# Comparison to Published data

Case	$\mu$	$\sigma$
20 GeV, real	0.5228	0.01579
20 GeV, mixed	0.5227	0.01510
62 GeV, real	0.5471	0.01439
62 GeV, mixed	0.5470	0.01310
130 GeV, real	0.5614	0.01423
130 GeV, mixed	0.5612	0.01282
200 GeV, real	0.5799	0.01347
200 GeV, mixed	0.5799	0.01190

PHYSICAL REVIEW C 72, 044902 (2005)

**\*This work**

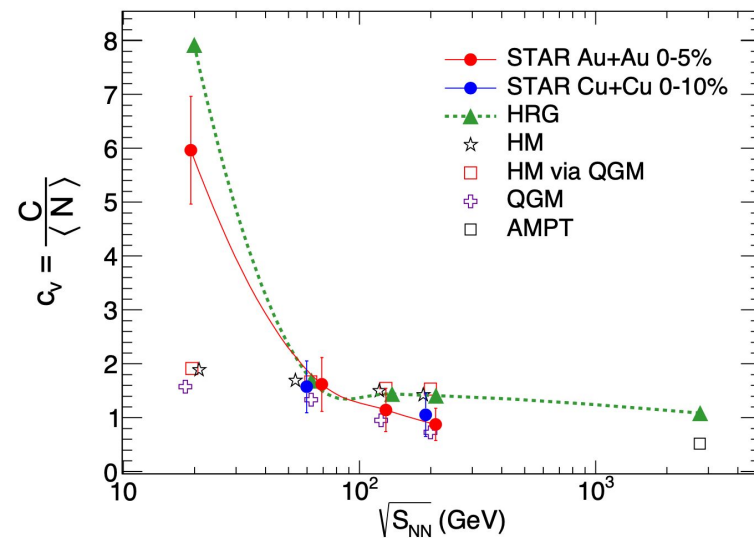
	$\mu$	$\sigma$
3 GeV, real	0.5593	0.0423
3 GeV, mixed	0.5593	0.0428





# STAR Outlook

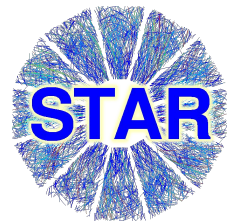
1. Study the effect of primordial protons on  $\langle p_T \rangle$  fluctuations.
2. Study the eta window acceptance effect.
3. Error calculation, systematics



T.K.Nayak et al. PRC.94, 044901 (2016)

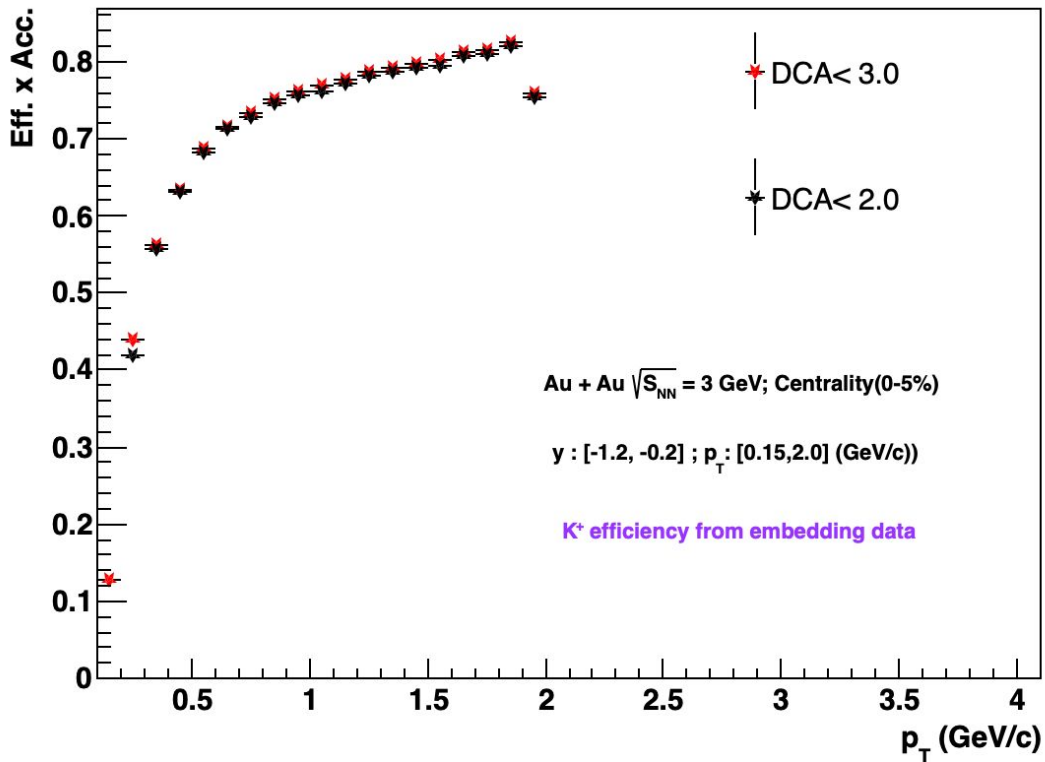


Thank you for your attention



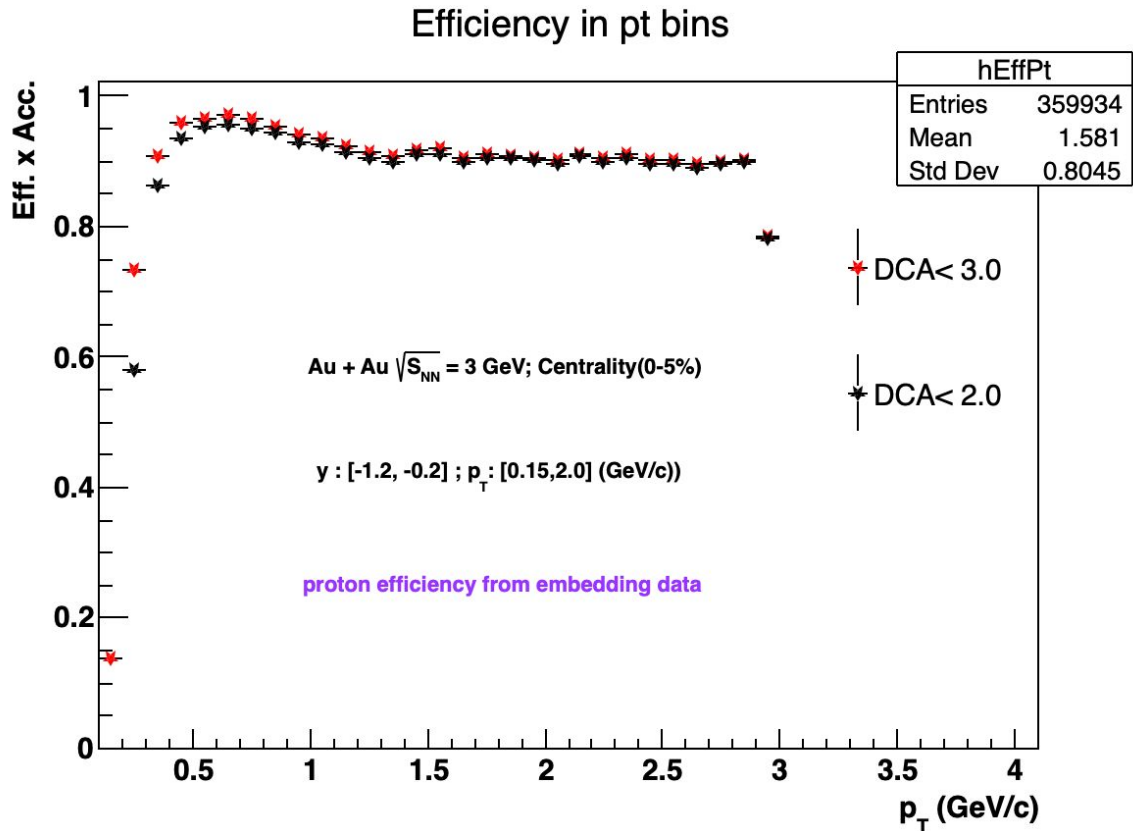
# Efficiency plots from embedding

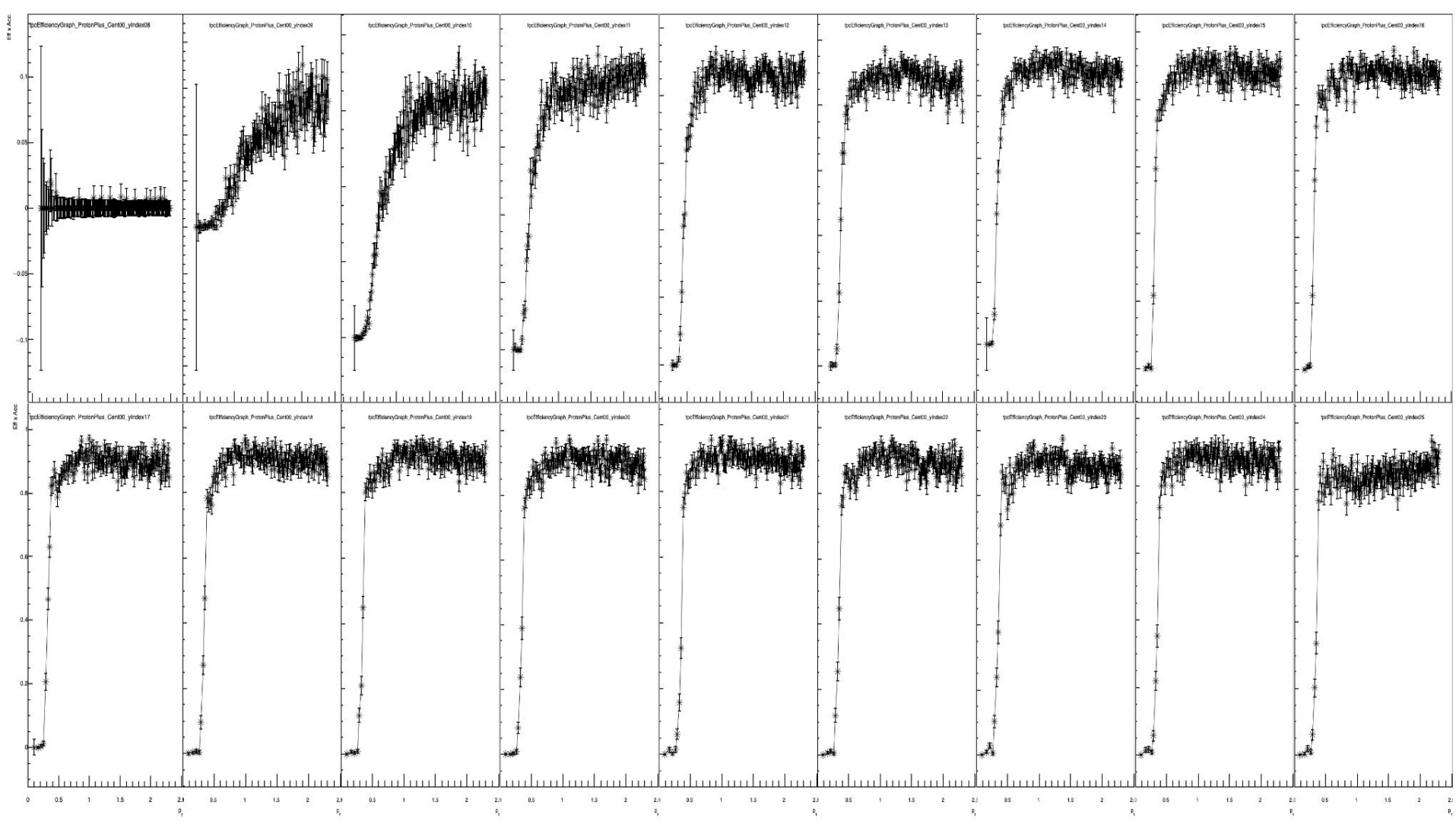
Efficiency in pt bins





# Efficiency plots from embedding





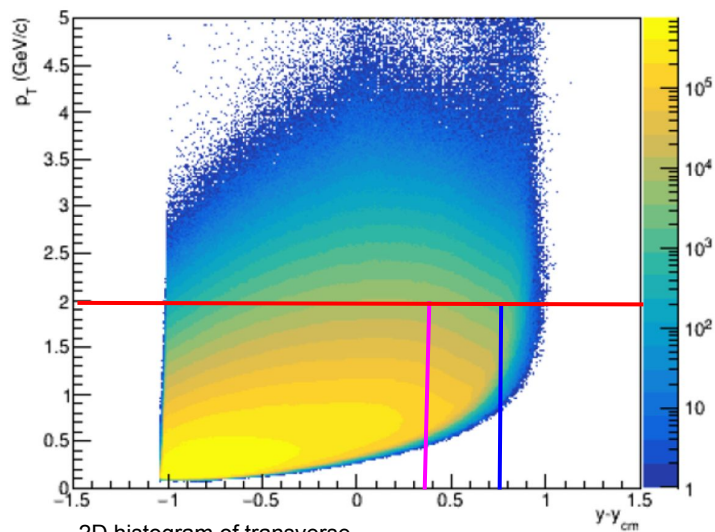


# Eta Study

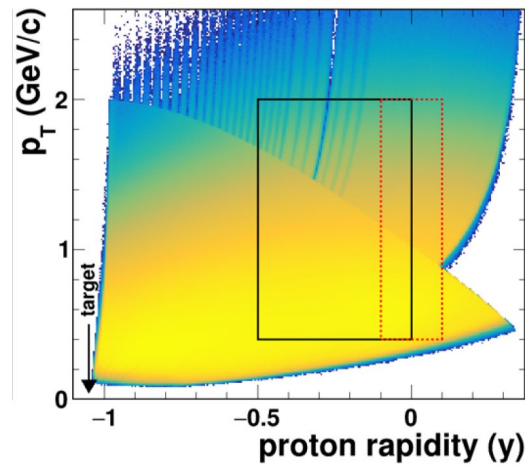
To suppress the spectator protons from entering the analysis, the maximum rapidity range is restricted to  $-0.5 < y < 0$ .

$$y_{\text{cm}} = 1.05$$

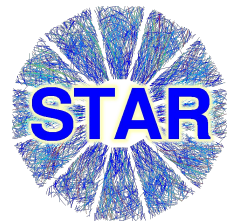
Physical Review Letters 128, 202303 (2022)



2D histogram of transverse momentum vs proton rapidity for proton tracks in the TPC.





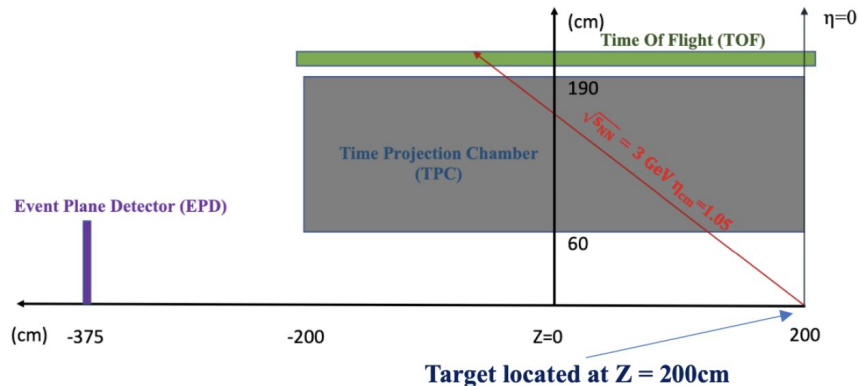


# Eta Study

Unlike collider mode collisions, the target is located at the edge of TPC, and the midrapidity is not zero in the FXT model collisions in the STAR coordinate system. In order to convert the measured rapidity  $y$  in the coordinate system to the rapidity  $y$  in the center of mass frame. One need to boost the measured rapidity by beam rapidity.

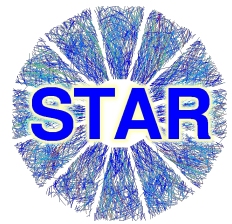
In our STAR convention, the beam-going direction is the positive direction (the target is located in the negative rapidity direction  $y_{\text{target}} = -1.045$ ).

In order to match the STAR conventions, when calculating rapidity in center of mass frame, in addition to shift by midrapidity, we also need to flip the sign of rapidity.



$$y_{\text{CM}} = -(y_{\text{lab}} - y_{\text{mid}})$$

In the STAR coordinate system, the target is located at Z=200cm the edge of TPC. The EPD is located at -375 cm. Red line indicates  $\eta_{\text{cm}} = 1.05$  and it is roughly midrapidity region.



# Eta Study

To suppress the spectator protons from entering the analysis, the maximum rapidity range is restricted to  $-0.5 < y < 0$ .

The target is located at  $y = -1.05$

Physical Review Letters 128, 202303 (2022)

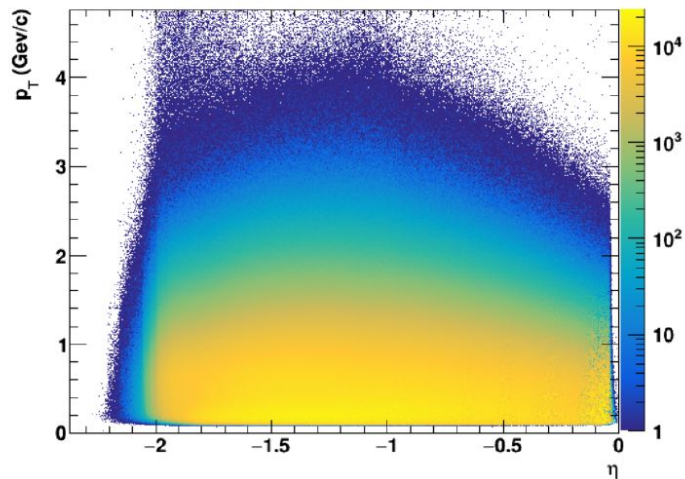
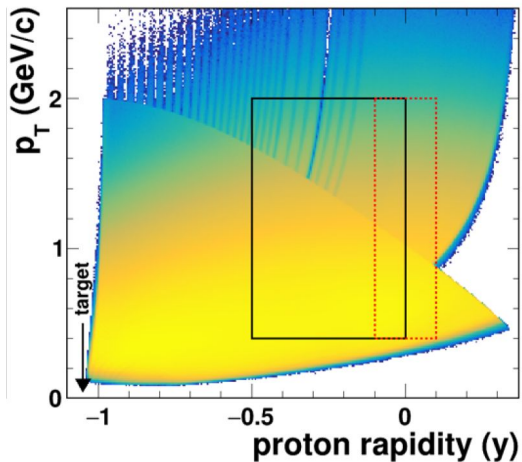


FIG. 1: Pseudorapidity acceptance of the TPC. For the FxtMult definition, primary tracks from  $-2 < \eta < 0$  with  $N_{HitsFit}/N_{HitsMax} > 0.51$  and tracks within  $198 < V_z < 202$  cm are included. FxtMult3 requires an additional cut of  $n\sigma_{Proton} < -3$  for positively charged tracks.



# Eta Study

To suppress the spectator protons from entering the analysis, the maximum rapidity range is restricted to  $-0.5 < y < 0$ .

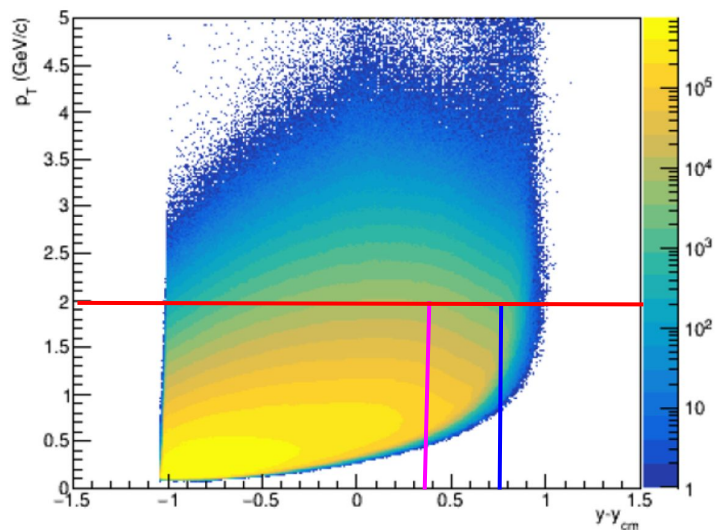
The target is located at  $y = -1.05$

Physical Review Letters 128, 202303 (2022)

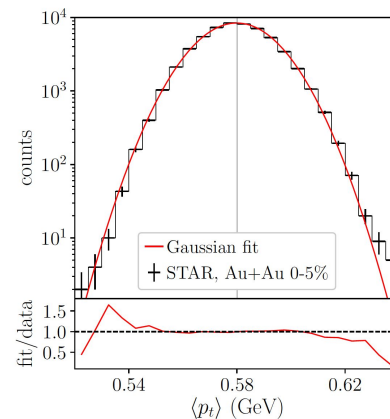
We predict, in particular, that  $p_T$  fluctuations have positive skew, which is significantly larger than if particles were emitted independently. We elucidate the origin of this result by deriving generic formulas relating the fluctuations of  $p_T$  to the fluctuations of the early-time thermodynamic quantities. We postulate that the large positive skewness of  $p_T$  fluctuations is a generic prediction of hydrodynamic models.

PHYSICAL REVIEW C 103, 024910 (2021)

FIG. 1. Distribution of  $\langle p_t \rangle$  for Au + Au collisions at  $\sqrt{s_{NN}} = 200$  GeV in the 0–5% centrality window. Data from the STAR Collaboration [2] are shown as a histogram. The solid line is a Gaussian fit to these data. The lower panel is the ratio between the Gaussian fit and the data. The data are above the Gaussian to the right and below the Gaussian to the left, which hints at a positive skew.



2D histogram of transverse momentum vs proton rapidity for proton tracks in the TPC.

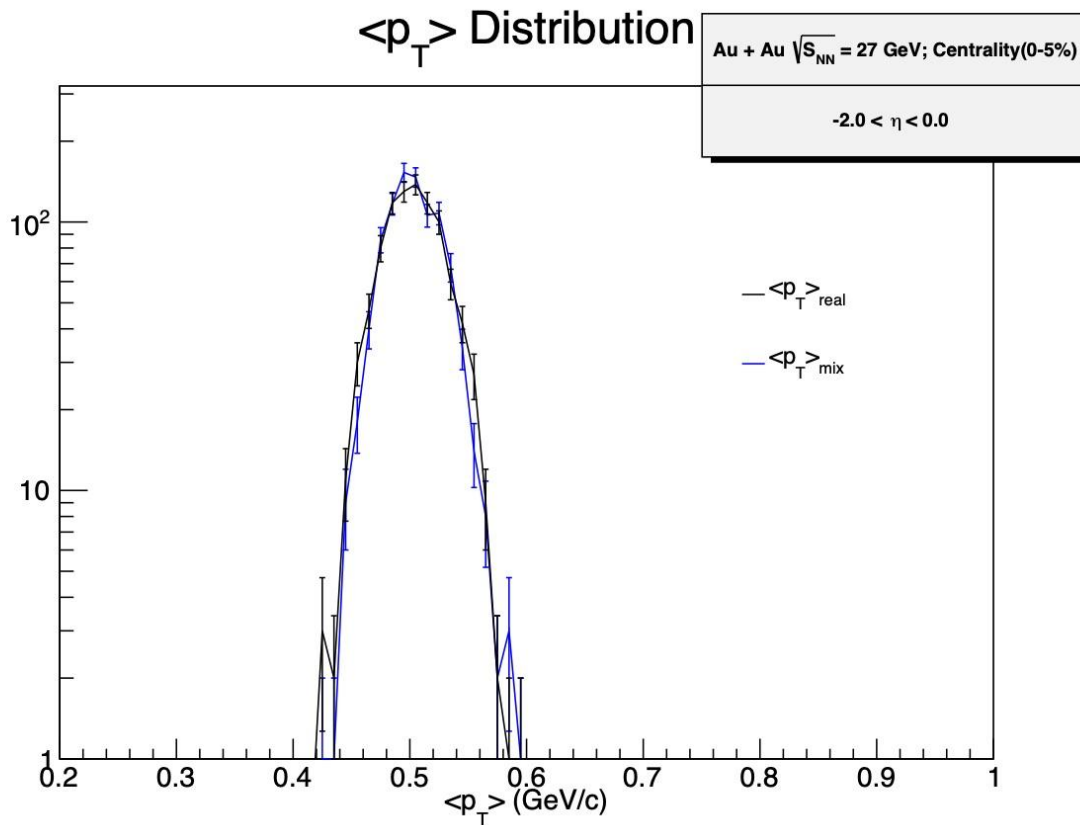


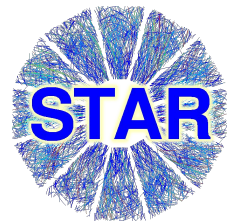


- ❖ A small subset of 27 GeV is taken for analysis.

- ❖ The event cuts and track cuts are from Chun-Jian.

([https://drupal.star.bnl.gov/SSTAR/system/files/BES\\_20054\\_27\\_meanpT\\_0119.pdf](https://drupal.star.bnl.gov/SSTAR/system/files/BES_20054_27_meanpT_0119.pdf))



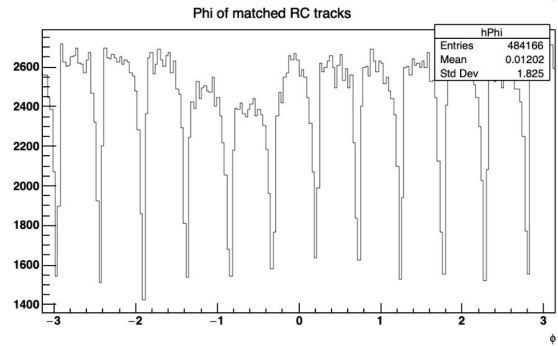
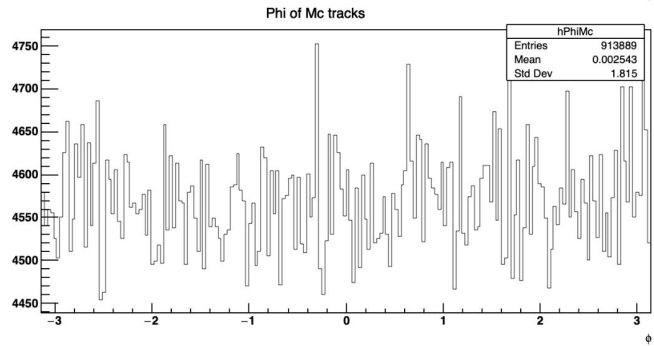
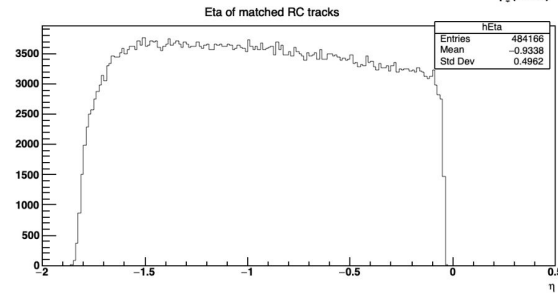
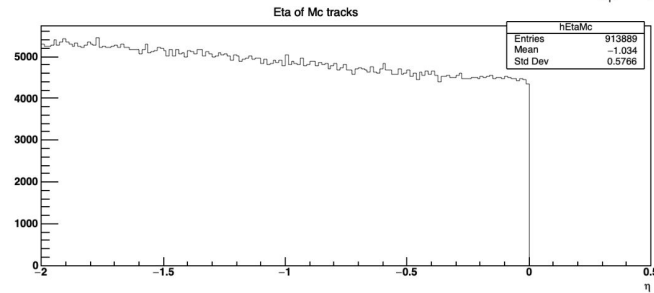
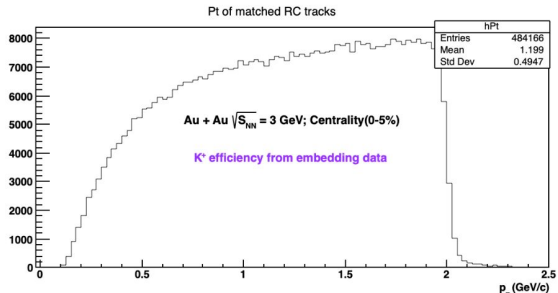
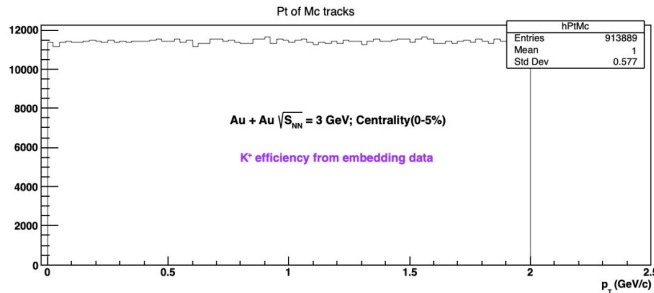


# Efficiency plots from embedding

- ❖ The changes in systematic cuts affect efficiency, hence have to calculate them for each set of systematic cuts.
- ❖ Embedding data are generally used in STAR experiments for detector acceptance & reconstruction efficiency study. In general, the efficiency depends on running conditions, particle types, particle kinematics, and offline software versions.
- ❖ Particles are generated flat in  $p_T$ ,  $\eta$  and  $\phi$  and then passed through the GEANT 3 simulation of the STAR Detector and then we calculate reconstructed and generated tracks in our acceptance to compute efficiencies.

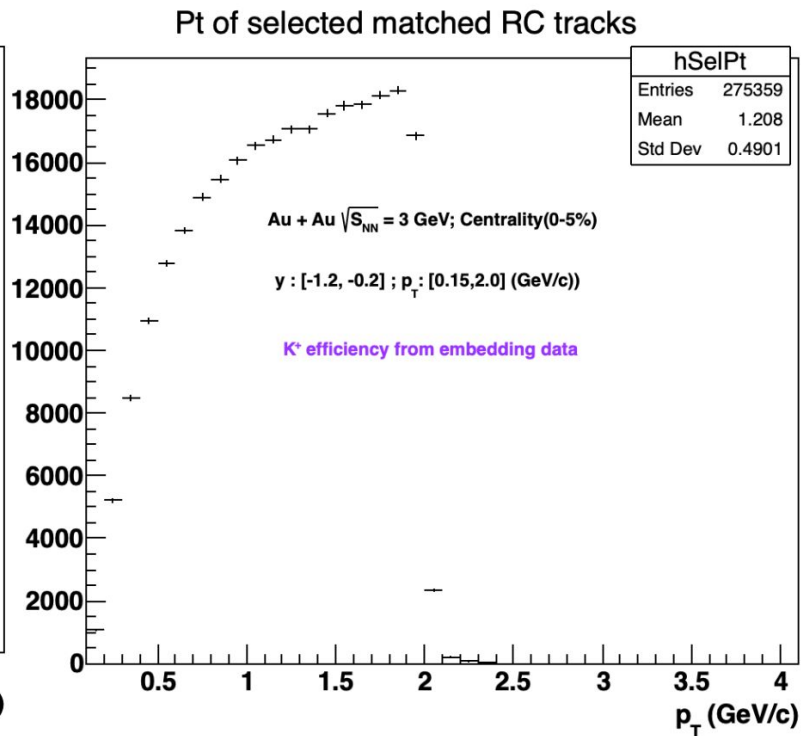
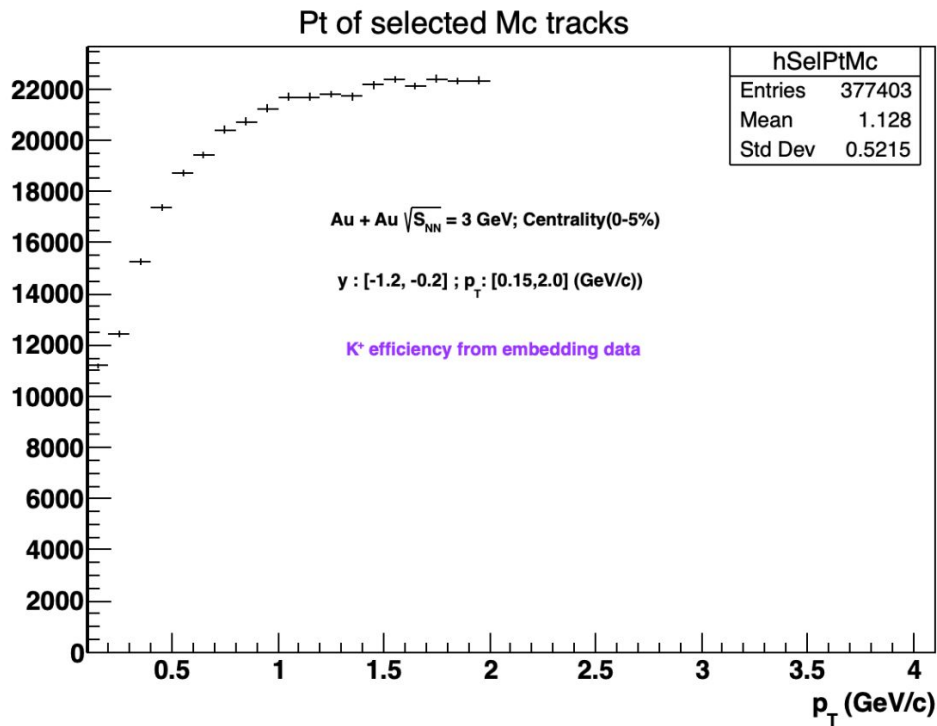


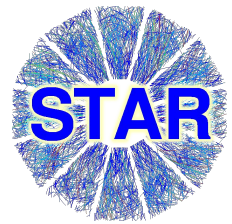
# Efficiency plots from embedding



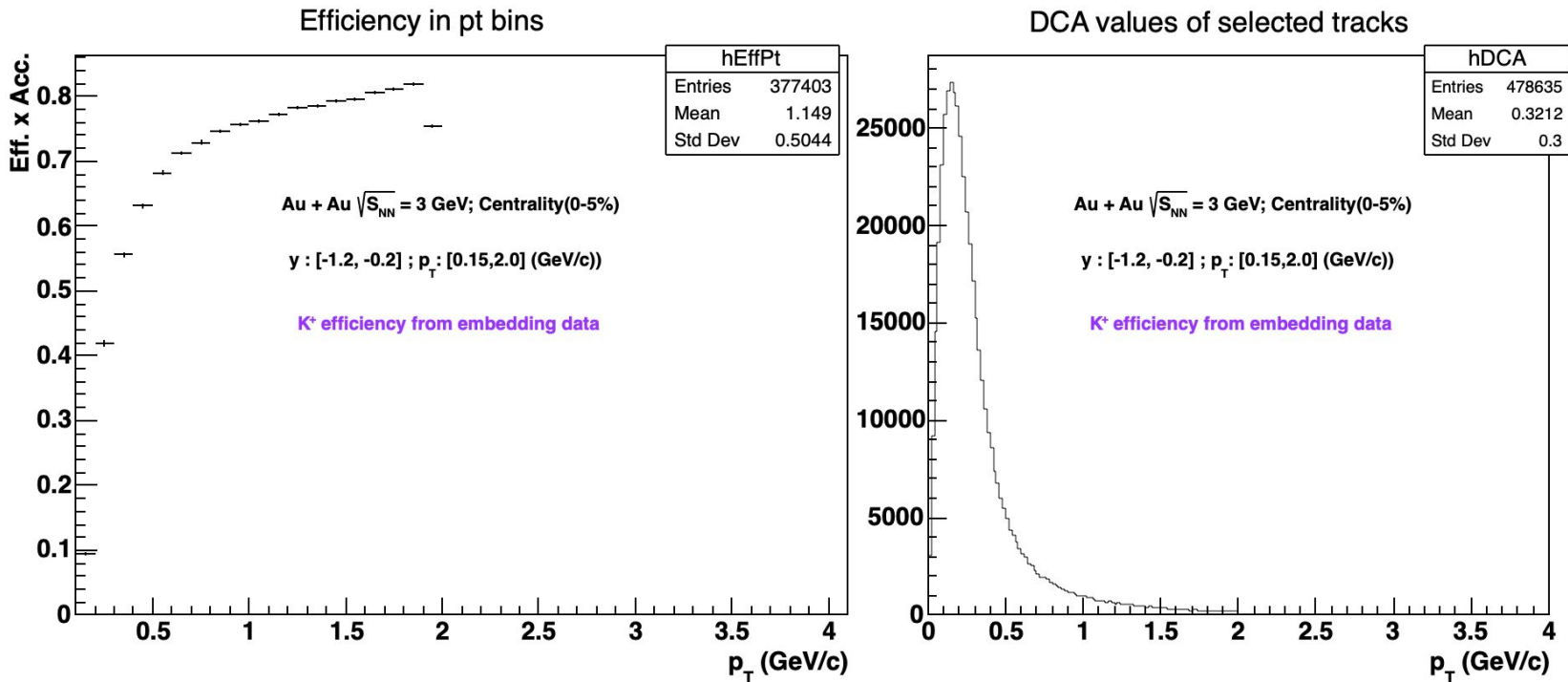


# Efficiency plots from embedding





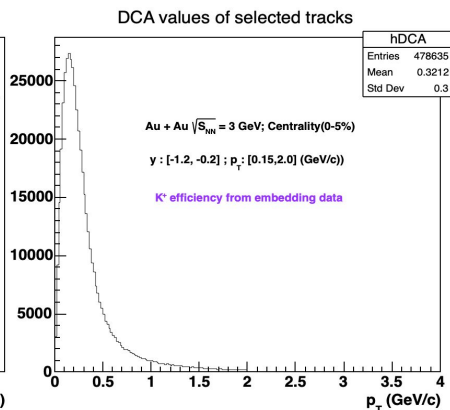
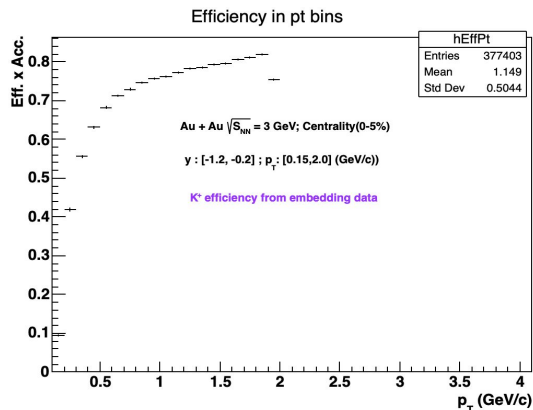
# Efficiency plots from embedding



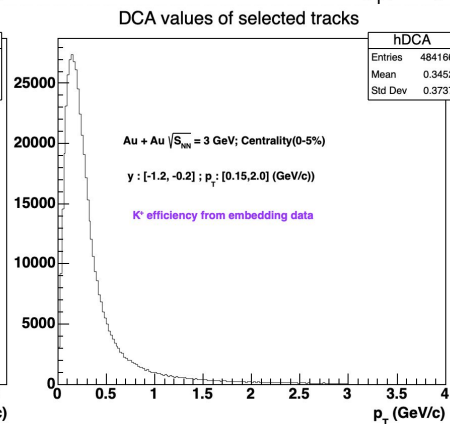
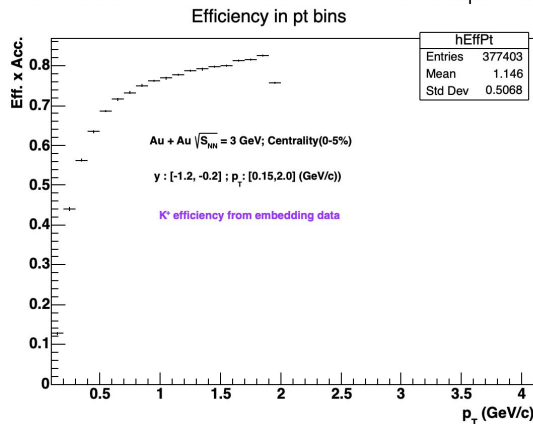




# Efficiency plots from embedding



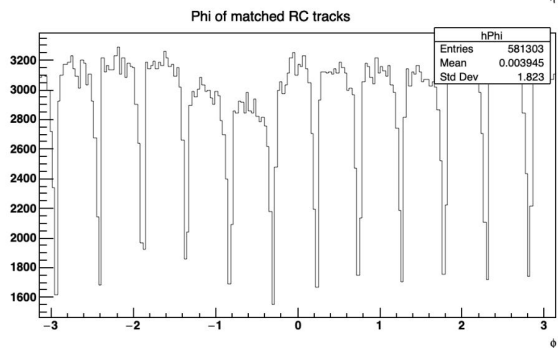
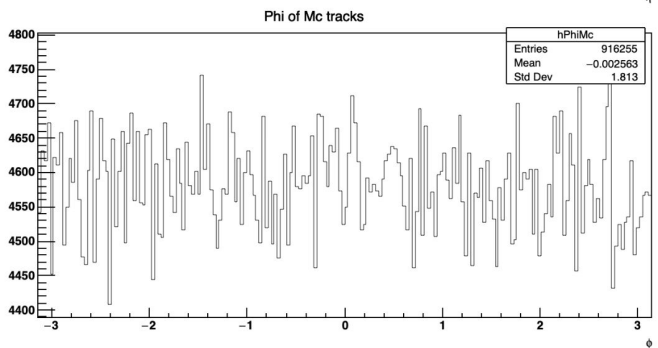
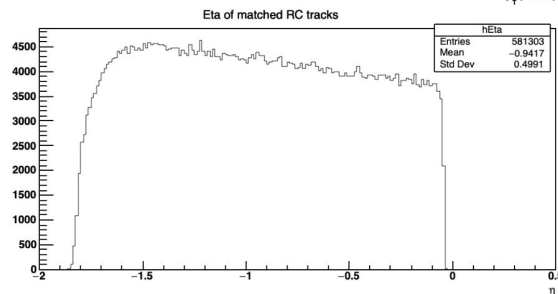
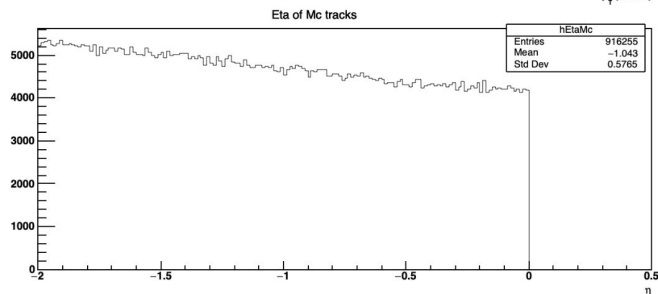
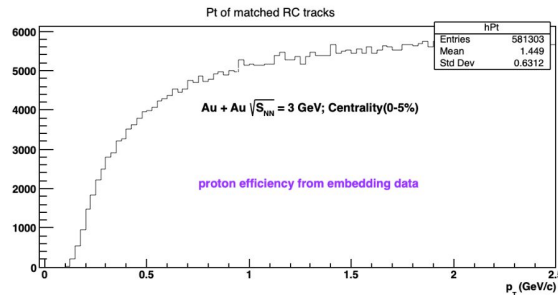
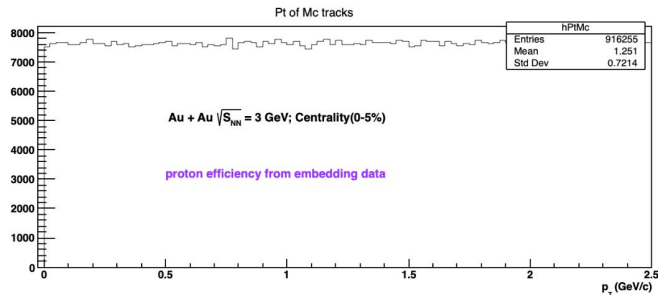
DCA < 2.0



DCA < 3.0



# Efficiency plots from embedding



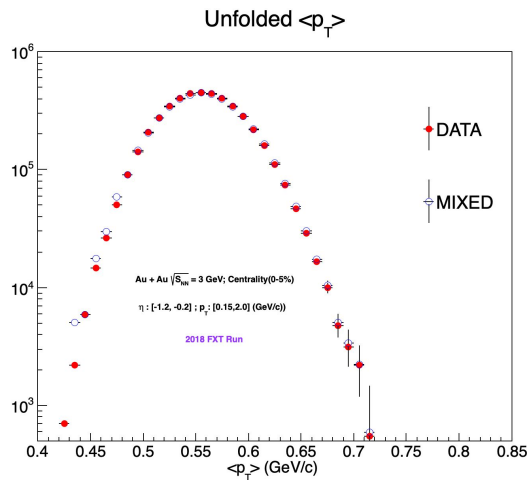


# Comparison to Published data

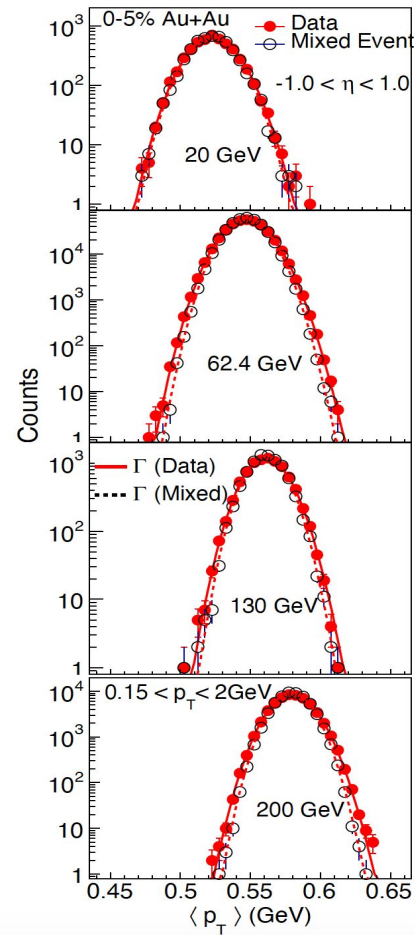
Case	$\mu$	$\sigma$
20 GeV, real	0.5228	0.01579
20 GeV, mixed	0.5227	0.01510
62 GeV, real	0.5471	0.01439
62 GeV, mixed	0.5470	0.01310
130 GeV, real	0.5614	0.01423
130 GeV, mixed	0.5612	0.01282
200 GeV, real	0.5799	0.01347
200 GeV, mixed	0.5799	0.01190

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	$\mu$	$\sigma$
3 GeV, real	0.5576	0.04038
3 GeV, mixed	0.5575	0.04111



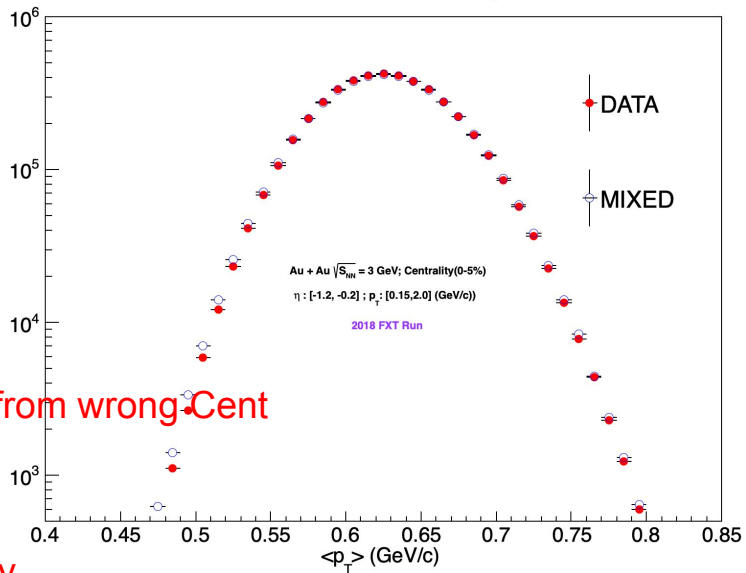
This is from wrong Cent def.





# Systematics

Reconstructed  $\langle p_T \rangle$



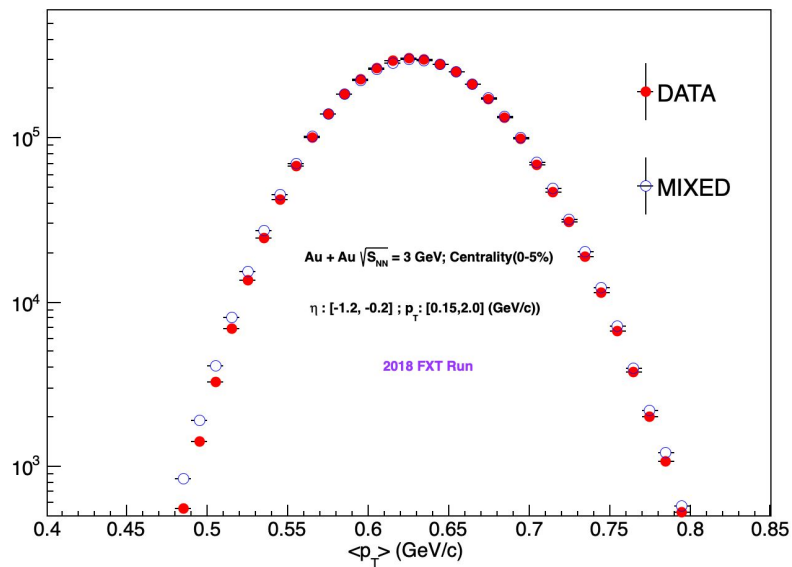
This is from wrong Cent def.

\*Efficiency  
Uncorrected !

DCA < 3.0 cm

	Data	Mixed
Mean	0.6271	0.6271
Sigma	0.04338	0.04402

Reconstructed  $\langle p_T \rangle$



DCA < 2.0 cm

	Data	Mixed
Mean	0.6304	0.6304
Sigma	0.04327	0.04409

CF PWG Meeting



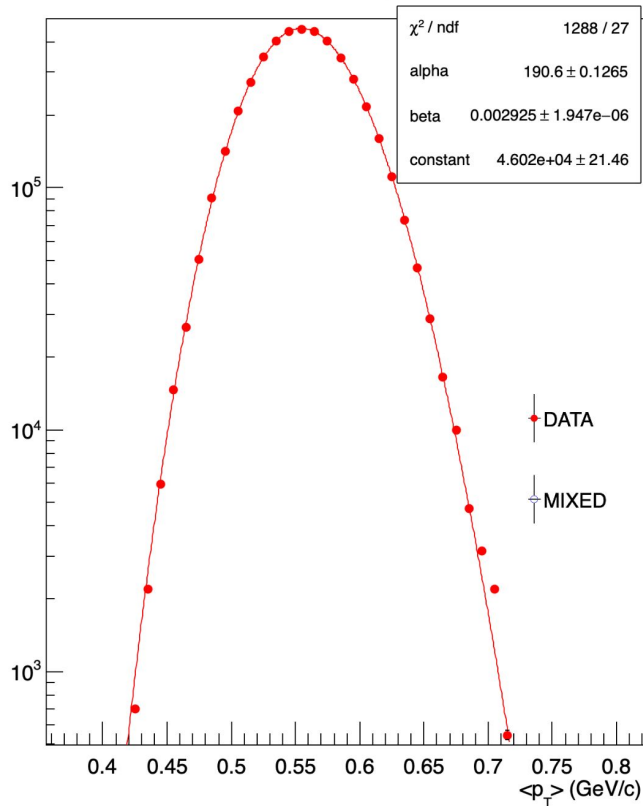
# Dynamical Fluctuations from $\langle p_T \rangle$

Data\_Mean =  $0.5576 \pm 0.0007$   
Data\_Sigma =  $0.04038 \pm 4e-5$

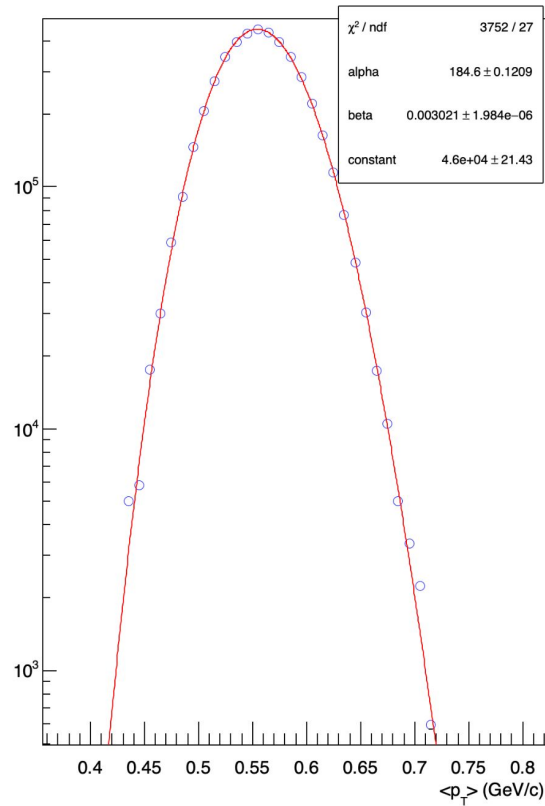
Mixed\_Mean =  $0.5576 \pm 0.0007$   
Mixed\_Sigma =  $0.04105 \pm 4e-5$

This is from wrong Cent  
def.

Unfolded  $\langle\langle p_T \rangle\rangle$



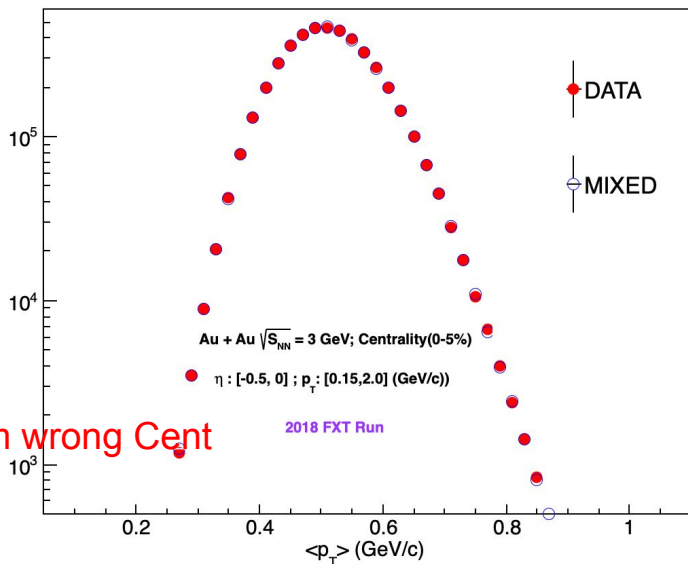
Unfolded  $\langle\langle p_T \rangle\rangle$





# Eta Study

Reconstructed  $\langle p_T \rangle$



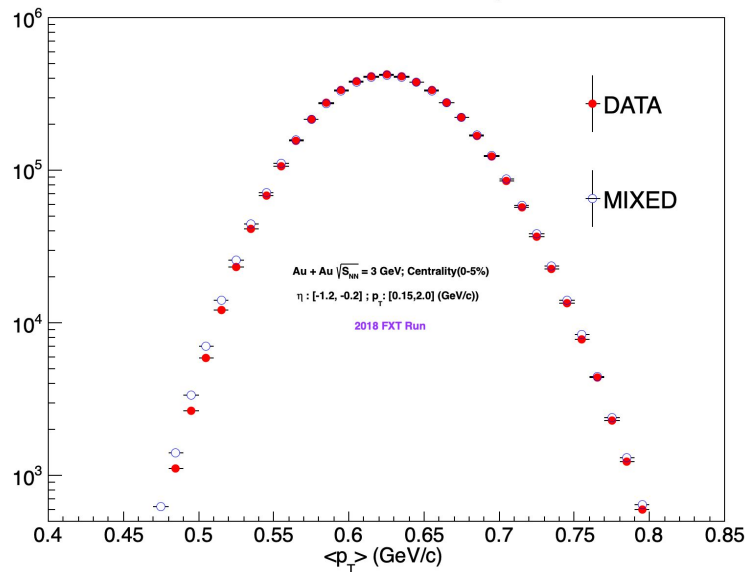
This is from wrong Cent def.

\*Efficiency  
Uncorrected !

$-0.5 < \eta < 0.0$

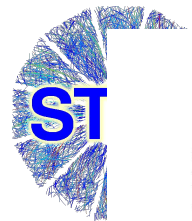
	Data	Mixed
Mean	0.5149	0.5148
Sigma	0.07925	0.07924

Reconstructed  $\langle p_T \rangle$

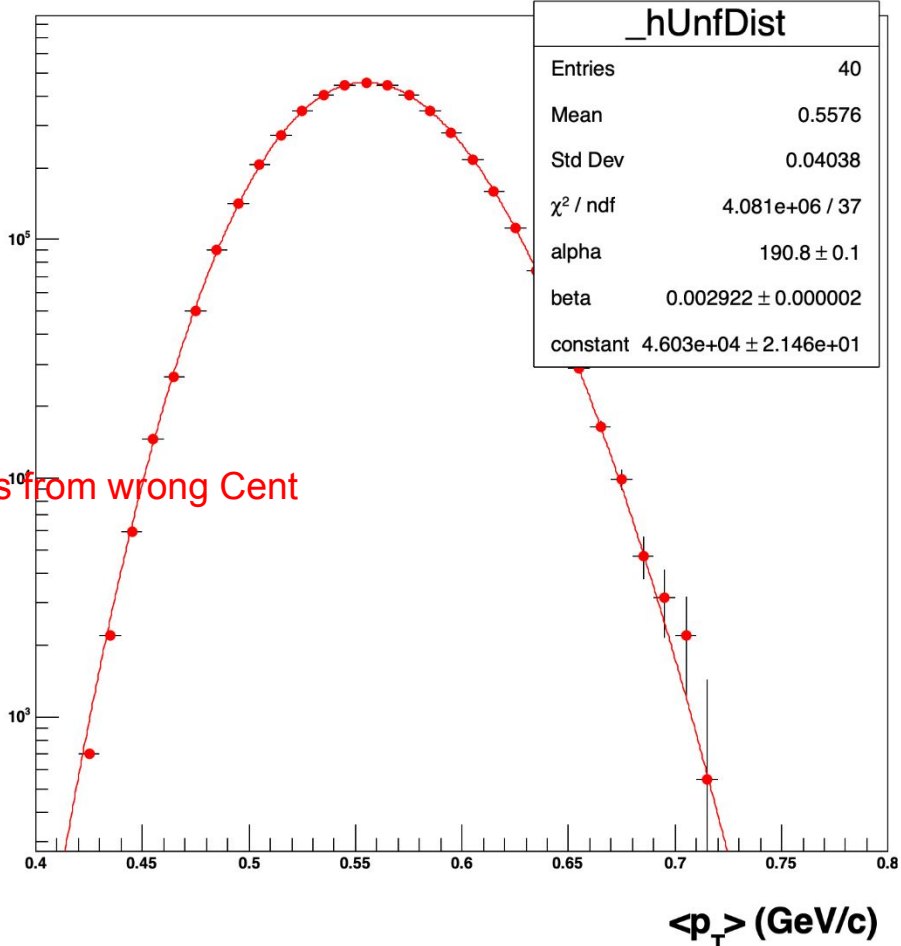


$-1.2 < \eta < -0.2$

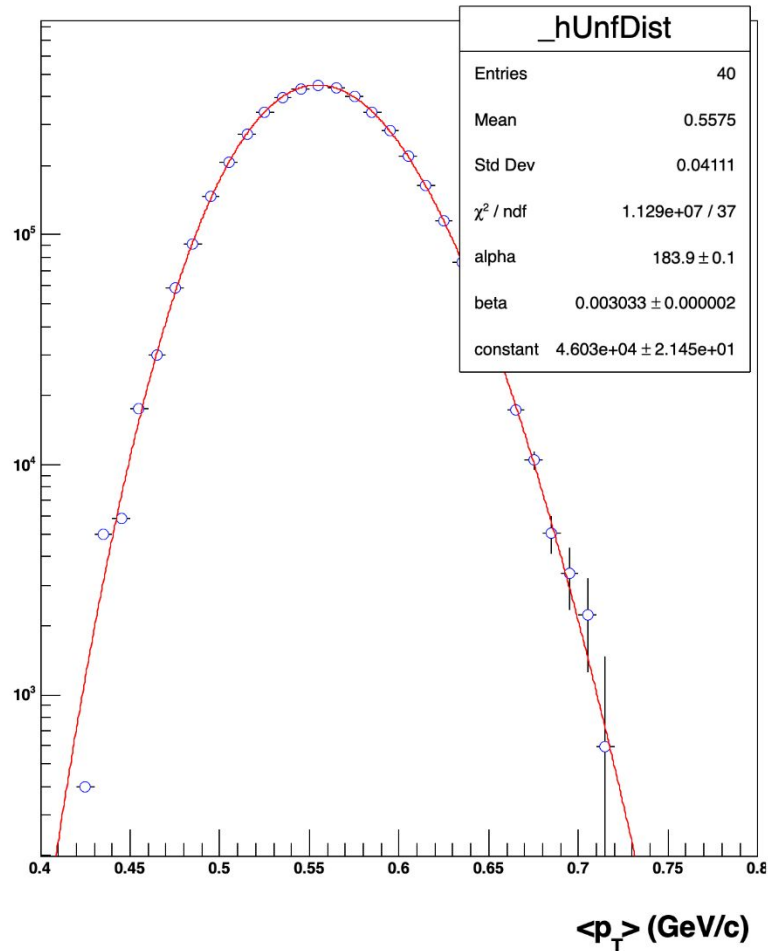
	Data	Mixed
Mean	0.6271	0.6271
Sigma	0.04338	0.04402



Unfolded  $\langle p_T \rangle$



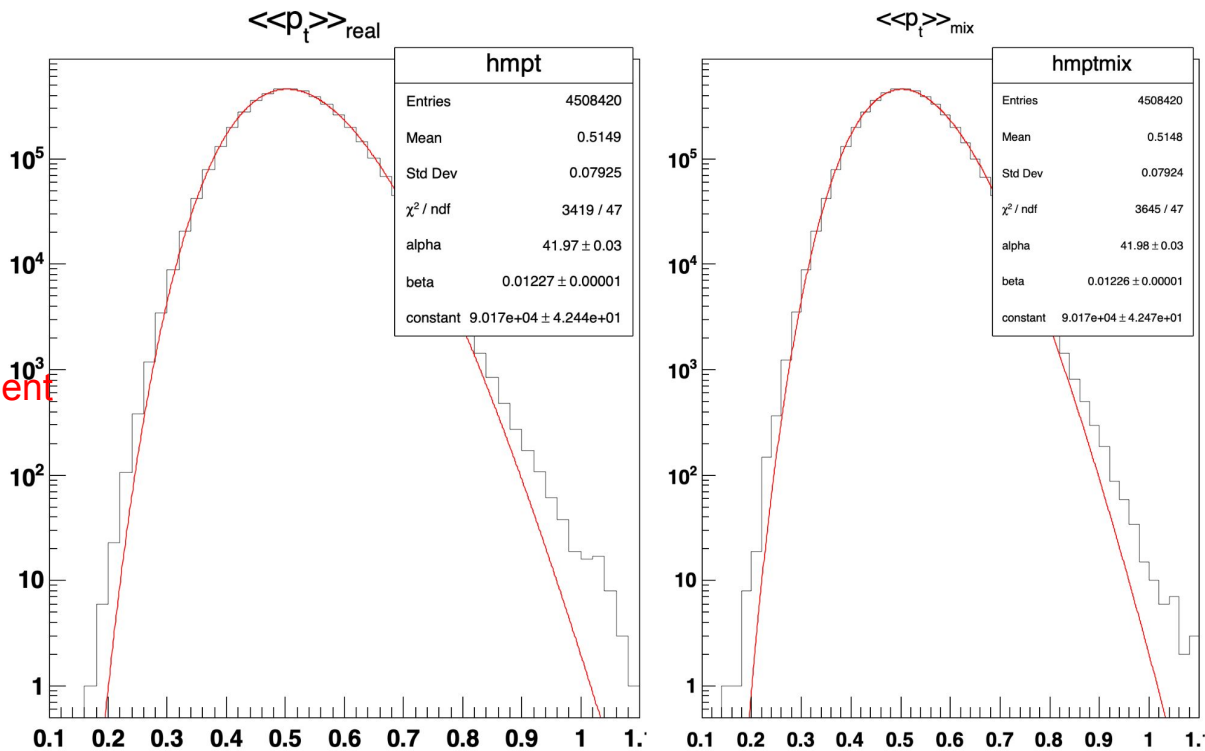
Unfolded  $\langle p_T \rangle$





# Eta Study

$$\sigma_{dyn} = \sqrt{\left(\frac{\sigma_{data}}{\mu_{data}}\right)^2 - \left(\frac{\sigma_{mix}}{\mu_{mix}}\right)^2}$$



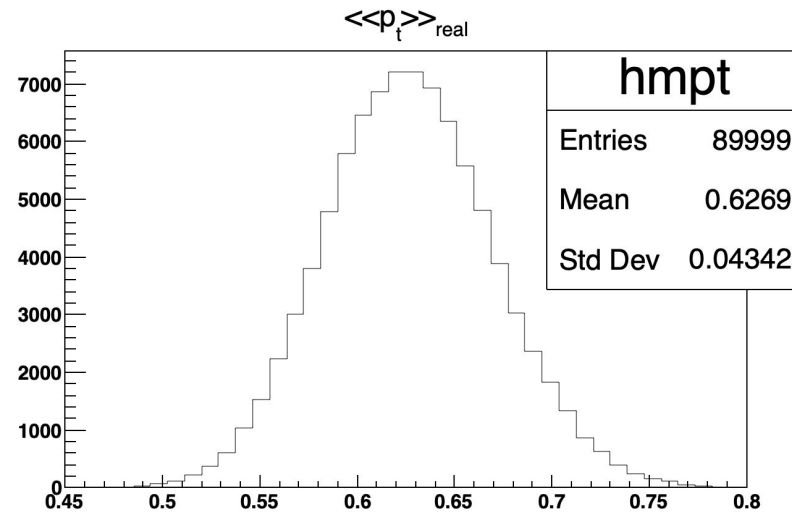
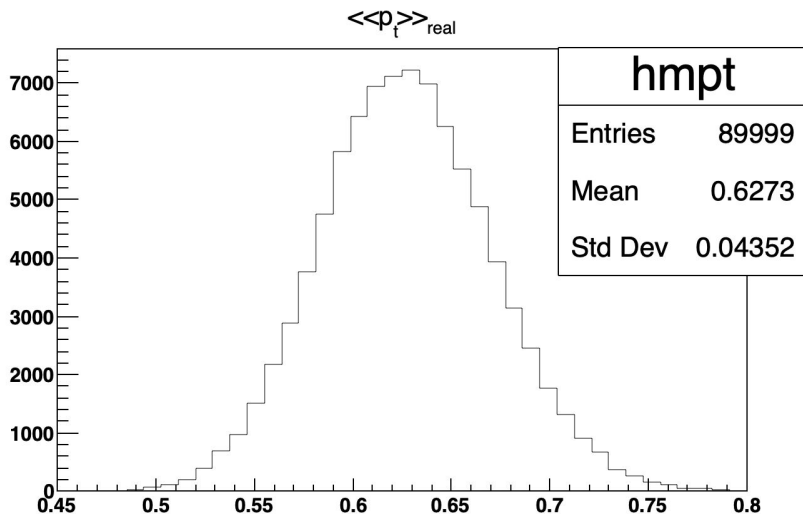
This is from wrong CentE def.





# Systematics for $\langle p_T \rangle$

For systematic contributions, BES-I analysis included DCA and nHitsFit uncertainties.  
Dan Cebra provided a list for Fluctuation analysis. 3.7M (0-5%) out of 140M events

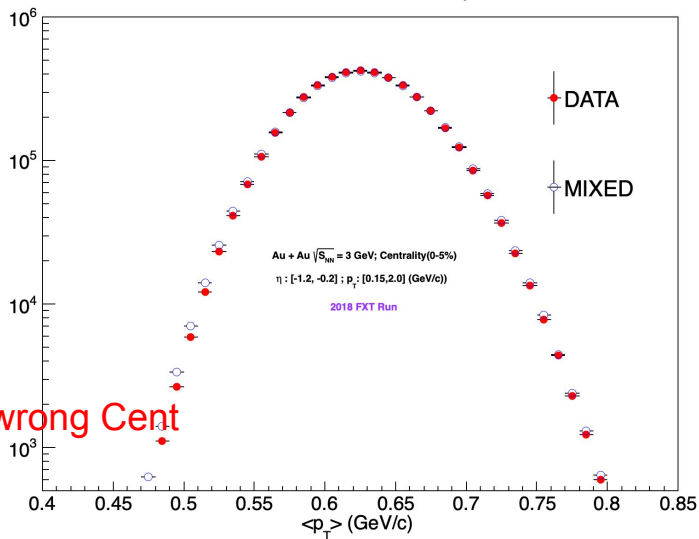




# Mixed Event Analysis for $\langle p_T \rangle$

4.6 M\*

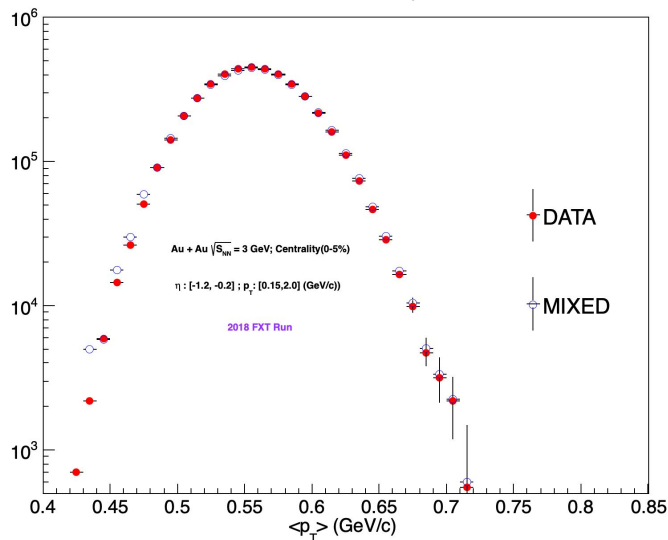
Reconstructed  $\langle p_T \rangle$



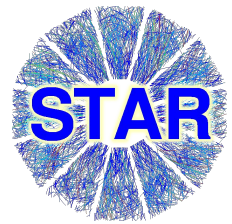
This is from wrong Cent def.

	Data	Mixed
Mean	0.6271	0.6271
Sigma	0.04338	0.04402

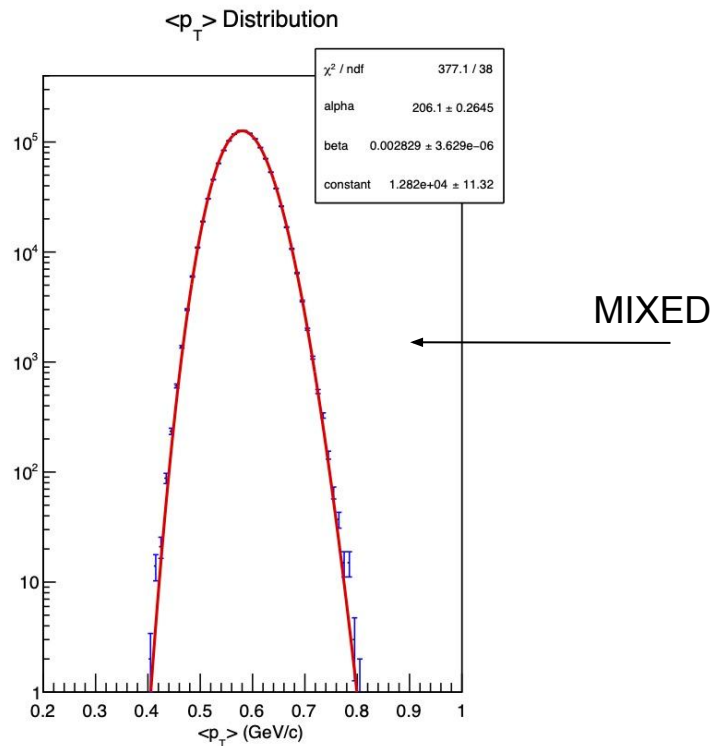
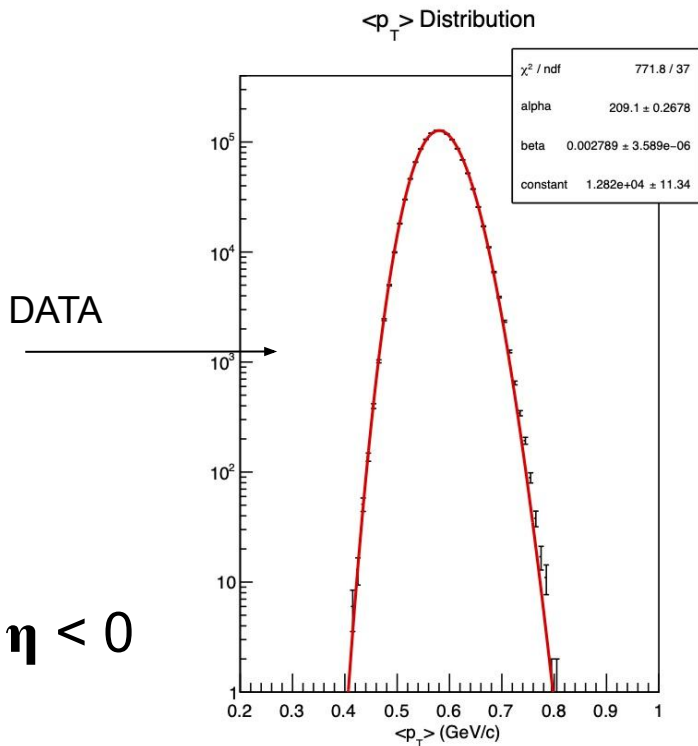
Unfolded  $\langle p_T \rangle$



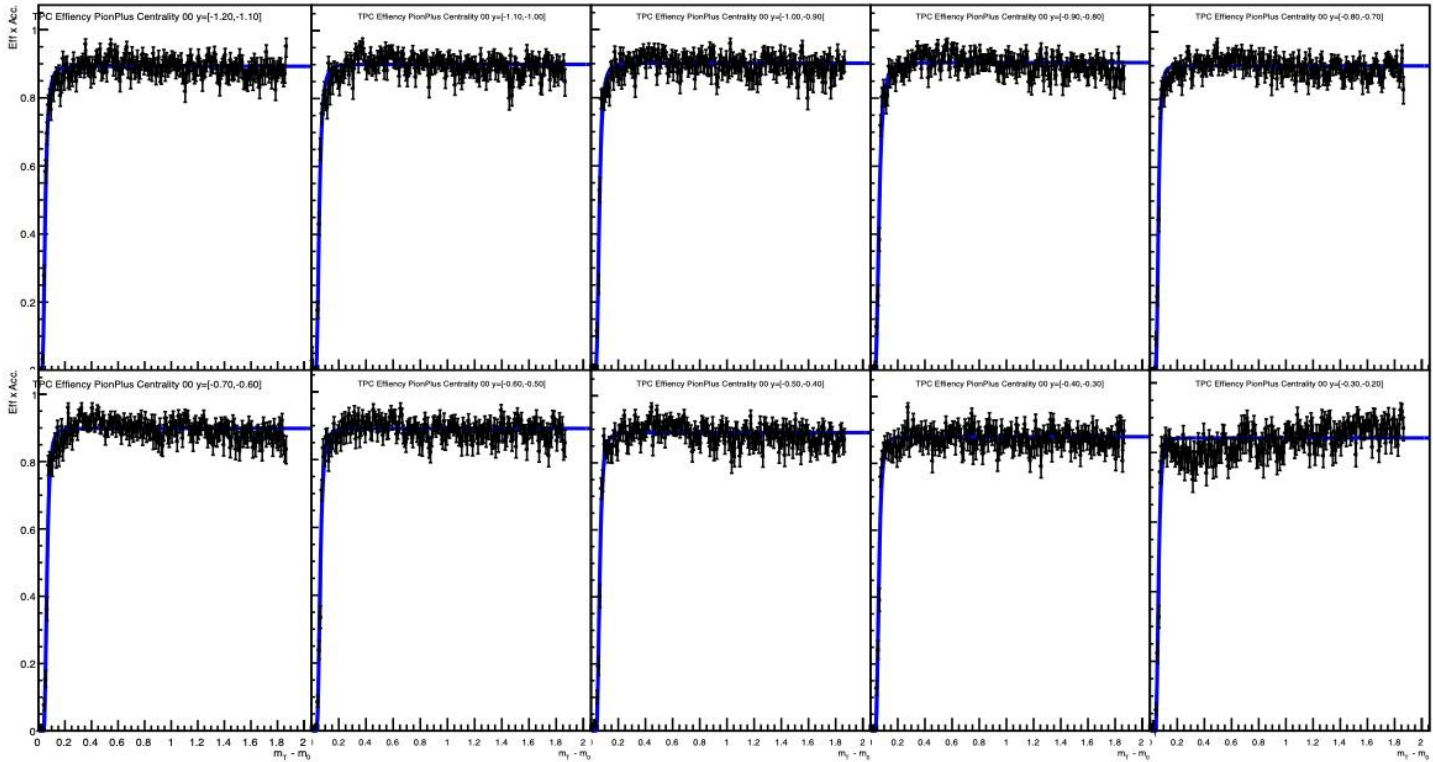
	Data	Mixed
Mean	0.5576	0.5575
Sigma	0.04038	0.04111



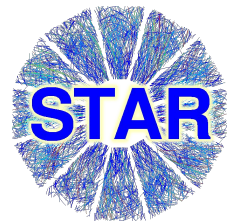
# Backup Slides



❖  $-1 < \eta < 0$



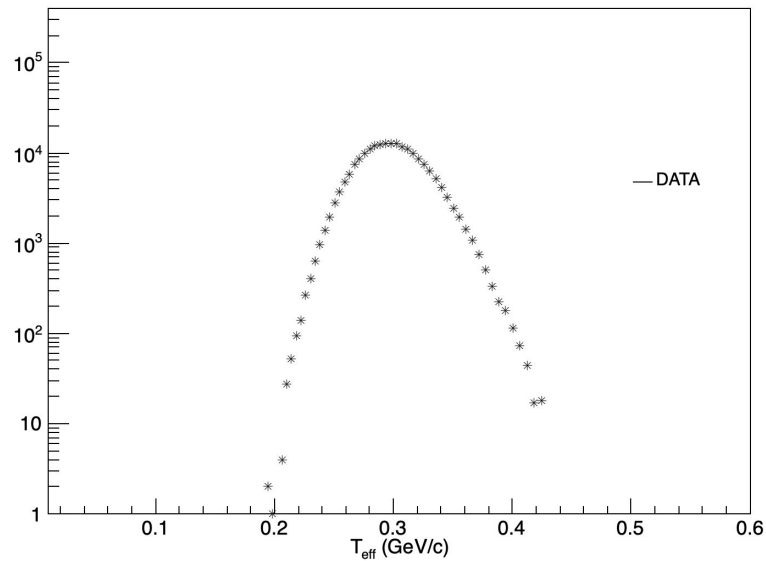
## Pion (+) Efficiencies

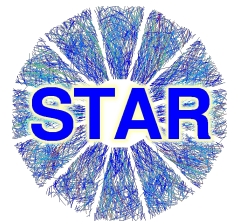


$$\langle p_T \rangle = 2T_{eff}$$

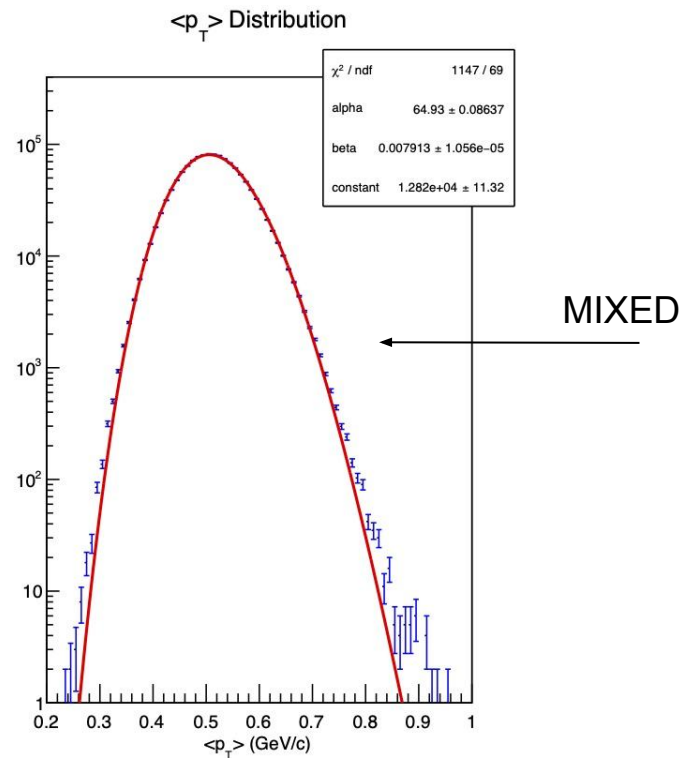
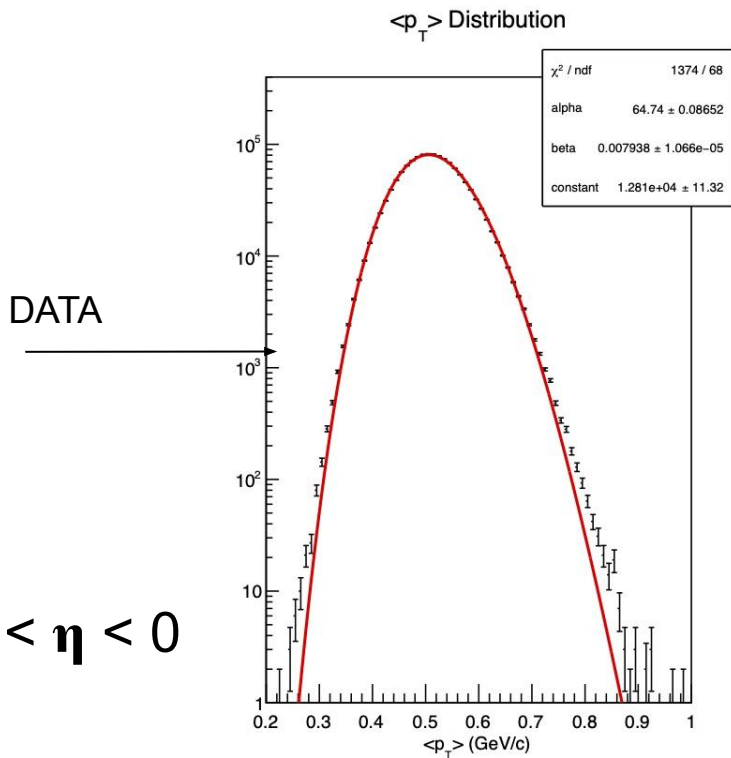
$$+ \frac{a^2 e^{-a/T_{eff}} - b^2 e^{-b/T_{eff}}}{(a+T_{eff})e^{-a/T_{eff}} - (b+T_{eff})e^{-b/T_{eff}}}$$

$T_{eff}$  Distribution,  $-1.2 < \eta < -0.2$





# Backup Slides



❖  $-0.5 < \eta < 0$



# Backup Slides

- ❖ A small subset of 27 GeV is taken for analysis.

- ❖ The event cuts and track cuts are from Chun-Jian.

[https://drupal.star.bnl.gov/STAR/system/files/BES\\_200\\_54\\_27\\_meanpT\\_0119.pdf](https://drupal.star.bnl.gov/STAR/system/files/BES_200_54_27_meanpT_0119.pdf)

$\langle p_T \rangle$  Distribution

Au + Au  $\sqrt{s_{NN}} = 27$  GeV; Centrality(0-5%)  
-2.0 <  $\eta$  < 0.0

